

RF Academy Measurement Fundamentals

What Is a Vector Signal Transceiver (VST)?

Super-Heterodyne Signal Analyzers Description and Applications

Optimizing IP3 and ACPR Measurements Tips for Getting the Best Performance from Vector Signal Analyzers

Introduction to Network Analyzer Measurements Fundamentals and Background



What Is a Vector Signal Transceiver (VST)?

Software-defined RF test system architectures have become increasingly popular over the past several decades. Almost every commercial off-the-shelf (COTS) automated RF test system today uses application software to communicate through a bus interface to the instrument. As RF applications become more complex, engineers are continuously challenged with the dilemma of increasing functionality without increasing test times, and ultimately test cost. While improvements in test measurement algorithms, bus speeds, and CPU speeds have reduced test times, further improvements are necessary to address the continued increase in the complexity of RF test applications.

To address the need for speed and flexibility, COTS RF test instruments have increased their usage of field-programmable gate arrays (FPGAs). At a high level, FPGAs are reprogrammable silicon chips that you can configure to implement custom hardware functionality through software development environments. While FPGAs in RF instrumentation is a good first step forward, typically these FPGAs are closed with fixed personalities designed for specific purposes and allow little customization. This is where user-programmable FPGAs have a significant advantage over closed, fixed-personality FPGAs. With user-programmable FPGAs, you can customize your RF instrument to the pin so that it is specifically targeted toward your application needs.

A vector signal transceiver (VST) is a new class of instrumentation that combines a vector signal generator (VSG) and vector signal analyzer (VSA) with FPGA-based real-time signal processing and control. The world's first VST from National Instruments also features a user-programmable FPGA, which allows custom algorithms to be implemented directly into the hardware design of the instrument. This software-designed approach allows a VST to have the flexibility of a software-defined radio (SDR) architecture with RF instrument class performance. Figure 1 (below) illustrates the difference between traditional approaches to RF instrumentation and a software-designed approach with a VST.

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Vector Signal Transceiver Advantages

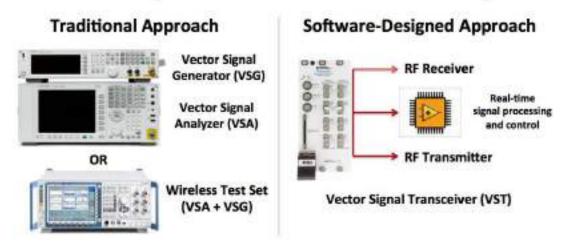


Figure 1. Compare the software-designed approach of a VST with traditional approaches.

NI VST: Built on LabVIEW FPGA and NI RIO Architecture

The NI LabVIEW FPGA Module extends the LabVIEW system design software to target FPGAs on NI reconfigurable I/O (RIO) hardware, such as the NI VST. LabVIEW is well suited for FPGA programming because it clearly represents parallelism and data flow, so users who are both experienced and inexperienced in traditional FPGA design can productively apply the power of reconfigurable hardware. As a system design software, LabVIEW is uniquely capable of blending processing done on an FPGA and a microprocessor (in your PC environment) in a way that does not require extensive knowledge of computing architectures and data manipulation. This is crucial for assembling modern communications test systems.

NI VST software is built on this powerful LabVIEW FPGA and NI RIO architecture, and features a multitude of starting points for your application including application IP, reference designs, examples, and LabVIEW sample projects. These starting points all feature default LabVIEW FPGA personalities and prebuilt FPGA bitfiles to help you get started quickly. Without these out-of-the-box capabilities, the productivity of LabVIEW, and the well-crafted application/firmware architecture, the software-designed nature of the VST would be challenging for many classes of users. With these traits, however, it brings unprecedented levels of customizability to high-end instrumentation.

Enhancing Traditional RF Test

NI VSTs feature both the fast measurement speed and small form factor of a production test box combined with the flexibility and high-performance expectation of instrument-grade box instruments. This gives the VST the ability to test standards such as 802.11ac with an error vector magnitude (EVM) of better than -45 dB (0.5%) at 5.8 GHz. In addition, the transmit, receive, baseband I/Q, and digital

inputs and outputs all share a common user-programmable FPGA, making the VST much more powerful than traditional box instruments.

Data reduction is a prime example, where decimation, channelization, averaging, and other custom algorithms allow the FPGA to perform the computationally intensive tasks. This decreases test time by reducing necessary data throughput and processing burden on the host, and it allows for increased averaging, which gives users a higher confidence in their measurement. Other examples of FPGA-based, user-defined algorithms include custom triggering, FFT engines, noise correction, inline filtering, variable delays, power-level servoing, and much more.

Software-designed instruments such as the VST can also help to bridge the gap between design and test, allowing test engineers to incorporate or validate aspects of the design before it is complete, while allowing design engineers to use instrument-class hardware to prototype their algorithms and evaluate their designs earlier in the design flow.

Example: FPGA-Based DUT Control and Test Sequencing

In addition to the baseband I/Q data of the RF receiver and transmitter, the PXI VST also features highspeed digital I/O directly connected to the user-programmable FPGA. This allows users to drastically reduce test times by implementing custom digital protocols to control the device under test (DUT). See Figure 2 below for an example. In addition to this, test sequencing can be performed on the FPGA, allowing the DUT to change states and sequence through tests in real time.

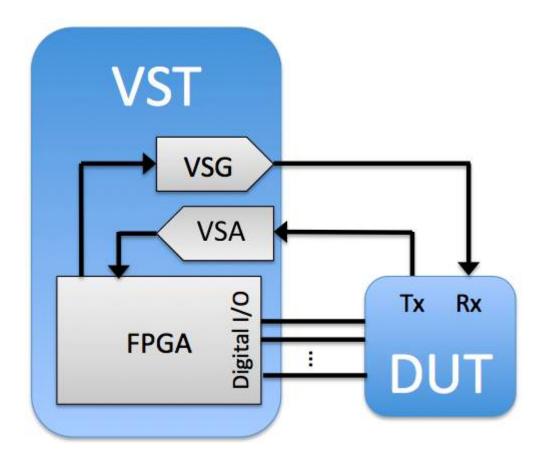
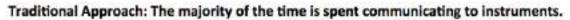


Figure 2. The flexible digital I/O capability of a VST can control the state of an RF transceiver.

Example: Power-Level Servoing for Power Amplifier Test

It is important for power amplifiers (PAs) to have an expected output power, even outside their linear operating modes. To accurately calibrate a PA, a power-level servo feedback loop is used to determine the final gain. Power-level servoing captures the current output power with an analyzer and controls the generator power level until desired power is achieved, which can be a time-consuming process. In simplest terms, it uses a proportional control loop to swing back and forth in power levels until the output power-level converges with the desired power. A VST is ideal for power-level servoing because the process can be implemented directly on the user-programmable FPGA, resulting in a much faster convergence on the desired output power value (see Figure 3).



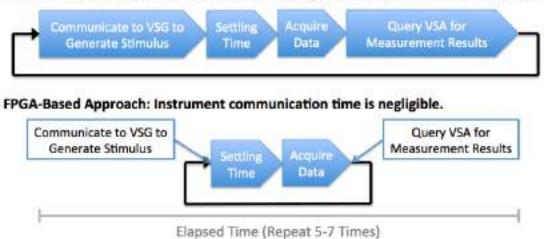


Figure 3. Using a VST for power-level servoing results in much faster convergence on the desired output power level during PA test.

Other RF Applications

A VST is more than just an incredibly fast and flexible vector signal analyzer and vector signal generator. The RF receiver, RF transmitter, and user-programmable FPGA also allow a VST to go beyond the traditional VSA/VSG paradigm. For example, the VST can be completely redesigned by the user to perform complex processing for other RF applications such as prototyping new RF protocols, implementing a software defined radio, and channel emulation among others.

Example: Radio Channel Emulator for MIMO RF Signals

In recent years, multiple input, multiple output (MIMO) RF technology has grown significantly, especially in cellular and wireless standards. In addition to this, RF modulation schemes are growing in complexity, RF bandwidth is increasing, and radio spectrums are becoming more crowded. With these

advances in technology, it is important to not only test wireless devices in a static environment, but to understand how these devices behave in a dynamic real-world environment as well.

A radio channel emulator is a tool for testing wireless communication in a real-word environment. Fading models are used to simulate air interference, reflections, moving users, and other naturally occurring phenomenon that can hamper an RF signal in a physical radio environment. By programming these mathematical fading models onto the FPGA, a VST implements a real-time radio channel emulator. Figure 4 below shows a 2x2 MIMO radio channel emulator implemented using two VSTs in LabVIEW. Settings for the fading models are shown on the left and in the center of the screen. The resulting RF output signals from the fading models were acquired with spectrum analyzers and are displayed on the right. These spectral graphs clearly show the spectral nulls that have resulted from the fading models.

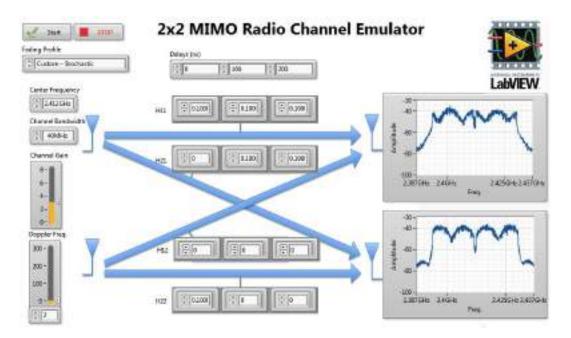


Figure 4. An example LabVIEW front panel shows the effect of MIMO channel emulation implemented using two VSTs.

Multiple Possibilities for Software-Designed Instrumentation

The VST represents a new class of instrument that is software designed, with capabilities limited only by the user's application requirements—not the vendor's definition of what an instrument should be. As RF DUTs become more complex and time-to-market requirements become more challenging, this level of instrument functionality shifts control back to the RF designer and test engineer. The examples shown in this document barely scratch the surface of what a VST is capable of. To answer the "What is a vector signal transceiver?" question, you have to first answer the question of "What RF measurement and control problem do you need to solve?" With the flexibility of an accurate RF transmitter, RF receiver, and digital I/O connected to a user-programmable FPGA, the VST is more than likely up to the challenge.



Super-Heterodyne Signal Analyzers

Description and Applications

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1. Signal Analysis Background

Introduction

Signal analyzers encompass test and measurement receivers known as spectrum analyzers and vector signal analyzers. The signal analyzer is to the frequency domain what the oscilloscope is to the time domain: a general purpose test instrument that can measure and display electrical signals. Like oscilloscopes, signal analyzers come in varying grades of performance and feature set offerings. For the most part, bandwidth distinguishes one oscilloscope from another. However, signal analyzers can vary in tuning range, bandwidth, dynamic range and measurement accuracy (both frequency and amplitude). The signal analyzer performance criterion or feature set offering needed is dependent on the measurement requirement. In addition, signal analyzers vary by their internal architectures. Some architectures are better suited than others for a particular measurement requirement. Lastly, performance alone does not guarantee success in the measurement. Like any precision instrument, knowing how to optimize the settings goes a long way toward enabling the best performance capability of the signal analyzer.

This application note serves as an introduction to the architecture and proper use of one particular signal analyzer structure: the super-heterodyne signal analyzer. Subjects covered in this application note:

- Link between time domain and frequency domain signal analysis
- Super-heterodyne principle: how the mixing process creates wanted and unwanted responses
- Architectural differences of various super-heterodyne signal analyzers
- RF Chain signal processing: how the gain control elements affect the measurement
- IF Chain signal processing: digital signal processing for detection, bandwidth setting and averaging
- Dynamic range optimization
- Amplitude accuracy
- Spectrum monitoring applications

Time and Frequency Domain Representations of Signals

Signal analyzers are a frequency domain class of test instrumentation. However, before introducing the signal analyzer, it is illustrative to review the more familiar time domain representation of signals.

Suppose the signal of interest is a sinusoid given by the equation

$$s(t) = Asin(2\pi f_o t + \theta)$$

Equation 1-1.

Time, t, is the independent variable, whereas frequency, f_0 is a fixed constant value. Everything inside the bracket () is the instantaneous phase. For simplicity, equate the instantaneous phase to $x:x = (2\pi f_o t + \theta)$. As time, t, advances so too does the instantaneous phase, x. When $x=\pi/2$, sin(x) is at its maximum and when $x = 3\pi/2$, sin(x) is at its minimum value. For x = 0, π , and 2π , sin(x) = 0. This process repeats itself with a period, $T = 1/f_0$ as shown in **Figure 1-1**.



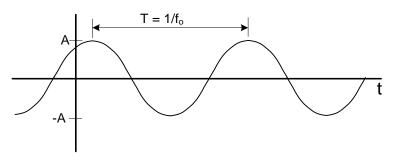


Figure 1-1. Time Domain Representation of a Sinusoid

The sinusoid represented in the time domain is a very familiar to any user of an oscilloscope. The sinusoid has further significance in that a *Fourier series expansion* of any periodic signal is a sum of sine and cosine terms [¹]:

$$x(t) = a_0 + 2\sum_{n=1}^{\infty} \left[a_n \cos\left(\frac{2\pi nt}{T}\right) + b_n \sin\left(\frac{2\pi nt}{T}\right) \right]$$

Equation 1-2.

For example, a square wave represented as a Fourier series expansion has the form [²]:

$$x(t) = \frac{4}{\pi} \sum_{n \text{ odd}}^{\infty} \frac{1}{n} \sin\left(\frac{2\pi nt}{T}\right)$$

Equation1-3.

As shown in **Figure 1-2**, one can see how a square wave begins to form as the odd harmonics add together.

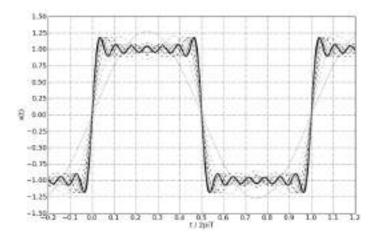


Figure 1-2. Square Wave Fourier Series Expansion



The ability to deconstruct any periodic signal into a series of sinusoids plays an important role in making the leap between the time domain and the frequency domain. Another link between the time and frequency domains is the Fourier transform, itself. The Fourier transform operates on non-periodic time domain signals and is given by **Equation 1-4**.

$$x(t) = \int_{-\infty}^{\infty} X(f) e^{j2\pi ft} df$$
$$X(f) = \int_{-\infty}^{\infty} x(t) e^{-j2\pi ft} dt$$

Equation 1-4.

X(f) is the frequency domain representation of the time domain signal, x(t). Some example transform pairs are shown in **Figure 1-3**. A sinusoid in the time domain translates to a pair of impulses in the frequency domain **Figure 1-3a**. The locations of the impulses are +/- the frequency value of the sinusoid, f_o . A single rectangular pulse in the time domain translates to a sinc function in the frequency domain **Figure 1-3b**. Finally, the pairing of the Fourier series expansion and the Fourier transform is evident in the pulse train. In the frequency domain, this pulse train is a series of impulse functions (sinusoids) in which the envelope of these discrete sinusoids results in the sinc function **Figure 1-3c**.

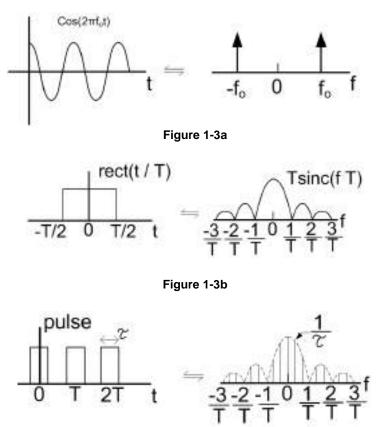


Figure 1-3c

Figure 1-3. Fourier Transform Pairs for Cosine, Single Rectangular Pulse, Periodic Pulse Train



By applying the Fourier transform to each sinusoid component of the square wave represented in **Equation 1-3**, the frequency domain view clearly shows the odd harmonics. **Figure 1-4** shows the time and frequency domain representation of the square wave.

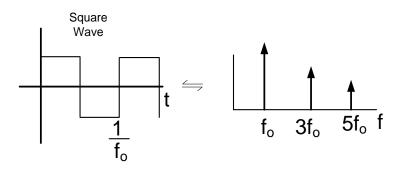


Figure 1-4. Fourier Transform Pair for the Square Wave

Why View Signals in the Frequency Domain?

Just because the Fourier transform allows a signal to be viewed in the frequency domain, why is this necessary? The answer is that examining the signal in the frequency domain allows some added insight in that is just not possible in the time domain. As an example, **Figure 1-5** shows a sinusoid with harmonic perturbations added. Both the time domain and frequency domain views are shown.

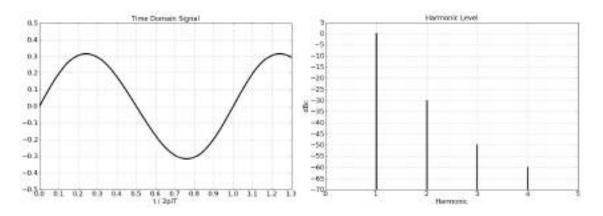


Figure 1-5. Sinusoid with Harmonics

When viewed in the time domain, the signal in **Figure 1-5** appears to be an undistorted sinusoid. It is only in the frequency domain that relatively large harmonics are apparent. In many situations, a -30 dBc 2nd harmonic, which can only be deciphered in the frequency domain, may be unacceptable.

Another situation where frequency domain information is crucial is in diagnosing circuit and system problems. **Figure 1-6** shows the time domain view of a heavily distorted signal. The time domain view only gives the user confirmation that there is in fact a heavily distorted waveform. No information is readily apparent on what could be causing this distortion.



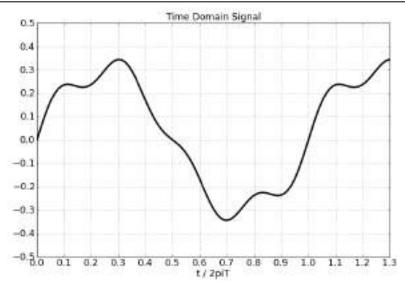


Figure 1-6. Heavily Distorted Signal in the Time Domain

Only in the frequency domain, shown in **Figure 1-7**, is it apparent that most of the problem is a result of large 3rd and 4th harmonic levels.

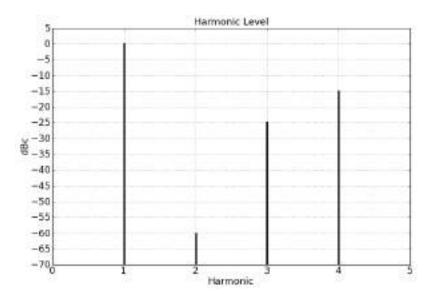


Figure 1-7. Frequency Domain View of the Signal in Figure 1-6

Another situation where the frequency domain might give more insight is in the examination of jitter. Normally associated with digital signals, jitter results from phase noise causing time crossing fluctuations in the rising and falling edges of the digital signal. Jitter in the time domain is shown in the top graph of **Figure 1-8**. Viewing the digital signal in the frequency domain shows the phase noise associated with the jitter. The phase noise as represented in the bottom graph in **Figure 1-8** has a frequency dependent amplitude pedestal.



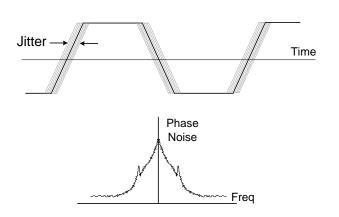


Figure 1-8. Jitter and Phase Noise.

Examining the phase noise pedestal at particular frequency offsets allows better diagnosis of issues leading to large value jitter problems. For example, in **Figure 1-8** the phase noise viewed in the frequency domain shows evidence of a spurious signal. This unwanted spur could be the source of the poor jitter performance. Whether the spur is due to power supply noise, fan vibrations, etc., the frequency domain view gives the user information that is just not possible in the time domain view.

Using the Oscilloscope for Frequency Domain Analysis

Discrete Fourier transform (DFT) and its more efficient implementation, the fast-Fourier transform (FFT), use sampled data to make the transformation between time and frequency domains. Modern oscilloscopes mostly operate on sampled data, so why not just use the oscilloscope and perform an FFT and present the data in the frequency domain? In fact, many oscilloscopes have software that does precisely this operation.

The answer to the question of why not use an oscilloscope lies in the performance requirements of the measurement. Frequency, bandwidth, and dynamic range are the main considerations.

Until recent years the bandwidth of the oscilloscope did not approach RF and microwave frequencies. Hence, a signal analyzer was the only solution for frequency domain examination of signals in the RF and Microwave range. At present, oscilloscopes are available with greater than 30 GHz bandwidth, making bandwidth not as strong of a differentiator between signal analyzers and oscilloscopes. However, price would need to be considered as signal analyzers can be the more cost effective solution for measurements at higher frequencies.

The big differentiator between signal analyzers and oscilloscopes is dynamic range, specifically the spurious-free dynamic range. This is the dynamic range specification needed for the measurement of distortion such as harmonics. Sampled data systems all require Analog-to-Digital Converters (ADC), sometimes referred to as digitizers. Spurious free dynamic range is a function of the digitizer's amplitude resolution, operating frequency, and the sample rate. Because the signal analyzer can frequency shift the RF signal down to a lower digitizer input frequency, the signal analyzer's digitizer can operate at a lower sample rate. Availability of higher resolution digitizers as well as simply operating at a lower digitizer input frequency gives an advantage to the signal analyzer over the oscilloscope in terms of its dynamic range performance. Typically, the differences in dynamic range between the signal analyzer and the oscilloscope can exceed 50 dB.



Oscilloscopes do not normally specify spurious-free dynamic range, whereas signal analyzers highlight this performance specification. This demonstrates the strengths and weaknesses of these two platforms. The oscilloscope is optimized for wide bandwidth to accurately characterize fast slew-rate signals in the time domain, whereas the signal analyzer is optimized to achieve as high a dynamic range as possible for the measurement of signal distortion.

Spectrum Analysis versus Vector Signal Analysis

As previously mentioned, the signal analyzer broadly defines the class of instrumentation known as spectrum analyzers and vector signal analyzers. In the early days of signal analyzers when only narrowband analog modulation formats existed, the spectrum analyzer was the only category. With the advent of digital modulation formats, their relatively wide modulation bandwidths brought forth the need for the vector signal analyzer. The key differentiator between the spectrum analyzer and the vector signal analyzer is the measurement bandwidth.

However, the terms spectrum analyzer and vector signal analyzer, which depict the label of the test instrument, are not always consistent from vendor to vendor. To make matters even more confusing, some spectrum analyzers are capable of making vector signal analyzer type measurements. And some so called vector signal analyzers have spectrum analyzer type traits. The following sections will define both the spectrum analyzer and the vector signal analyzer architectural differences. Because of the inconsistency in nomenclature, the user should rely more on the specifications rather than the label of the instrument.

Rather than dwell on the terms spectrum analyzer and vector signal analyzer, which are labels given by the instrument manufacturer, it is perhaps more instructive to define the measurements. *Vector signal analysis* in the context of this document refers to the measurement of the entire bandwidth of the modulated signal. This type of analysis requires that the analog IF bandwidth of the test receiver be at least as wide as the modulation bandwidth of the signal being measured. The top of **Figure 1-9** shows how the shaded region depicting the test receiver's analog IF bandwidth encompasses the entire signal. Digital modulation metrics such as error vector magnitude (EVM) and complementary cumulative distribution function (CCFD) require capturing the entire modulation bandwidth.

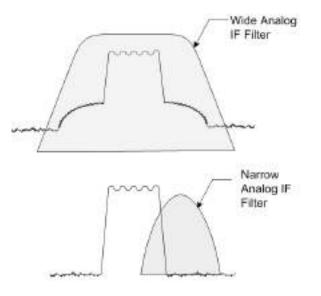




Figure 1-9. Top: Vector Signal Analysis. Bottom: Spectrum Analysis

Spectrum analysis will be used in this document to mean the measurement of the signal using the test receiver's narrow analog IF bandwidths. This is depicted in the bottom portion of **Figure 1-9**. Later sections in this document will explain why this is true, but limiting the IF bandwidth allows for higher dynamic range performance. Measurements such as adjacent channel leakage ratio (ACLR), intermodulation distortion, and harmonic distortion are greatly enhanced with a relatively narrow receiver IF bandwidth.

Super-Heterodyne versus Direct Conversion Architectures

This introductory section concludes with a mention of two major categories used in signal analyzers. The super-heterodyne receiver architecture is the subject of the rest of this document. The other architecture now being used in signal analyzers is the direct conversion receiver, which also goes by the names homodyne receiver and zero IF (ZIF) receiver. This document will not cover the direct conversion receiver in detail; however, a brief mention of this receiver will be given to highlight its difference with the super-heterodyne receiver.

Figure 1-10 shows the basic structure of the direct conversion receiver. A single local oscillator is used to shift the incoming RF signal down to baseband. Baseband contains two paths, I-path and Q-path, corresponding to In-phase and Quadrature paths. Each path is then digitized separately.

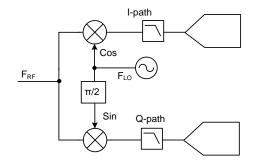


Figure 1-10. Direct Conversion Receiver.

The direct conversion receiver has benefits over the super-heterodyne receiver in terms of bandwidth and compactness, such as only one local oscillator is needed and there are fewer requirements on the RF path filtering. However, the super-heterodyne receiver, in general, is capable of more dynamic range than the direct conversion receiver.



2. Super-Heterodyne Principle

Brief History of the Super-Heterodyne Receiver

The heterodyne principle was coined by its inventor, Reginald Fessenden in 1901 [⁴]. The term has the Greek roots, heteros, meaning "other" and dynamis, meaning "force". The heteros part refers to the translation to another frequency and the dynamis part refers to the apparent amplification of the detected signal during the heterodyne process. This first instantiation did not resemble modern day receivers; rather, it used two antennas to receive two RF signals. When combined in an envelope detector, these two signals created a beat note: a signal at the difference frequency. This was an audio frequency beat note.

Little progress in this receiver was made until Edwin Howard Armstrong in 1918 was able to develop the idea of using the frequency conversion of higher frequency signals down to the range of the then common heterodyne receiver. He was able to observe that the modulation on a signal did not alter during this frequency conversion process. The super part of super-heterodyne refers to super-sonic, meaning that the heterodyne process was extended above the audio frequency range.

Frequency Shift Property

One property of the Fourier transform is the shift theorem which states that if X(f) is the Fourier transform of x(t), then the Fourier transform of x(t) multiplied by a complex exponential, $e^{j2\pi f_c t}$ results in the frequency domain signal shifted in frequency by the amount -fc [³].

$$\mathcal{F}\{x(t)e^{j2\pi f_c t}\} = X(f - f_c)$$

Equation 2-1

Manipulating Euler's Identity allows a sinusoid to be represented by a pair of complex exponentials [3]. For example, the cosine signal representation is shown in **Equation 2-2**.

$$\cos(2\pi f_c t) = \frac{e^{j2\pi f_c t} + e^{-j2\pi f_c t}}{2}$$

Equation 2-2

By multiplying a time domain function, x(t), with a sinusoid and applying the Fourier transform shift theorem, the result is shown in **Equation 2-3**.

$$\mathcal{F}\left\{x(t)\left[\frac{e^{j2\pi f_c t} + e^{-j2\pi f_c t}}{2}\right]\right\}$$
$$= \frac{1}{2}[X(f - f_c) + X(f + f_c)]$$

Equation 2-3

In the context of the Fourier transform, **Equation 2-3** is known as the modulation theorem [³]. Multiplying a signal by a sinusoid whose frequency is f_c results in two copies, each with amplitude multiplier of $\frac{1}{2}$, being shifted by $\pm f_c$.



In receivers, the mixer is the device that performs the frequency conversion process. The block diagram for the mixer is shown in **Figure 2-1**.

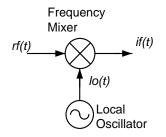


Figure 2-1. Frequency Conversion Process

The local oscillator (LO) is a sinusoid signal that can be tuned in frequency. To a first order approximation, the mixer performs straight time domain multiplication:

$$if(t) = rf(t) \ge lo(t) = rf(t) \ge cos(2\pi f_{L0}t)$$

Equation 2-4

Applying Equation 2-3 to Equation 2-4 results in the frequency domain representation:

$$IF(f) = \frac{1}{2} [RF(f - f_{LO}) + RF(f + f_{LO})]$$

Equation 2-5

Figure 2-2 shows how the frequency mixing appears in the frequency domain. The signal RF(f) centered at DC before mixing appears at the IF port as two instances of itself centered at frequencies $-f_{LO}$ and $+f_{LO}$

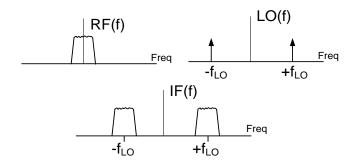


Figure 2-2. Modulation Theorem Applied to RF Signal Mixing

Note that the modulation content in the signal is preserved throughout this process.

Frequency Shift Property Applied to the Super-heterodyne Receiver

Now to see how the frequency shift property can be extended to the super-heterodyne receiver. The fundamental building blocks of the super-heterodyne receiver are shown in **Figure 2-3**.



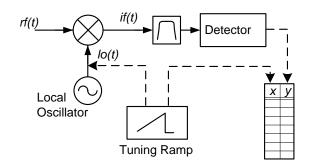


Figure 2-3. Basic Super-heterodyne Receiver

The full super-heterodyne structure adds some IF filtering and some means of converting the IF signal to magnitude and phase data by use of the block labeled *Detector*. The detected IF is converted to a digital value and recorded as y-data. A numeric tuning value used to control the LO frequency is also stored as x-data. The recorded x,y pairs corresponding to frequency / amplitude or frequency / phase data can be used directly to display the frequency domain spectrum of the signal or this data can be used for further signal processing.

Figure 2-4 shows the frequency shifting process in a super-heterodyne receiver.

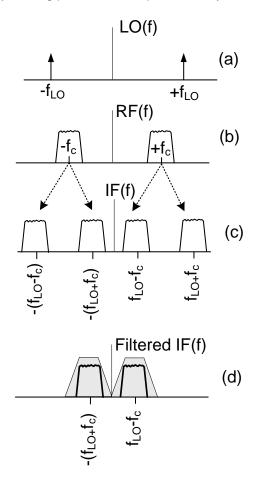




Figure 2-4. Mixing Process in a Super-heterodyne Receiver

Instead being centered at DC, the RF input signal, in general, is centered at some carrier frequency, f_c . The mixer functionally multiplies the LO signal with the RF signal. At the IF port of the mixer, the positive frequency component of the RF signal splits into two copies centered at $f_{LO} - f_c$ and $f_{LO} + f_c$. The negative frequency portion of the RF signal similarly splits into two copies at $-(f_{LO} - f_c)$ and $-(f_{LO} + f_c)$. After IF filtering, only the copies centered at $\pm (f_{LO} - f_c)$ are retained.

The negative frequency content is important for signal processing of complex signals, which is the case for the direct conversion receiver. However, for the analog portion of the super-heterodyne receiver, the signals are real, meaning that there is even magnitude symmetry and odd phase symmetry about DC as shown in the following:

$$|X(f)| = |X(-f)|$$

and
$$phase(X(f)) = -phase((X(-f)))$$

For analyzing the analog portion of the super-heterodyne receiver, concentrating on only the positive portion of the frequency spectrum does not lose any information about the signal.

Moving to a Non-ideal Receiver

The previous sections rely on idealized components to preserve the frequency shift process in a superheterodyne receiver. However, in actual receivers the components are far from ideal in terms of added noise and distortion to the measurement. Maintaining best measurement performance for dynamic range and measurement accuracy (both amplitude and frequency) requires a careful optimization of the system parameters. But, before discussing how to optimize system parameters, knowledge of the sources of the distortion and noise will be presented.

Mixing Process

Figure 2-5 once again reveals the basic pieces required to understand the frequency conversion (or mixing) process. Compared to **Figure 2-1**, a bandpass filter centered at the desired IF frequency has been added at the output port of the mixer.

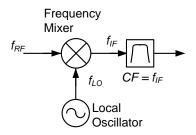


Figure 2-5. Basic Frequency Mixer

The *mixing equation*, shown in **Equation 2-6**, is a simplification of the frequency shift property.

$$Mxf_{RF} + Nxf_{LO} = f_{IF}$$

where, M and N = 0, ± 1 , ± 2 , ± 3 , . . .



Equation 2-6

The RF signal "mixes" or frequency shifts to an IF signal using a sinusoid LO signal. Whenever the mixing equation is satisfied, an IF signal passes through the bandpass filter and is recorded as a response. But, there are an infinite number of solutions that satisfy the mixing equation, resulting in an infinite number of possible wanted and unwanted signals converting to IF.

Normally, the receiver is calibrated to be accurate for only a single pair of M, N values. All other M,N combinations that cause a response to fall inside the bandpass filter are unwanted and are known as spurious responses, or spurs for short. Many of these spurs are given specific names such as image responses. There are far more spurs than wanted responses and by understanding the spur mechanisms the user can avoid having the receiver's spur mask the signal being measured.

First, the mixing process for desired mixing responses will be examined. Suppose the signal analyzer is calibrated to respond to M,N = -1,1 and let the IF be centered at 100 MHz. Assume that a fixed frequency sinusoid, which is referred to a continuous wave (CW) signal, at a frequency of 1 GHz is present at the RF input port. The mixing equation simplifies to:

$-f_{RF} + f_{LO} = f_{IF}$

Equation 2-7

Solving **Equation 2-7** shows that at an LO frequency of 1100 MHz, the mixing equation is solved, resulting in an IF signal at a frequency of 100 MHz:

-1000 MHz + f_{L0} = 100 MHz -1000 MHz + 1100 MHz = 100 MHz

The spectrum view of this mixing process is shown in Figure 2-6.

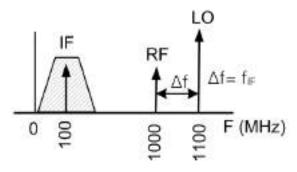


Figure 2-6. Spectrum Showing the Mixing Process for a Wanted Mixing Product

In a signal analyzer, often times the data is gathered over a range of frequencies. The LO is either stepped or swept in frequency. With a fixed frequency RF signal, the resulting IF, after applying the mixing equation, will also be swept or stepped. Furthermore, the shape of the IF filter will be traced out in the measured amplitude data. The process of stepping the LO over a range of frequencies with a fixed frequency RF input signal is shown in **Figure 2-7**.



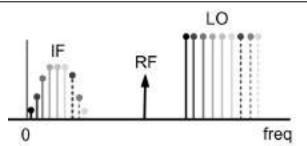


Figure 2-7. Mixing Process with Swept LO

Image Responses

Previously, the wanted M,N = -1,1 mixing product was analyzed. But the mixer will also respond to the M,N = 1,-1 product as well. Using the above example where the LO was tuned to 1100 MHz in order to convert a 1000 MHz RF signal down to an IF of 100 MHz, suppose another RF signal was present at 1200 MHz. In this case with M,N = 1,-1 the mixing equation simplifies to:

$$f_{RF} - f_{LO} = f_{IF}$$

or, 1200 MHz - 1100 MHz = 100 MHz

Equation 2-8

Figure 2-7 shows the spectrum with both the wanted RF at 1000 MHz mixing down to the 100 MHz IF as well as the unwanted RF signal at 1200 MHz also mixing down to the 100 MHz IF.

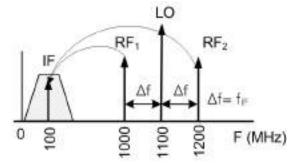


Figure 2-7. Desired Signal and its Image Response.

The unwanted spur at 1200 MHz is known as the image response. The image response, as evident from **Figure 2-7**, falls two times the IF frequency away from the desired response.

In **Figure 2-7**, the LO frequency is fixed and RF either one-IF below or one-IF above can cause a response. **Figure 2-8** shows another situation where the image response causes non-deterministic results. In this situation, the LO is tuning in order to create a swept spectrum display. In this example, a single RF signal at 1 GHz is present at the input port.



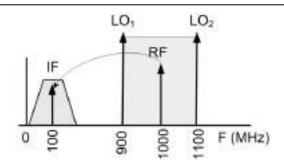


Figure 2-8. Single RF Causing an Image Response

When the signal analyzer is tuned to 800 MHz, the corresponding LO frequency is 900 MHz. The RF and LO signal mix using M,N = 1,-1 to create an IF response at 100 MHz. As the LO frequency is stepped, eventually it reaches 1100 MHz. This time the RF and LO mix using M,N = -1,1 once again creating a response at 100 MHz centered IF. The spectrum display of this scenario is shown in **Figure 2-9**.

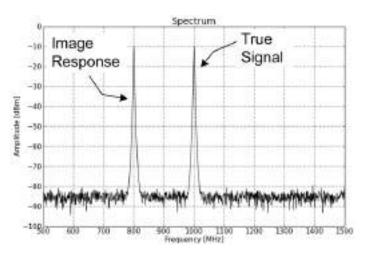


Figure 2-9. Spectrum of Single RF Causing Image

The two responses are separated by two IFs (200 MHz) in frequency and also note that their amplitudes are nearly the same. Remember, only a single tone at 1 GHz is input, yet two responses are present on the display. The 800 MHz displayed response, the image, is not real; however, the operator would not be able to decipher the image from the true response.

In signal analyzers with multiple frequency conversion stages, there will be images associated with each stage. For each stage, images will be present and will be spaced away from the desired signal by twice the IF frequency of that stage. Most often, the LO and IF frequencies in the later frequency conversion stages are fixed, so predicting the image frequency is not as difficult as with frequency conversion stages where the LO frequency is variable. The amount of suppression of the image signal amplitude is termed image rejection and applies to all the mixer stages in the system.

IF Subharmonics

Another spur mechanism results from signals mixing to sub-multiples of the IF frequency. The mixing equation, shown in **Equation 2-6**, can be modified to include IF subharmonics by adding the qualifier, Q:

$$Mxf_{RF} + Nxf_{LO} = Qxf_{IF}$$



where, M and N = $\pm 1, \pm 2, \pm 3, \ldots$ and Q = 1/2, 1/3, . . .

Equation 2-9

The RF and LO signals can combine in a way to create so called subharmonic signals at the output of the mixer that are at frequencies of $f_{\rm IF}/2$, $f_{\rm IF}/3$, etc. If not adequately filtered, the harmonics of these subharmonic signals may fall at the desired IF frequency. Harmonics can be generated in nonlinear stages in the IF such as an IF Amplifier. **Figure 2-10** shows this process.

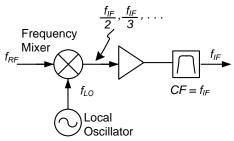


Figure 2-10. Subharmonics in the IF

Take the example where $f_{IF} = 100$ MHz, and $f_{LO} = 1100$ MHz. In this case, an RF signal at 1050 MHz will mix with the LO to create a 50 MHz response at the IF port. A second harmonic of an IF amplifier is 100 MHz, which propagates through the IF filter and records as a response. **Figure 2-11** shows the spectrum for the scenario described here.

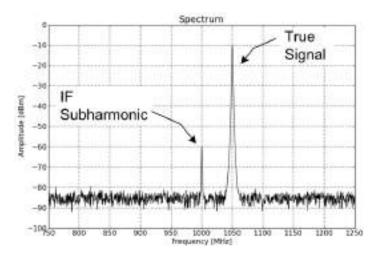


Figure 2-11. Spectrum Showing IF Subharmonic at a Displayed Frequency of 1000 MHz

In this example, the true signal is at 1050 MHz. However, when the analyzer is tuned to 1000 MHz, this 1050 MHz signal mixes to the IF such that a subharmonic is created. Harmonic distortion of the IF chain multiplies the subharmonic to the IF where it is displayed as an unwanted spur.

General M,N Spurs

Considering other M,N combinations within the context of the mixing equation **2-6**, many other spurious responses can be generated. Careful design of the receiver can try to minimize their amplitudes or better



yet filtering can be added to remove these spurs completely. However, the most careful design will still have conditions that can lead to the general spurious response.

Luckily, the amplitude of the spurious signal falls as a function of the spur order as shown in the following:

Spur Order =
$$|M| + |N|$$

Higher order spurs, in general, have lower amplitudes than lower order spurs. This at least bounds the problem of infinite number of possible spurs, since the vast majority will fall below the noise floor of the signal analyzer and will never be seen.

Consider the case of the 2,-1 spur created from a 600 MHz signal at the RF port. Again, the IF is centered at 100 MHz. When the signal analyzer is tuned to 1000 MHz, the corresponding LO frequency is 1100 MHz. The mixer equation with M,N = 2,-1 is satisfied as shown by the following:

 $2xf_{RF} - f_{LO} = f_{IF}$ 2x600 MHz - 1100 MHz = 100 MHz

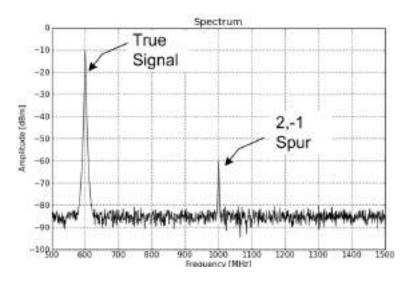


Figure 2-11 shows the spectrum of this scenario.

Figure 2-12. Spectrum Showing the 2,-1 Spur of a 1000 MHz Input Signal

When the signal analyzer is tuned to 600 MHz, the true response is displayed. However, when the analyzer is tuned to 1000 MHz, the 600 MHz signal generates an M,N = 2,-1 spur that is displayed at 1000 MHz.

There are also some redundant spurs. For example, consider the frequencies used to generate the IF subharmonic distortion in **Figure 2-11**: $f_{RF} = 1050$ MHz, $f_{LO} = 1100$ MHz, and $f_{IF} = 100$ MHz. The LO frequency corresponds to a tune frequency of 1000 MHz. The M,N = -2,2 spur is:

 $-2xf_{RF} + 2xf_{LO} = f_{IF}$ -2x1050 MHz + 2x1100 MHz = 100 MHz



The spectrum of this situation is exactly as depicted in **Figure 2-11**. So, a single RF signal is now responsible for multiple spur mechanisms.

IF Feedthrough

When the frequency at the RF port equals the IF port frequency, the RF signal can leak through the mixer, circumventing the mixing process. In other words, independent of the LO frequency, the RF signal is present at the IF port. **Figure 2-13** shows this spurious mechanism in the block diagram.

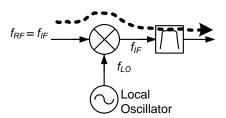


Figure 2-13. Block Diagram View of IF Feedthrough

The term given to this type of spur is IF feedthrough. Traditionally, this spur was also known as baseline lift as it manifested in a dramatic increase in the apparent displayed noise floor.

Using 100 MHz as the IF frequency, if the RF frequency is 100 MHz, no matter where the LO frequency is tuned, the spur will be present.

Signal analyzer structures put in place filtering to remove as much of the IF feedthrough as possible. However, filters do not have infinite rejection and some of the IF feedthrough does leak into the final IF. The performance specification given to the amount of suppression of the IF feedthrough signal is termed *IF rejection*.

LO Feedthrough

Using the mixer equation, when the tune frequency is set to 0 Hz, the LO frequency is set to the IF frequency. Using the IF and LO frequencies from the previous examples, the following shows this:

$$-f_{RF} + f_{LO} = f_{IF}$$

0 MHz + 100 MHz = 100 MHz
or, $f_{LO} = f_{IF}$

When this situation occurs, the LO signal can leak though the mixer and appear at the IF. Since these frequencies line up, the result is a spurious response displayed at DC. **Figure 2-13** shows this leakage path on the simplified signal analyzer block diagram.

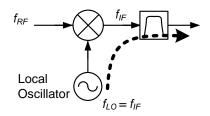






Figure 2-14 shows how LO leakage is manifested on the spectrum display as a response at a tune frequency of zero Hz.

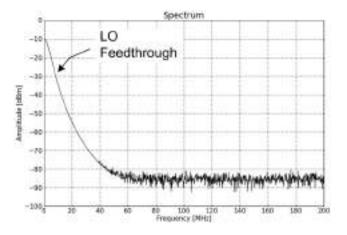


Figure 2-14. Spectrum Showing LO Feedthrough

LO Emissions

LO emissions is the term used to describe the LO signal leaking back through to the RF port of the mixer and showing up as an output signal at the RF input port of the receiver. This leakage path is shown in **Figure 2-15**.

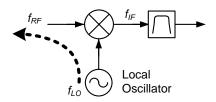


Figure 2-15. Block Diagram Showing LO Emissions.

Although not a spurious response per se, a signal being emitted from the input port of a receiver is normally an unexpected occurrence. In some measurement situations, this unexpected signal going into the output port of the device under test (DUT) may be undesirable. For spectrum monitoring where an antenna is connected to the input port, this leakage signal has now made a beacon out of the signal analyzer.

Residual Responses

Besides the LO, there are other sources of spurious signals in the signal analyzer. Reference oscillators for the LO phase lock loop (PLL), switching power supplies, and calibration signals can all generate signals. These signals can find various paths to the final IF, either through leakage in the signal paths or



even coupling between control and power supply routing traces. Once this signal falls into the final IF, a response is recorded. **Figure 2-16** shows one potential leakage path: one of the LO PLL reference oscillator signals.

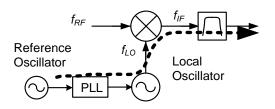


Figure 2-16. Residual Responses

These internally generated responses can occur when there is no signal present at the RF input port of the signal analyzer. These signals are termed *residual responses* and are usually specified with the RF input port terminated in the signal analyzer's characteristic impedance (normally 50 Ohms).

Cautionary Note on Nomenclature

Some of the spurious response terms used in the context of the super-heterodyne receiver are also used in direct conversion receivers, yet have different meanings. Image response and LO feedthrough are two of the terms that are used in both architectures, yet are defined differently. For the direct conversion receiver **Figure 1-10** the I and Q paths must be matched in amplitude response and their relative phases must be separated 90 degrees. Unbalance between relative amplitudes and phases between these two path results in spurious responses. The IF of the direct conversion receiver is split into two halves. The LO can leak through and cause a spurious response at the center of the IF. This is termed LO feedthru. A signal present in one half of the IF can result in an amplitude suppressed version of itself in the other half of the IF. This is termed image response. One should be aware that these terms have different meanings for the two different receiver architectures and care should be applied when comparing instrument specifications between these receiver architectures.



3. Super-Heterodyne Signal Analyzer Structures

Section 2 discussed the mixing process and many of its associated spurious response mechanisms. In this section, some of the more popular super-heterodyne systems will be shown. In these systems, attempts are made to minimize the impact of spurious responses. This section will concentrate on the frequency of the signal as it progresses though the signal analyzer, and later sections will concentrate on the amplitude of the signal.

Single Conversion Stage Structure

The most basic structure is the so called *single stage downconverter*. **Figure 3-1** shows the block diagram of this structure.

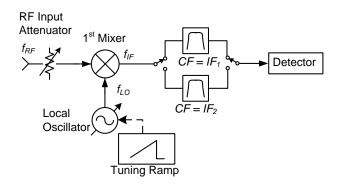


Figure 3-1. Single Stage Downconverter Block Diagram.

In the single stage downconverter, only one LO and one mixer stage are used to convert the RF input signal to the final IF. The final IF frequency is, in most cases, not at baseband, meaning the final IF is bandpass filtered rather than lowpass filtered. The purpose of the multiple filters in the IF is to suppress the IF feedthrough response. If IF_1 is the default path, at the tune frequency that equals the IF_1 frequency, then IF_2 path is selected.

The one major spur that is not suppressed is the image. Due to the lack of image signal suppression, the intent for this structure is not for general purpose applications. This structure is especially not appropriate for over-the-air (antenna connected) measurements where the spectrum contains unknown signals. However, this receiver is quite well suited for manufacturing test applications. In these applications, the signal under test is known in frequency and the connection to the DUT is normally though a shielded cabled which greatly attenuates the reception of unknown signal. Both the known frequency nature and the shielded connection make deciphering the signal analyzer spurious responses from the true response very predictable.

The advantage of this structure is its compactness and cost. Requiring a single LO and a single mixer results in a relatively simple design which tends to drive down both size and cost. In some applications, multiple downconverters are required and often phase coherence must be maintained. One way of ensuring phase coherence, meaning all receivers have known and predictable relative phase versus frequency behavior, is to share the LO where one downconverter is the master (LO source) and the downstream downconverters are the slaves (consumers of the LO). In such systems, routing only one LO signal via coaxial cables between downconverters is much simpler and less costly than systems with multiple LOs and mixer stages.



Multiple Conversion Lowband Structure

The multiple conversion structure attempts to overcome the lack of image rejection present in the single stage structure. The multiple stage structure is broadly divided into two different types: lowband and highband.

Figure 3-2 shows the block diagram of the lowband multiple stage super-heterodyne receiver. This architecture is designed to accept RF input frequencies ranging from near DC to RF. Typical maximum RF frequencies are 3 GHz to 7 GHz. True DC is impossible as the double balanced mixer used for the first conversion stage cannot accept more than a few hundred kilovolts without risk of damage. However, operation down to a few Hz is possible as long as the DC component is small.

The first LO has variable frequency that is configured to up convert RF signals to a fixed frequency first IF. The remaining LOs are fixed in frequency and are used with the later frequency conversion stages to progressively down convert the first IF to a final IF that can efficiently be detected.

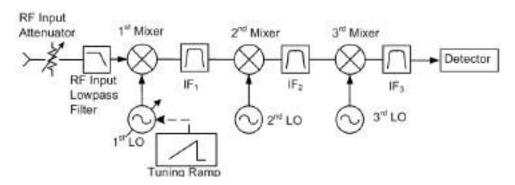


Figure 3-2 Triple Conversion Lowband Structure Block Diagram.

The main features of the lowband structure are:

- High image response suppression
- Low IF feedthrough response
- Low LO emissions

The first IF placed higher than the maximum input RF frequency enables all of these features. The first converter mixing equation governing this structure is shown in **Equation 3-1**.

$$-f_{RF} + f_{LO} = f_{IF}$$

Equation 3-1

With a high first IF, **Equation 3-1** implies that the LO frequency range is above both the IF and the maximum RF input frequency. **Figure 3-3** shows the relationship between the RF, LO and IF frequency ranges.



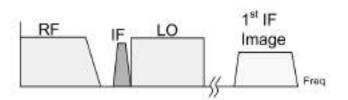


Figure 3-3. Spectrum of Lowband Frequencies

As an example, if the input RF frequency range is near DC to 3.6 GHz, and the IF is 4.6 GHz, the resulting LO frequency range is 4.6 GHz to 8.2 GHz.

Recall that the image responses are separated in frequency by twice the IF. Using the above frequency range values, the image of the first IF ranges from 9.2 GHz to 12.8 GHz.

In **Figure 3-2**, the lowpass filter in the RF input path is responsible for reducing the spurious signal content resulting from the following three mechanisms:

- 1. First IF image. The images are pushed far enough out in frequency in relation to the input RF frequency range that the lowpass filter can easily attenuate signals in the image frequency range.
- First IF feedthrough. The challenge for the lowpass filter is to minimize the attenuation of the RF signals in the RF input range and to provide a great deal of attenuation at the first IF frequency. The ratio between the first IF frequency and the maximum RF input frequency governs the lowpass filter complexity.
- 3. First LO emissions. Since the first LO frequency range is above the RF input frequency range, the LO signals fall into the stopband of the input lowpass filter and cannot pass on out through to the RF input port.

Why are at least three frequency conversion stages required? The answer is that the first IF frequency is too high to directly down convert to the final IF in one hop. At a 4.6 GHz first IF, using the above example, it is too difficult to attenuate the most challenging M,N spur: the 2,2 spur (actually M,N = -2,2 or 2,-2) generated in the latter mixer stages. This spur is present at one-half the IF frequency away from the input frequency. With low final IF frequencies, this places an unusually high burden on the complexity of filters possessing higher center frequencies. Multiple conversion stages are required to progressively mix the high first IF frequency down to the low final IF frequency.

The filters in the first and second IFs are carefully designed to mostly attenuate the challenging 2,2 spur as well as images and LO emissions. Inter-stage LO emissions can lead to residual responses: the LO and harmonics of the LO from a following stage can leak into the preceding stage. These LO harmonics can mix, possibly converting to one of the IF frequencies. Once converted to one of the IFs, this becomes an unwanted spurious response.

Multiple Conversion Highband Structure

Extending the lowband idea to higher input frequencies requires ever higher first IF and first LO frequencies. Challenges with filtering, dynamic range, first LO phase noise and above all cost become relevant at higher frequencies. The highband structure is normally used in conjunction with the lowband structure. The highband's minimum frequency starts at the lowband's maximum frequency.



The block diagram of the highband structure is shown in Figure 3-4.

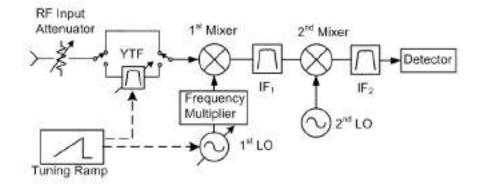


Figure 3-4. Highband Structure Block Diagram.

The main difference between the lowband and the highband structure is the tunable bandpass filter in the RF input path. At present the most common tunable filter type used in signal analyzers is the YIG tuned filter (YTF). YIG is a ferromagnetic material whose resonant frequency is directly proportional to an applied DC magnetic field [⁵]. The YTF exploits the properties of YIG creating a tunable bandpass filter. YTFs can tune from a few GHz to greater than 50 GHz with bandwidths ranging from a few 10's of MHz to a few 100's of MHz.

Because of the relatively narrow bandwidths, the YTF is very effective at attenuating images of the first IF and the very challenging 2,2 spurs. Instead of requiring a very high first IF frequency as in the case of the lowband structure,

the highband structure uses a relatively low first IF frequency. Typical first IF frequencies of highband structures in signal analyzers range from 300 MHz to 800 MHz, which is always below the minimum tune frequency of the highband path.

Figure 3-5 shows the spectrum of a highband structure with a first IF of 600 MHz. In this case, the signal analyzer is tuned to 4 GHz. The YTF also tunes to the center frequency of the signal analyzer, allowing a 4 GHz signal to pass. The YTF frequency response is represented by the dotted line in **Figure 3-5**. A signal at twice the IF (2 x 600 MHz) below the center frequency, would be considered an image if unfiltered. However, the narrow bandwidth of the YTF quite easily filters out this potential image response.



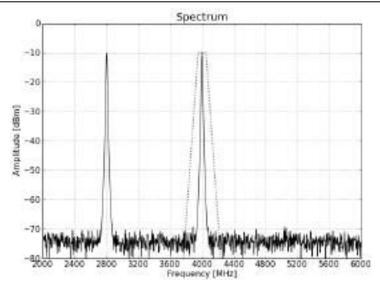


Figure 3-5. Image Suppression in the Highband Structure.

An added benefit of the highband structure is fewer conversion stages. The first IF frequency is low enough that normally only two conversions are required. Most often, the first IF frequency of the highband structure matches the second IF frequency of the lowband structure, allowing sharing of the final conversion stages. Further benefits provided by the YTF are attenuation of both the LO emissions and IF freedthrough.

However, the YTF is not ideal in all measurement circumstances. The YTF uses open loop tuning, meaning that the center frequency of the bandpass response is prone to tuning drift. This center frequency instability translates to relative amplitude and phase shift when the YTF is tuned to a constant frequency. For vector signal analysis, the lack of phase and amplitude stability due to the YTF normally precludes its use for the analysis of digital modulation metrics.

Another problem with the YTF may be the bandwidth. At lower tune frequencies, the YTF bandwidth is usually limited to 100 MHz, sometimes less. Modern communication formats have bandwidths that may be wider than the signal being measured.

To overcome the YTF's degradation of digital modulation analysis performance, the YTF, can be bypassed in some structures. However, once the YTF is bypassed the signal analyzer is now exposed to many of the spur mechanisms that the YTF prevents: images of the first IF, first IF feedthrough and first LO emissions. In a controlled environment, these spurs may not be an issue. Using the YTF path to decipher actual signals from spur signals before applying the bypass can at least reduce the chances that the measured signal is in fact not a spurious response.

Multiple Conversion Block Converter Structure

The block converter is one more structure; however, this structure is not commonly seen in most signal analyzer structures. The block converter's block diagram is shown in **Figure 3-6**.



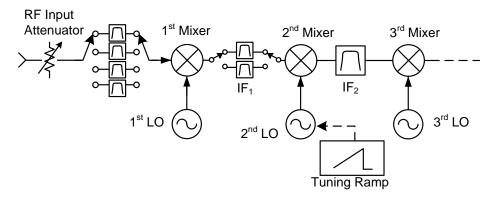


Figure 3-6. Block Diagram of the Block Converter Structure.

The main feature of this structure is that the first LO is fixed in frequency and the first IF and second LO frequency is variable. Often a lowband structure will be constructed and then a block converter will be placed in front of the lowband structure. The block converter, which usually operates at high RF and microwave frequencies, downconverts a segment, or block, of spectrum to the lowband frequency range.

Final IF Frequency Selection

Almost all modern signal analyzers rely on the Analog-to-Digital Converter (ADC) to digitize the final IF signal. Choosing the IF frequency and determining the parameters of the final IF filter are completely dependent on the ADC's sample clock rate. A very detailed analysis of the ADC can be found in reference [⁶]. Only the ADC as it relates to the signal analyzer is discussed here. **Figure 3-7** shows the relevant pieces of the ADC system.

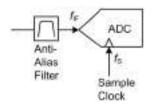


Figure 3-7. Analog-to-Digital Converter Components

The anti-alias filter (AAF) limits the bandwidth of the input signal before it enters the ADC. The filtered signal is sampled for a very short duration once every $1/f_s$ seconds, where f_s is the sample clock frequency. This sampling is depicted in **Figure 3-8**.

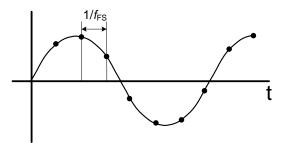




Figure 3-8. Sampling in the Time Domain

The *Nyquist theorem* states that in order to accurately reconstruct a signal using digital sampling, the sample rate must be two times the frequency of the input signal [⁷]. As long as the bandwidth of the input signal can be constrained, the Nyquist theorem can be violated. This is referred to as bandpass sampling. Bandpass sampling is best explained in the frequency domain. See **Figure 3-9**.

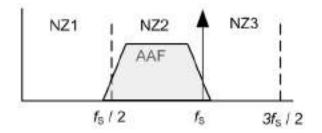


Figure 3-9. Bandpass Sampling in the Frequency Domain

Nyquist zones (NZ) are segments of the frequency spectrum, each with a range that spans one-half the ADC clock sample rate. As long as the input signal is band limited to be within one Nyquist zone, then aliasing will not occur. Aliasing can be explained in terms of the frequency domain representation of sampling. After sampling, the signals in any Nyquist zone appear in Nyquist zone one (NZ1) as shown in **Figure 3-10**.

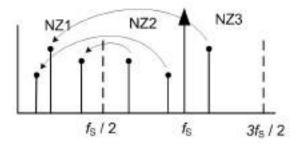


Figure 3-10. Signals Sampling Down to Baseband.

Nyquist zone one is often referred to as baseband. Failure to constrain the input signal frequencies to one Nyquist zone leads to ambiguity in the baseband. For example, if NZ2 is defined as the valid Nyquist zone, a signal present in NZ3 can appears baseband as well as the valid signals in NZ2. This creates ambiguity in trying to decipher the valid signal that appear in NZ1 from the invalid signals. The anti-alias filter as shown in **Figure 3-9** performs the task of constraining the input signal to one Nyquist zone NZ2 in the example shown in **Figure 3-9**.

Fifty percent of the ADC sample rate is the theoretical maximum bandwidth, however ractical constraints on analog AAF filter design limit the maximum bandwidth to approximately 40% of the ADC sample rate.

Variable Bandwidth Final IF Filters

The distinction between vector signal analysis and spectrum analysis was introduced in section 1. These terms, in the context of this document, refer to whether the final IF bandwidth is wide enough to capture



the entire modulation bandwidth of the measured signal. For spectrum analysis, the final IF is intentionally narrow banded to enhance dynamic range performance. This subject is explained further in section 6.

Figure 3-11 shows a possible implementation of the narrowband IF filters in the signal analyzer chain.

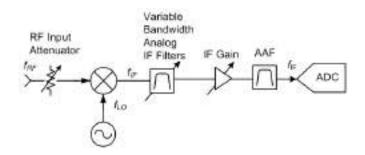


Figure 3-11. Signal Analyzer Block Diagram Highlighting the Analog IF Filters

The variable bandwidth filters can either be a bank of fixed tuned filters that are switched or it can be a single filter with tuning elements. In either case, ideally these filters are placed in the signal chain such that as many devices as possible are prevented from being exposed to the fundamental signals when the distortion components are being measured.

Some signal analyzers do not possess narrow band IF filters. These signal analyzers are only capable of making vector signal analysis measurements.



4. RF Chain Signal Processing

In this section, an analysis of the signal analyzer elements used in controlling the signal amplitude is presented. This is the RF chain signal processing, which does not include the processing of the final IF signal. The next section will discuss the IF signal processing. As part of the RF chain, LO phase noise will also be discussed in this section.

Amplitude Representation in Signal Analyzers

Even though the amplitude units for the signal analyzer can be in terms of power, the signal analyzer is actually a voltage measuring instrument. Post processing converts measured voltage to a variety of amplitude scales, some power based and some voltage based.

Power units by convention, are based on the mean squared value of a sinusoid as shown in the following:

$$P = \frac{V_{RMS}^2}{Z_o} Watts,$$

where V_{RMS} is the root mean square voltage of a sinusoid and Z_o is the nominal input impedance of the signal analyzer (typically 50 Ohms).

Figure 4-1 shows how the root mean square value of a sinusoid is related to the peak amplitude by the following:

$$V_{RMS} = \frac{V_{peak}}{\sqrt{2}}$$

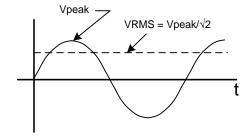


Figure 4-1. Sinusoid RMS Value

Power is expressed in decibels on a log scale, exploiting the compressive nature of logarithms. Further, the most common normalization is the milliwatt, resulting in log units of dBm, dB relative to 1 milliwat. One milliwatt corresponds to 0 dBm and one watt corresponds to +30 dBm For a signal analyzer whose nominal RF port impedance is Z_o , the conversion from V_{RMS} to log power in units of dBm is given in **Equation 4-1**.

$$P\left[dBm\right] = 10 x \log_{10}\left[\frac{V_{RMS}^2}{Zo * 0.001}\right]$$

Equation 4-1

Other amplitude units may be dBuV (dB relative to a microvolt) and dBmV (dB relative to a millivolt) as shown in **Equation 4-2**.



dBuV = dBm + 90 + 10log(Zo); for Zo = 50 Ohms, dBuV = dBm + 107dBmV = dBuV - 60

Equation 4-2

The amplitude units differ between oscilloscopes and signal analyzers. Oscilloscopes tend to have high impedance inputs so as to have minimal effect on the voltage being measured. Signal analyzers are intended to be power measurement devices. Furthermore, the impedance of the signal analyzer is intended to match the source impedance of the device being tested. This is normally 50 Ohms for most RF applications and 75 Ohms for cable television (CATV) devices.

RF/IF Path Amplitude Control Elements

Figure 4-2 shows a very simplified block diagram of the super-heterodyne signal analyzer highlighting the amplitude control elements. In most signal analyzers there are many stages of variable gain IF amplification, but here they are all represented with one element.

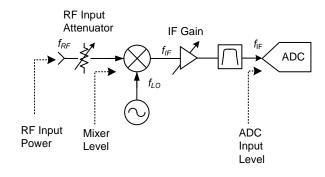


Figure 4-2. RF/IF Signal Path Amplitude Control

RF input power is the total signal power expressed in dBm at the input port of the signal analyzer. If the signal is a CW tone, then this is RMS power of the signal. If the signal contains modulation, then this power is computed by integrating average power over the modulation bandwidth. For multiple signals, this is the average power summation of all signals.

The *RF input attenuator* can be directly controlled by the operator. The amount of attenuation varies between 0 dB and some maximum value (ranging between 50 dB and 70 dB). Resolution varies between 1 dB and 10 dB.

Mixer level is related to the input power of the first mixer by:

Mixer Level = RF Input Power – RF Input Attenuator setting

For instance, if the RF input power is -10 dBm and the RF input attenuator setting is 10 dB, the mixer level is -20 dBm. This term is not to be taken too literally as it does not include the effects of frequency response between the RF input port and the mixer. The true power incident on the first mixer may be slightly off from -20 dBm due to frequency response, but the true power is not the definition of mixer level.



ADC input level is the average power expressed in dBm of a CW tone at the input of the ADC or digitizer. *IF output power level* is also a term used to express this value.

Reference Level and Gain Setting Equations

Reference level refers to the maximum power level of an input CW tone that can be measured without overdriving the IF backend of the signal analyzer. Reference level is also the name given to the vertical axis control of the signal analyzer display. In most signal analyzers, the reference level is the top graticule on the vertical display axis. The user has direct control of this reference level setting. When a CW signal at the RF input has an amplitude equal to the reference level setting, the signal analyzer's internal gains are configured such that the signal level at the ADC is maximized. Letting the input signal power level exceed the reference level value runs the risk of gain compressing component in the RF frontend. Gain compressing the frontend leads to more distortion and degradation of amplitude accuracy. The other risk at letting the signal level exceed the reference level setting is that the ADC can be overdriven. Overdriving the ADC results in distortion so dramatic that no information of the signal can be recovered. Most signal analyzers do not have warning mechanisms for frontends being overdriven. However, an ADC being overdriven often results in a warning such as "IF Overload."

The gain setting parameters that the user can directly control are as follows:

- Reference level
- RF input attenuation
- Maximum mixer level
- IF output level, which is the same as the ADC input level

Maximum mixer level defaults to a value such that the mixer level for any given reference level and input attenuator combination is at least 10 dB to 15 dB below the 1 dB gain compression level of the RF frontend. IF output level is normally set 3 dB to 10 dB below the full scale input level of the ADC. Most often there are default values for the maximum mixer level and IF output level that are optimum. Exposing these parameters allows the experienced user to experiment with measurement specific custom settings.

IF gain as shown in **Figure 4-2** is not directly controllable by the user; however, it is indirectly controlled via the reference level setting. The gain setting equations are constructed such that the displayed amplitude represents the CW signal amplitude at the RF input port. When the RF attenuator is *auto-coupled* to the reference level, the gain equations are as follows:

RF Input Attenuation setting = Reference Level – Maximum Mixer Level Signal Analyzer Gain = IF Output Level – Reference Level IF Gain = Signal Analyzer Gain + RF Attenuation Setting

Often the RF attenuator under auto-coupled conditions has a minimum value (5 dB to 10 dB) so that the user does not inadvertently allow the attenuation to become low enough such that there is risk of damaging the front end with too high input power. For manual RF attenuator control, the Signal Analyzer Gain and IF Gain equations still apply.



The IF gain also compensates for frontend frequency response. Recall that the signal analyzer tries to faithfully reflect the power of the input signal. However, the amplitude response of the components in front of the first mixer and the first mixer itself are not flat with frequency. Compensation for frequency response using IF gain will tend to

display the noise floor as an inverse of the frequency response curve. As shown in **Figure 4-3**, roll-off in the frontend frequency response results in roll-up of the noise floor after gain compensation in the IF.

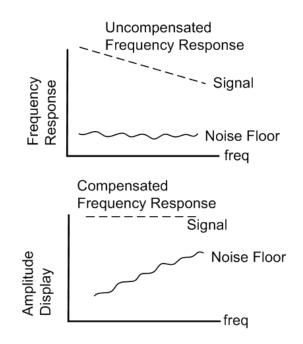


Figure 4-3. Frequency Response Compensation

Mixer Level Effect on Frontend Distortion

Regarding first mixer M,N spurs, lowering the incident power at the mixer changes the distortion and amplitude as a function of the mixing equation's RF signal multiplier, M. Every dB decrease in the mixer level decreases the spur amplitude by the absolute value of M in dB. For example, for a spur with an M multiplier whose absolute value is 2, a 10 dB drop in the mixer level lowers the spur amplitude by 20 dB. Conversely, an increase in mixer level raises the spur amplitude by [M] dB.

Mixer level is varied by changing the power level of the signal being measured or by changing the RF input attenuator setting. Changing the test signal's power is most often not very convenient, which leaves the RF input attenuation as the primary means of improving the signal analyzer's distortion performance.

Mixer Level Effect on Frontend Noise

The noise performance specification for the signal analyzer is the *average noise level*, or sometimes DANL for *displayed average noise level*. DANL is normally specified for an RF input attenuation value of 0 dB. As the RF input attenuation is increased, the signal-to-noise ratio (SNR) decreases. This is shown in **Figure 4-4**.



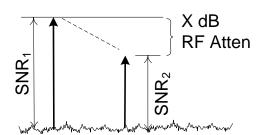


Figure 4-4. Signal-to-Noise Ratio at Two Different RF Input Attenuator Settings.

The signal to the left in **Figure 4-4** is with one RF input attenuator setting and the signal to the right is with X dB more RF input attenuation. At the input of the first mixer, no amplification has taken place, so the noise floor, no matter the RF input attenuator setting, is kTB or -174 dBm/Hz. Thermal noise, or kTB noise, is the noise power generated by a 50 Ohm resistor under the condition that this source resistor is attached to a matched impedance load. This noise power is normalized to a measurement bandwidth, B, of 1 Hz. With entirely passive loss elements in the signal chain, the noise power, assuming matched source/load conditions, can never drop below the kTB level.

Because the noise floor cannot drop below kTB, as the signal is attenuated, the SNR drops dB per dB of attenuation value. However, a constant noise floor is not what is recorded. The signal analyzer is calibrated to try to ensure that the measured signal level at the input port is accurately portrayed. When the RF attenuator is increased, the signal level at the input port does not change, only the signal level at the input of the first mixer. IF gain is increased to compensate for RF attenuation, which increases the displayed noise floor. **Figure 4-5** shows the two signals from **Figure 4-4** as they would appear on the display. The signal powers are now the same, but the noise floor associated with the right-hand signal is X dB higher than the noise associated with the signal on the left. Compensating IF gain accounts for both equal displayed signal levels and noise floor offsets between the two measurement conditions.

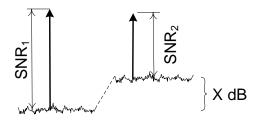


Figure 4-5. Display Adjusted for RF Input Attenuator Setting.

In regards to the analog frontend of the signal analyzer, increasing the RF input attenuation value lowers the amplitudes of the distortion components generated by the signal analyzer. However, increasing the RF input attenuation value also degrades the signal analyzer's frontend noise performance. A balance must be made between the signal analyzer's noise and distortion performance using the RF input attenuator setting. Section 6 is entirely dedicated to describing this tradeoff.

ADC Dynamic Range

The ADC used at the very end of the signal chain has a very different optimization for dynamic range. **Figure 4-6** shows the ADC SNR and spurious-free dynamic range (SFDR) versus ADC input level.



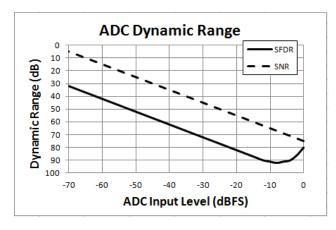


Figure 4-6. ADC SNR and SFDR versus ADC Input Level

For both ADC SNR and SFDR, these values have their highest values near the maximum ADC input power level. Maximum ADC input power level is termed full scale input level. The x-axis in **Figure 4-6** is dB relative to full scale input level (dBFS).

Consequently, two optimizations must take place in the signal analyzer. The RF input attenuator must be used to set the optimum mixer level for best tradeoff between frontend generated noise and distortion. The IF gain must then be set such that the signal at the end of the signal analyzer's IF chain is close to the ADC's full scale input level. More on this optimization technique in section 6.

Preamplifier

The preamplifier is placed in front of the first mixer in the signal analyzer. Most often, but not always, the preamplifier is placed after the RF input attenuator. **Figure 4-7** shows the block diagram of the signal analyzer with the preamplifier present. Note that the preamplifier path is either selected or bypassed.

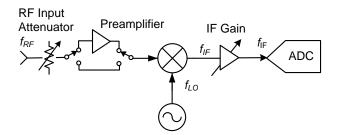


Figure 4-7. Preamplifier in the Signal Analyzer Block Diagram

The purpose of the preamplifier is to lower the noise floor of the signal analyzer. Caution must be exercised when enabling the preamplifier path as it does add gain in front of the first mixer. The maximum input power level must be lowered by the preamplifier gain value. Also, the preamplifier will generate harmonic distortion and intermodulation distortion. In most cases, this distortion degrades more than the improvement in noise floor, resulting in degraded signal analyzer performance. However, if the application relies on improved noise performance alone, enabling the preamplifier is very effective.



Explaining how the preamplifier improves system noise figure adds some insight for low noise measurements as well as gives some foundation in the event that the user may want to add external preamplification for further noise performance improvement.

Noise figure is the logarithmic representation of noise factor, *f*, as shown in **Equation 4-3**.

Noise Figure
$$[dB] = 10 * log_{10}(f)$$

Equation 4-3

Noise factor is defined for a two-port device and is the ratio of input SNR to output SNR [⁸] as shown in **Equation 4-4**.

$$f = \frac{SNR_{in}}{SNR_{out}}$$

Equation 4-4

Using the nomenclature in Figure 4-8, the noise factor, shown in Equation 4-4, can be reduced to Equation 4-5.

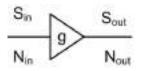


Figure 4-8. Terms Used to Describe Noise Figure.

$$f = \frac{\frac{S_{in}}{N_{in}}}{\frac{S_{out}}{N_{out}}} = \frac{N_{out}}{N_{in}g}$$

where S_{in} and S_{out} are the input and output signal powers respectively. N_{in} and N_{out} are the input and output noise powers respectively.

Equation 4-5

Gain, g, is defined as S_{out}/S_{in} . Input noise, N_{in} , is defined to be kTB. Normally, noise factor for a device is provided by the manufacturer or is a parameter that can be measured. Once noise factor is known, **Equation 4-5** can be solved for output noise.

The signal analyzer can be thought of as a two port device. The input is the RF input port and the output is the recorded amplitude data, which is often the display data. From this point of view, the signal analyzer can be considered to have a gain of 0 dB. The output amplitude is calibrated to equal the input amplitude. With G=1, which is the linear value corresponding to 0 dB, and expressing **Equation 4-5** in logarithmic terms, the signal analyzer noise figure is expressed as shown in **Equation 4-6**.

Equation 4-6



Output noise power is the average noise level specification given for the signal analyzer. Input noise is defined to be kTB or -174 dBm/Hz. So, as an example, if the average noise level specification for the signal analyzer is -154 dBm/Hz, the signal analyzer's noise figure equates to 20 dB.

To overcome the signal analyzer's noise figure, the preamplifier is placed in front of the signal analyzer as shown in **Figure 4-9**.

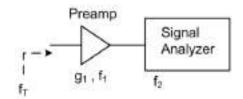


Figure 4-9. Cascaded Noise Figure

The system's total noise factor, f_{T} , follows the cascade noise figure equation [8]:

$$f_T = f_1 + \frac{f_2 - 1}{g_1}$$

Equation 4-7

If the gain of the preamplifier is sufficiently larger than the noise figure of the signal analyzer, as shown in **Equation 4-7**, the system noise figure approaches the noise figure of the preamplifier alone. Two important parameters define the preamplifier: low noise figure and high gain. In dB terms, to be effective, the gain of the preamplifier should be at least as large as the noise figure of the signal analyzer.

Phase Noise

As mentioned in section 1, phase noise is the frequency domain representation of jitter or phase fluctuation versus time. The signal analyzer contributes phase noise by means of its LO signals. Oscillators used in signal analyzers have an electronic signal to control the frequency of oscillation. Inevitably, noise will be present on the tune control which in turn implants phase modulation on the oscillator signal.

In the frequency domain, phase noise on an LO signal appears as shown in Figure 4-10.



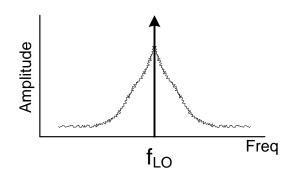


Figure 4-10. Phase Noise on an LO Signal.

A double-sideband pedestal of noise surrounds the LO carrier signal. Most often the phase noise power increases for frequencies close to the carrier.

When specifying phase noise, the phase noise is normally the single-sideband noise pedestal plotted versus frequency offset from the carrier. Single-sideband noise is shown in **Figure 4-11**.

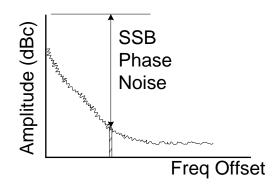


Figure 4-11. Single-sideband Phase Noise versus Offset Frequency.

Because of the variation with offset frequency, phase noise is specified at a particular offset frequency. Phase noise behaves like broadband noise in that phase noise power is a function of the measurement bandwidth. Phase noise is normalized to a 1 Hz measurement bandwidth resulting in units of dBc/Hz.

The frequency shift property of the Fourier transform, as shown in **Equation 2-1**, applies to phase noise. With phase noise superimposed on the LO signal, the LO's phase noise will implant on the IF signal after frequency mixing. If there is phase noise on the RF input signal from the DUT, the LO phase noise will add to the DUT's phase noise during the frequency mixing process.

The signal analyzer's phase noise performance determines whether or not certain measurements can be made. One measurement example is shown in **Figure 4-11**.



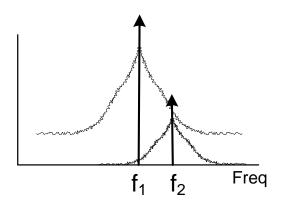


Figure 4-11. Phase Noise Masking.

In this example, a small amplitude signal at frequency f_2 , is located close in frequency to a large amplitude signal located at a frequency of f1. The large amplitude signal's phase noise pedestal potentially hides the small signal amplitude. This problem is especially prevalent in the measure of two-tone intermodulation distortion when the two tones are closely spaced in frequency. The resulting intermodulation distortion products, which can be quite small, can easily be buried in the phase noise pedestal surrounding the fundamental tones.

Another measurement where the signal analyzer's phase noise is important is in the measurement of the DUT's phase noise. This situation is shown in **Figure 4-12**.

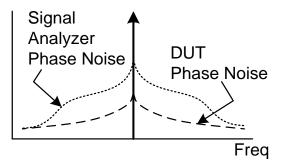


Figure 4-12. Signal Analyzer and DUT Phase Noise.

Often the signal analyzer is used to measure phase noise of the RF input signal from a DUT. If the signal analyzer's phase noise is larger than the DUT's, then this measurement is impossible.

The signal analyzer's phase noise at wide offset frequencies (greater than 1 MHz) is important in the measurement of adjacent channel power leakage ratio (ACLR) of digitally modulated communications signals. Section 6 will show how far offset phase noise can limit ACLR measurement performance.



5. IF Chain Signal Processing

IF Chain Signal Processing refers to the signal conditioning steps after the signal is digitized by the ADC. This includes realization of the *resolution bandwidth filters*, trace averaging, implementation of trace detector types and also FFT processing to convert the signal into frequency spectrum components. For spectrum analysis, only the magnitude versus frequency information is of interest. For vector signal analysis, both magnitude and phase are of interest. This section will concentrate on spectrum analysis IF chain signal processing.

Nearly all modern signal analyzers digitize the final IF signal and use digital signal processing (DSP) routines to perform IF chain signal processing. The DSP is performed in either dedicated integrated circuits (ASICs) or field-programmable gate arrays (FPGAs). Software on the host central processing unit (CPU) often augments the ASIC and FPGA processing. In some cases the entire signal processing is software oriented.

The digital signal processing algorithms, regardless of where the computation takes place, tries to mimic what was once done with analog hardware components. First, a brief review of the analog hardware implementation will be presented. Then, steps for mapping the analog signal processing components to the DSP will be discussed.

Analog IF Signal Processing

Figure 5-1 shows the block diagram of the IF signal processing section of pure analog signal analyzers. The main components are: the *resolution bandwidth filters, logarithmic amplifier, envelope detector,* and the *video bandwidth filter.*

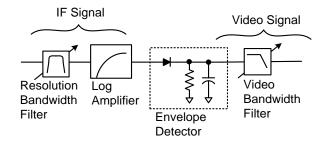


Figure 5-1. Analog IF Block Diagram

Resolution Bandwidth Filter

The *resolution bandwidth (RBW) filter*, in the context of the analog IF signal processing chain, is a variable bandwidth bandpass filter whose center frequency is the frontend's final IF. In most traditional topologies, the bandwidth is adjustable in a 1:3:5 progression, repeating once a decade: for instance 100 Hz, 300 Hz, 500 Hz, 1 kHz, 3 kHz, and so on. Bandwidths range from less than 100 Hz to greater than 1 MHz.

The RBW filter provides the means of distinguishing signals closely spaced in frequency. **Figure 5-2** demonstrates the resolving capability of the RBW filter.



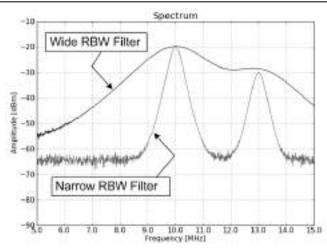


Figure 5-2. Resolving Power versus RBW Setting.

In **Figure 5-2**, with the narrow RBW filter setting, the two signals can clearly be distinguished. However, the wide RBW setting does not provide enough resolving power to separate these two signals.

Another feature of the RBW filter is its influence on the noise floor. In section 4, it was remarked that the thermal noise power is given by the equation $N_o = kTB$. The 'B' part is the measurement bandwidth. Since, the RBW filter is normally the narrowest filter in the entire signal analyzer, the 'B' can be regarded as being the bandwidth of the RBW filter. On a log scale, the noise power using two different RBW settings is given by:

$$N_{02} - N_{01} = 10 \log \left(\frac{RBW_2}{RBW_1}\right)$$

where RBW is the 3 dB BW of the RBW filter

Equation 5-1

Equation 5-1 demonstrates that a decade change in the RBW setting results in a 10 dB change in the noise floor. **Figure 5-3** shows the influence that the RBW setting has on the noise floor.

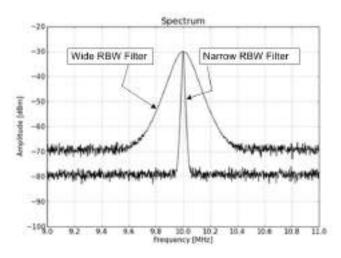


Figure 5-3. Noise Floor versus RBW Setting.



As demonstrated in **Figure 5-3**, the noise floor changes with RBW setting, but the amplitude of a CW signal does not. Therefore, signal-to-noise ratio is a function of RBW setting when measuring CW and narrow bandwidth signals.

Equation 4-6 gives the relationship between the signal analyzer's average noise level specification and the actual noise floor. Average noise level in units of dBm/Hz uses a 1 Hz RBW setting for this measurement. The average noise level generalized to any RBW setting follows **Equation 5-2**.

Signal Analyzer Noise Floor [dBm] = Average Noise Level specification [dBm/Hz]+ 10log(RBW) [dB]

Equation 5-2

For instance, if the average noise level specification for the signal analyzer is -150 dBm/Hz, the signal analyzer's noise floor in a 1 kHz RBW setting is:

Noise Floor [dBm] = -150 [dBm/Hz] + $10\log(1 \text{ kHz})$ = -120 dBm

Logarithmic Amplifier

Traditional signal analyzers commonly had a default display vertical axis calibrated in log power, and usually referenced to 1 mW [dBm]. Further, the default range was commonly 90 dB to 100 dB. Using **Equation 4-1**, if one were to try to view the voltages of two signals separated in amplitude by 100 dB, the voltage ratio would be 100,000:1. This voltage ratio would prove impractical for viewing on a single display. To accommodate viewing of large dynamic range on a single display some sort of signal amplitude compression is required. The logarithmic amplifier fulfils this need.

The *logarithmic amplifier*, or log amp, operates on the IF signal, whose voltage transfer function is logarithmic: Vout = $A^* \log(Vin)$, where A is a scaling factor. **Figure 5-4** shows an approximation to the log amp transfer function. From **Figure 5-4**, the compression that the log amp provides is clearly evident. A miniscule change on the input voltage translates to many dBs of change in the output. Simultaneously viewing two signals 100 dB or more apart in amplitude is possible with the log amp.

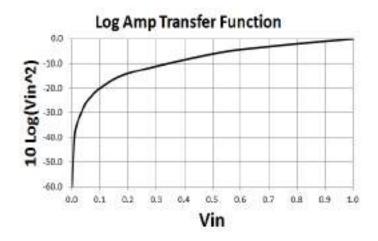


Figure 5-4. Log Amp Transfer Function.



Envelope Detector

Consider a stationary CW signal at the input of the signal analyzer shown in **Figure 2-7**. As the LO is swept, the resulting signal at the IF of the mixer is a sinusoid whose frequency is also swept. A swept frequency signal is sometimes called a chirp signal. As this chirped IF signal passes through the RBW filters, the amplitude of this signal has the shape of the RBW's frequency response. **Figure 5-5** depicts the IF signal as it exits the RBW filter.

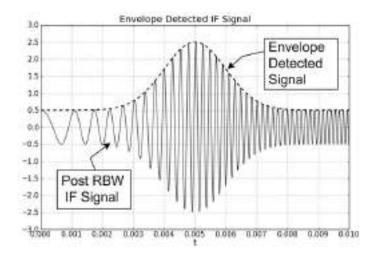


Figure 5-5. Time Domain View of IF Signal After the Resolution Bandwidth Filters.

The *envelope detector* is a simple diode rectifier, whose output traces only the peaks of the IF signal and is represented by the dashed line in **Figure 5-5**. Once detected, the waveform is referred to as the video signal. The detected signal's maximum frequency is no longer that of the IF signal, but rather the modulation riding on top of the IF signal.

Why convert to video? One answer is that is that traditional cathode ray tube (CRT) displays, which were driven directly by the video signal, could only respond to low frequency waveforms. More importantly, however, is that the information in a signal is its modulation characteristics. The central IF signal is merely a carrier for the modulation waveform. Stripping away this carrier still retains the useful information about the signal.

Video Bandwidth Filter

The other IF filter in the analog IF signal processing chain is the *video bandwidth (VBW) filter*. The VBW filter is a lowpass filter with user controlled bandwidth. As the name implies, this filter operates on the video signal, which is the modulation waveform, not the IF waveform. This distinction is important. Filtering the IF signal, which is the function of the resolution bandwidth filter, lowers the overall noise floor. The VBW filter reduces the variance of the noise. Another way of expressing it is that the VBW decreases the noisiness of the displayed noise.

Figure 5-6 demonstrates the effect of the VBW filter on the envelope detected trace. The lighter trace depicts the trace with a relatively wide VBW setting, whereas the dark trace represents the same signal with a narrower VBW setting.



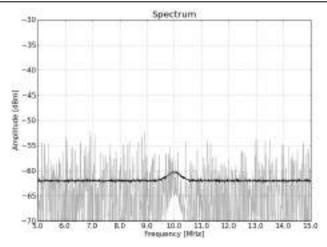


Figure 5-6. Signal with Two Different Video Bandwidth Filter Settings.

Note how the average value of the noise floor itself does not shift with VBW setting, just the fluctuation of noise. With a signal close to the noise floor, as shown in **Figure 5-6**, the signal is undetectable with the wide VBW setting. It is only with the trace smoothing provided by the narrower VBW setting can the near-noise signal be deciphered.

The VBW filter is implemented as a single pole resistor-capacitor (RC) lowpass filter whose corner frequency, or bandwidth, can be selected by the user.

Sweep Speed Considerations

Narrowing either or both the RBW filter setting and/or the VBW filter setting is not without consequences. Both filters have rather dramatic influence on sweep speed. In the time domain, the settling time of a filter grows longer as its bandwidth is decreased. When the IF signal's frequency matches the center frequency of the RBW filter, the RBW responds as if it were hit with an impulse in the time domain. A narrower RBW filter setting takes longer to respond to this impulse. Decreasing sweep time allows the filter to respond and settle to a final amplitude value.

The VBW filter has similar settling constraints in that the sweep speed of the signal analyzer must be slowed down to allow the filter response to settle.

Viewing Modulation

Figure 5-7 shows the frequency domain spectrum of a signal with amplitude modulation (AM)



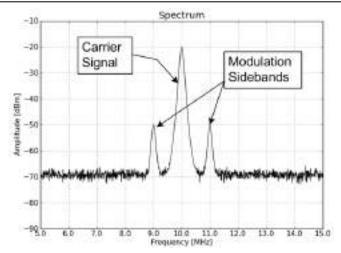


Figure 5-7. Amplitude Modulation in the Frequency Domain.

To resolve the AM sidebands in the frequency domain, the RBW filter setting must be narrower than the AM signal's modulation frequency. However, the signal analyzer can also be used to analyze the modulated signal in a time domain view. This is done with the signal analyzer's frequency span set to zero with the center frequency set to the signal's carrier frequency 10 MHz in the case of **Figure 5-7**. Zero span implies that the LO is not sweeping, but rather fixed at a constant frequency.

To view the modulation, the RBW filter setting must be wide enough to capture both the carrier and the modulation sidebands. In **Figure 5-7**, the AM rate is 1 MHz, so to capture both sidebands, the RBW setting would need to be at least 2 MHz. **Figure 5-8** shows the AM signal when the signal analyzer is set to zero span and the RBW filter setting is much wider than the modulation bandwidth of the test signal.

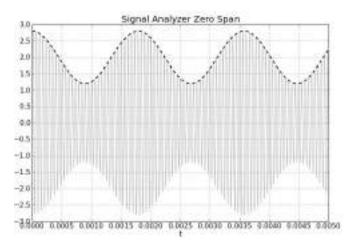


Figure 5-8. Amplitude Modulation in the Time Domain.

Note that the x-axis is time rather than frequency when in zero span.

The high frequency sine wave (light colored trace) in **Figure 5-8** is the carrier. The dashed line represents the modulation signal -- a sinusoid in this case. Envelope detection strips away the carrier, leaving just the modulation.



Detector Modes

At some point between the all analog and all digital IF evolution, the ADC was introduced to the analog IF to digitize the video signal (early ADCs did not have enough bandwidth to digitize the IF signal). Once implemented, a few features related to signal detection were introduced. The digitized video signal is quantized into a finite number of frequency and amplitude points – these are the data points eventually displayed. The data represented on these y-axis points is the subject of the detection modes. Some of these modes are as follows:

- Max peak: the maximum amplitude that occurs between the discrete frequency points, sometimes also referred to as the *peak detector*. With a relatively narrow RBW filter setting, the peak of the signal may occur between the discrete data points. Without the max peak detector, the signal data could potentially be lost.
- Min peak: sometimes also referred to as the *pit detector*. Same function as the peak detector, but records the minimum signal amplitude between the discrete frequency points.
- Sample: the amplitude data at the discrete frequency data point is recorded. This mode is the appropriate one for the measurement of noise.

Figure 5-9 shows the relationship between the continuous data and the detector modes.

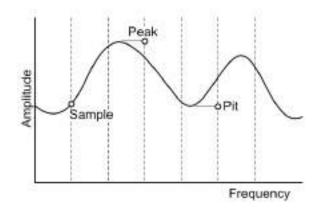


Figure 5-9. Analog IF Detector Modes

Challenges with the All Analog IF

Size and power are two downsides to the analog IF signal processing chain. The circuitry required to realize all the components is not small and certainly not power efficient.

However, the biggest mark against the analog IF is the amplitude accuracy. Among some of amplitude error terms associated with the analog IF are as follows:

- Log fidelity: the transfer function from Vin to Log(Vin) is not perfect. The wider the range of the amplitude scale, the larger this error.
- Bandwidth switching: changing either the RBW or the VBW filter settings results in imperfect relative amplitude accuracy between filter settings.
- Noise floor: the actual 3 dB bandwidth of the RBW filter has inaccuracies compared to the desired bandwidth setting. Since noise power is a function of the bandwidth, errors in the true bandwidth yield errors in the displayed noise floor.



• Temperature stability: holding all the analog circuitry stable over temperature is not an easy problem. No matter how well designed, the analog signal chain almost always yields amplitude error versus operating temperature.

IF Signal Processing with the Digital IF

Analog signal processing has largely been replaced with digital signal processing (DSP) algorithms. DSP techniques can now perform all the necessary signal conditioning of the IF signal. As mentioned earlier, the algorithms can either be implemented in ASIC, FPGA or in software; however, the fundamental DSP algorithms apply no matter which physical device is used. Due to the predictability of DSP algorithms, nearly all of the amplitude uncertainty error terms associated with the analog IF signal chain are simply not present in the digital IF implementation.

The DSP algorithms do vary amongst the different signal analyzer vendors. There are even differences amongst different products within the same vendor. This discussion will concentrate on the DSP routines used by National Instruments and even more specifically the routines associated with the NI-RFSA driver software used by many signal analyzers at National Instruments.

Another distinction is whether or not the end use is for spectrum analysis or vector signal analysis. Spectrum analysis uses magnitude versus frequency data, requiring a conversion from the time domain IF data to the frequency domain data. Vector signal analysis often uses only time domain data; however, the data is complex, containing magnitude and phase. Spectrum analysis is the main focus of this discussion.

Signal Processing Chain: Digital Hardware

In **Figure 5-10**, the block diagram of the signal processing chain portion that resides in the digitizer is shown. For more information, refer to Reference [⁹]. In the digitizer, the ADC samples the analog IF signal. As discussed in section 3, sampling signals that fall in any Nyquist zone appear as if they are in Nyquist zone one, the baseband. At the end of the processing, the frequency axis will be re-scaled to the appropriate analog IF Nyquist zone, but at this point in the processing chain, the digitized data is in the baseband as shown in **Figure 5-10**.

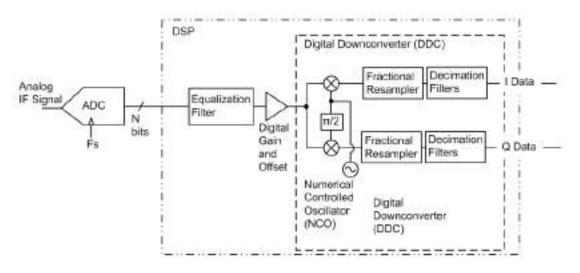


Figure 5-10. Digital Downconverter Block Diagram

The first block in the digital signal processing chain is the *equalization filter*. This is implemented as a finite impulse response (FIR) filter whose filtering coefficients, or taps, can be reconfigured for each separate data acquisition. The purpose of the equalization filter is to remove the magnitude and phase



response of the analog frontend components. The equalization is over one Nyquist zone, which implies that this filter removes magnitude and phase frequency

response over the IF bandwidth of the signal analyzer. Factory calibration data and/or signal analyzer self-calibration data corresponding to IF channel magnitude and phase is stored in the signal analyzer's memory. The signal analyzer's driver software uses this measured data to compute the FIR filter coefficients and applies them to the equalization filter.

The frontend will have some gain and offset errors mostly resulting from the finite amplitude step size of front end gain compensation circuitry. These offsets are compensated digitally on the equalized data.

Note that some of the frontend paths may not be able to be equalized due to a large amount of passband amplitude and phase ripple. Paths with SAW filers may fall into this category. In these cases, the number of FIR taps is too large for practical implementation in digital hardware. For magnitude only data, the equalization for these situations may take place entirely in software.

After the data is equalized and corrected for frontend gain offset, the data enters the digital downconverter (DDC). The DDC uses a digital implementation of an IQ demodulator to convert the real data to complex I and Q data streams. A *numerically controlled oscillator* (NCO) as part of the IQ demodulator is used to correct for finite frontend frequency errors. These frequency errors can result from the frontend LOs having limited frequency resolution.

In many modulation formats, the data must be processed with specific data rates. These data rates most likely do not match the sample clock rate of the ADC. Instead of adjusting the ADC's clock rate, which can lead to unwanted spurs and noise degrading sample clock signal, the digitizers in most National Instruments signal analyzers use a *fractional resampler* in the DDC. The fractional resampler effectively changes the data rate of the I and Q data streams.

When the requested bandwidth is less than the entire Nyquist zone, it is beneficial to remove the unwanted spectrum digitally. This procedure is accomplished with the decimation filters. These filters not only filter out the unwanted

spectrum, but also reduce the data rate. Lower data rate leads to more efficient signal processing in the chain following the decimation filters. The Instantaneous Bandwidth property in the NI-RFSA driver software influences the decimation filter bandwidth.

For measurements requiring complex data, modulation analysis for example, the I and Q data streams are exported to software. For real data analysis, magnitude only spectrum analysis for example, only the I Data stream is exported to software.

Signal Processing Chain: Software

For magnitude only spectrum analysis, the flow diagram of the signal processing that NI currently uses is shown in **Figure 5-11**. The National Instruments implementation of the software signal processing is housed within the NI Spectral Measurement Toolkit



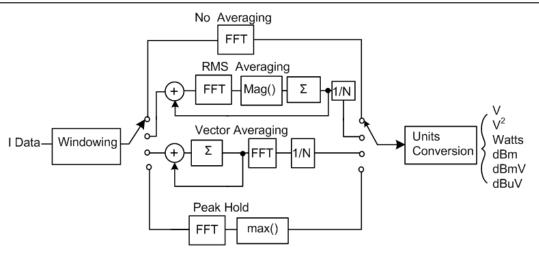


Figure 5-11. Spectrum Mode Signal Processing Block Diagram

Again, for magnitude only processing (spectrum analysis), only the real data from the hardware signal processing chain is required. This is the I Data digital bit stream. The blocks are: *windowing*, followed by *FFT/averaging*, and finally *amplitude unit conversion*. Each of these blocks will be discussed. But first, some background on FFT and spectral leakage is presented before leading into details on the windowing process.

Spectral Leakage in the FFT Process

The discrete Fourier transform (DFT) and the more computationally efficient fast-Fourier transform (FFT) operate on time domain data that has been sampled. The sampling process of converting the time domain signal, x(t), into the sampled signal, x(n), is shown in **Figure 5-12**:

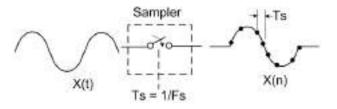


Figure 5-12. Sampling a Sinusoid Prior to FFT Processing.

The ADC acts as a momentary switch that outputs a sample of the incoming continuous time signal. The samples occur spaced at every 1/Fs in time, where Fs is the clock rate of the ADC. The sampling occurs over a finite period of time – this is the *data acquisition* duration. After the FFT processing, the frequency domain data also falls at discrete points. If there are N time domain samples, then there will be N frequency domain points. The frequency domain points are sometimes called *bins*.

The sampled signal appears as a pulsed signal where the samples have zero valued amplitude values outside the acquisition duration. From a signal processing perspective, the signal has rectangular pulse, or a rect(t), function superimposed on it. From **Figure 1-3**, a rectangular pulse in the time domain translates to a sinc function in the frequency domain.

For a sinusoid input signal, if its frequency is an integral multiple of Fs/N, then the frequency domain sinc function superimposed on the CW signal is as shown in **Figure 5-13**.



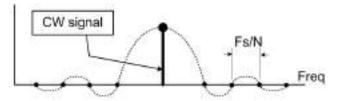


Figure 5-13. Spectrum of CW Signal <u>without</u> Spectral Leakage.

The sinc function is centered on the CW signal. The zero crossings of the sinc envelope fall on FFT frequency bins. Only the spectral line associated with the CW signal results.

Recall that all periodic signals can be deconstructed into a summation of sinusoids. Even if the signal is not periodic, FFT processing requires that the one data acquisition record be considered as periodic – that is, the FFT treats the one data acquisition as if it were replicated infinitely in time. So, the treatment of a single sinusoid can be generalized to any signal that is being processed with the FFT.

If the frequency of the sinusoid is restricted to an integral sub-multiples of the data clock rate (i.e. the signal frequency falls exactly on one of the FFT frequency bins), then the frequency domain is as expected: a single frequency domain spectral line positioned at the sinusoid's frequency. However, this is a very narrow restriction. In nearly all cases, the frequency of the input signal will not line up with one of the frequency bins. Once this occurs the sinc pulse, which is still centered on the frequency of the input signal, no longer has its zero crossings line up with multiples of Fs/N. This is depicted in **Figure 5-14**.

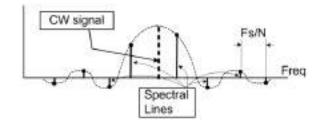


Figure 5-14. Spectrum of CW Signal *with* Spectral Leakage.

Now the zero crossings of the sinc function do not fall on FFT frequency bins. The result is not the single spectral line at the CW signal frequency, but rather a series of spectral lines that fall at the bin frequencies as shown. This phenomenon is known as spectral leakage.

The real damage of *off bin* sampling or spectral leakage is shown in **Figure 5-15**. A CW signal at approximately 50 MHz is shown. One version is centered on a frequency bin (dark line), the other is shifted in frequency so that it is no longer falling on a frequency bin (light line).



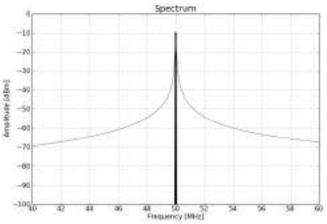


Figure 5-15. Signal On and Off Frequency Bin

The spectral leakage associated with this rectangular window causes a severe restriction on the dynamic range. **Figure 5-16** shows how a nearby signal that is 40 dB down from the central signal is nearly masked by the spectral leakage of the central signal.

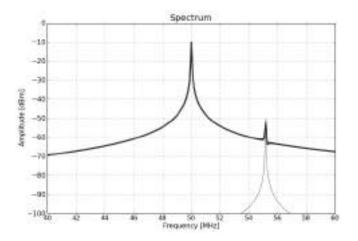


Figure 5-16. Spectral Leakage Masking a Nearby Signal.

The abruptness of the rectangular pulse edges in the time domain is what leads to the excessive leakage in the frequency domain. By multiplying a windowing function with the time domain signal before the FFT processing, the abruptness can at least be minimized [¹¹]. **Figure 5-17** shows the time domain signal with (dark trace) and without (light trace) an applied windowing function (Hamming window in this case).



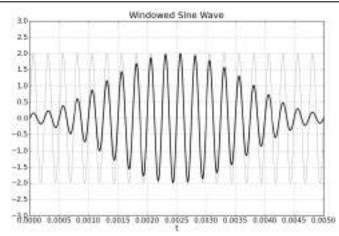


Figure 5-17. Signal Processed with Windowing Function in the Time Domain.

In the frequency domain, the effects of applying various windowing functions are shown in **Figure 5-18**. In this example, a signal has a frequency that does not line up with a frequency bin. The results with four different windowing functions are shown. The difference in the amount of spectral leakage varies greatly between the windowing functions.

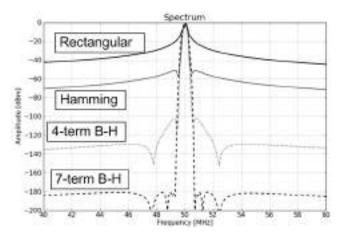


Figure 5-18. Spectral Leakage with Rectangular, Hamming, 4-Term Blackmann-Harris, and 7-Term Blackmann-Harris Windowing Functions.

Resolution Bandwidth Using Windowing Functions

Recall from the discussion of the analog IF signal processing that the RBW filter is used for frequency domain selectivity and SNR enhancement. The digital IF implementation used by National Instruments uses the windowing function to realize the RBW filter. The RBW bandwidth value is proportional to the number of time domain sample points. A higher number of sample points results in a narrower RBW bandwidth.

Figure 5-19 shows one example of the RBW realization using the Hamming windowing function.



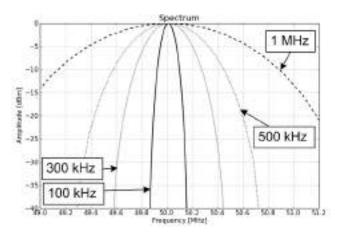


Figure 5-19. Hamming Windowing Function with Various Effective Resolution Bandwidths.

The number of time domain samples is adjusted to synthesize the different 3dB bandwidth settings; the bandwidth is inversely proportional to the number of samples. For example, the 500 kHz bandwidth filter requires twice the number of sample points as the 1 MHz bandwidth filter. Since the ADC has a constant sample rate, the data acquisition time required for the 500 kHz bandwidth is twice as long as that for the 1 MHz bandwidth.

Just as displayed noise floor is a function of RBW bandwidth for the analog only IF, so too is this behavior for digitally implemented RBW filters. Again signal selectivity and lower noise floor are reasons for choosing a narrower RBW filter.

Windowing Function Figures of Merit

There are dozens of windowing functions from which to choose. Some of the criteria for choosing the best windowing type for a particular measurement depend on: bin width, equivalent noise bandwidth scalloping loss, and side lobe attenuation. First, these terms are defined and then a summary table of the more popular windowing functions is given.

Bin Width

Bin width is the frequency resolution of the spectrum after FFT processing. If the number of time domain samples is held constant, then the bin width remains unchanged. **Figure 5-20** shows the spectrum of a CW signal with various windowing functions applied. For each window, the number of time domain samples is held constant and the control setting for the filter is specified using a constant bin width value.



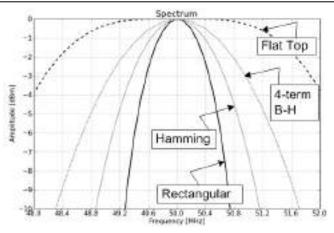


Figure 5-20. Windowing Functions Specified by Bin Width

Note that the 3 dB bandwidth associated with each windowing function varies. Between the widest and narrowest bandwidths of the windowing function in **Figure 5-20**, the ratio is 4.2 to 1. Holding the number of time domain samples constant for each windowing function, which implies a constant frequency domain bin width, does not result in constant filter bandwidth.

In **Figure 5-21**, the number of samples for each window is adjusted such that they all have the same 3 dB bandwidth.

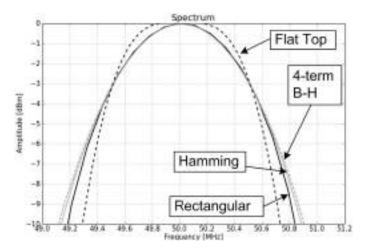


Figure 5-21. Windowing Functions Specified by 3 dB Bandwidth

Now the bandwidths are consistent, but the number of sample points for each windowing function is not constant. The 4.2:1 ratio of widest BW to narrowest BW in **Figure 5-20** becomes the ratio of sample points required to satisfy the consistent 3 dB BW requirement of **Figure 5-21**. Data acquisition time is directly proportional to the number of required sample points. The ADC is still sampling at the same rate for each windowing function, so in order to affect the number of sample points the acquisition time must be adjusted. So there is a tradeoff between selectivity of the windowing function and measurement time. Windows with lower selectivity require more sample points and hence longer acquisition time. For instance the Flat Top filter requires 4.2 times longer acquisition time over the rectangular filter to realize the same 3 dB bandwidth.



Equivalent Noise Bandwidth

Equivalent noise bandwidth (ENBW) is also referred to as *effective noise bandwidth*. Integrating the area under the curve of the magnitude squared of the windowing function in the frequency domain gives the *power gain*[¹¹]. Multiplying the power gain by the noise power yields the filter's *noise power*. A perfect brick-wall filter (infinitely sloped skirts) has a noise power equal to its 3 dB bandwidth. However, the actual filter will have a slightly higher noise power value as there is more area under the curve than the brick-wall filter.

ENBW is defined as the area under the curve of the filter's frequency response divided by the frequency range used to compute the power gain measurement as shown in **Equation 5-1**.

$$ENBW = \frac{\int_{f_1}^{f_2} |X(f)|^2 df}{f_2 - f_1}$$

Equation 5-1

Figure 5-22 shows that if a perfect brick-wall filter is constructed such that is has the same noise power as the actual filter, the ENBW is the bandwidth of that brick-wall filter. ENBW is always higher than the 3 dB BW of the filter response.

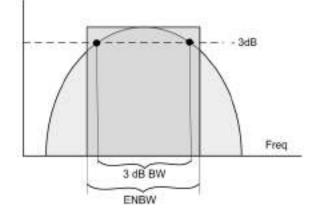


Figure 5-22. Equivalent Noise Bandwidth

What ENBW means to the user is that the noise floor of the signal analyzer is slightly affected by the choice of windowing function. The average noise specifications for the signal analyzer are usually given for a specific windowing function. If a different windowing function is used, then the noise floor can have an offset.

The windowing function bandwidth can be chosen in terms of ENBW instead of 3 dB BW. The number of samples is adjusted to ensure that the noise power of each windowing function is constant. **Figure 5-23** shows the filter response for the Hamming window when specified for 3 dB BW and ENBW.



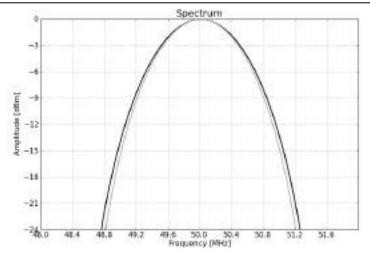


Figure 5-23. Hamming Window. Dark Trace = 3 dB BW Setting, Light Trace = ENBW Setting.

The ENBW filter is narrower than the 3 dB BW filter to compensate for the noise power gain.

Scalloping Loss

Signal frequencies off bin lead to amplitude errors as well as the generation of spectral leakage. The amplitude error is termed *scalloping loss* [⁷] and is depicted in **Figure 5-24**.

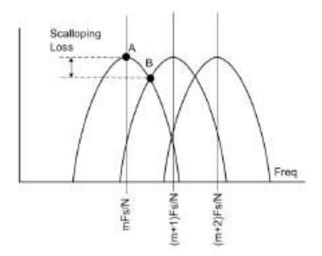


Figure 5-24. Scalloping Loss

If the signal frequency is positioned at point 'A' in **Figure 5-24**, the signal is on bin and there is no amplitude error. As the signal frequency drifts away from being on bin the amplitude will reach a minimum when the frequency is half way between bins (point 'B'). The amplitude difference between maximum and minimum is the scalloping loss. The amount of scalloping loss varies amongst the various windowing functions and must be considered when selecting windowing type.



Sidelobe Attenuation

Figure 5-18 shows a dramatic difference in the stopband attenuation of a few different widowing functions. By zero padding the FFT (adding trailing zero amplitude points to the time domain sample waveform), more stopband detail of the windowing function is exposed. This is shown in **Figure 5-25** for the Hamming windowing function.

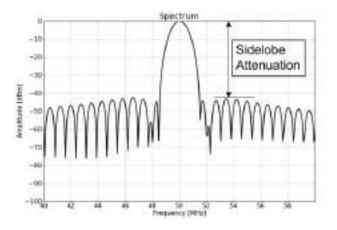


Figure 5-25. Hamming Widowing Function Sidelobe Attenuation

The amplitude difference from peak to one of the local maxima in the stopband is the sidelobe attenuation value. Again, the choice of windowing function dictates the amount of sidelobe attenuation.

Windowing Function Comparison

Table 5-1 shows some figures of merit for a few windowing functions.

 $3 \, dB \, BW$ is defined as the number of bins required to realize a RBW filter's 3 dB bandwidth. The number of bins is scaled to a 1 Hz BW. Use these as a relative ranking on the number of sample points required. For example, the Hanning window uses 1.451/0.886 = 1.63 times more sample points as compared to the rectangular window.

ENBW is also defined in terms of bins for a 1 Hz bandwidth.

Noise Power Gain is the dB difference in the noise power when specifying the filter in terms of its 3 dB bandwidth versus it ENBW, that is this is the noise power error when using the 3 dB BW specification.

Sidelobe Attenuation is the attenuation beginning 10 bins away from the center. Sometimes the closest sidelobe is used for this benchmark; however, this may be overly pessimistic for most signal analysis use cases.



| Windowing Function | 3 dB BW (bins) | ENBW (bins) | Noise Power Gain (dB) | Scalloping Loss (dB) | Sidelobe Attenuation (dB) |
|-------------------------|-------------------|-------------|--------------------------|-------------------------|---------------------------------|
| Rectangular | 0.886 | 1.0 | 0.519 | -3.905 | 30.3 |
| Hanning | 1.451 | 1.511 | 0.174 | -1.396 | 71.1 |
| Hamming | 1.309 | 1.370 | 0.195 | -1.727 | 47.3 |
| Blackmann | 1.656 | 1.739 | 0.214 | -1.077 | 78.9 |
| 4-term Blackmann-Harris | 1.913 | 2.019 | 0.234 | -0.809 | 99.2 |
| 7-term Blackmann-Harris | 2.501 | 2.652 | 0.253 | -0.475 | 173.7 |
| Flat Top | 3.751 | 3.797 | 0.052 | -0.011 | 91.9 |

Table 5-1 Windowing Function Comparison

| Lowest acquisition time | Rectangular | | |
|--|-----------------------------|--|--|
| Highest acquisition time | Flat Top | | |
| Least Noise Power Error | Flat Top | | |
| Lowest Scalloping Loss | Flat Top | | |
| Highest stopband attenuation | 7-term Blackmann- Harris | | |
| Bestoverallcompromise of speed,scallopingloss,stopbandattenuationand noise power error | 4-term Blackmann- Harris | | |

Trace Averaging

The next block in the signal processing chain of **Figure 5-10** is *trace averaging*. In the previous discussion, the windowing in the signal processing chain emulates the analog IF's RBW filters. Trace averaging emulates the VBW filter functionality as well as some of the detector functions.

Vector Averaging

This is also called coherent averaging [⁷]. In vector averaging, the time domain data from a specified number of data acquisitions is averaged. The averaged date is then processed with the FFT. In this manner the noise floor drops by 20log(N), where N is the number a trace averages specified. The intended application is for the measurement of time triggered events. If the signal is triggered at the same amplitude and phase from trace to trace, the averaged signal amplitude remains constant, but the noise floor decreases. This improves the signal to noise ratio of the measurement.

Another implementation is to operate on the spectral components by first applying the FFT operation. The real part and the imaginary part of each spectral component are then averaged separately.

RMS Averaging

RMS averaging is also known as *incoherent averaging* as it does not use an input signal to trigger the data acquisitions. The data acquisitions are free running which allows for a more general use case of measurement signals. RMS averaging operates on the power spectrum of the data and consequently the phase information is lost in this averaging process. As indicated in **Figure 5-10**, the time domain data is first processed with the FFT. For each spectral component, X_i , the magnitude is computed using:



 $|X_i| = \sqrt{(X_i)(X_i^*)}$ where X_i^* is the conjugate of X_i

Equation 5-2

RMS averaging closely resembles the analog IF's VBW filter functionality. The variance of the noise floor is reduced, but its average value is not affected.

Peak-Hold Averaging

Peak-hold averaging is not actually an averaging function. The user specifies a number of trace averages – these are independent data acquisitions. The peak-hold detector retains the maximum value at each frequency bin of all the data acquisitions. This mode is useful detecting transient events.

Averaging Mode Comparisons

Figure 5-26 show the comparison of the RMS, Vector, Peak-hold averaging functions. For each one, 100 trace averages are specified.

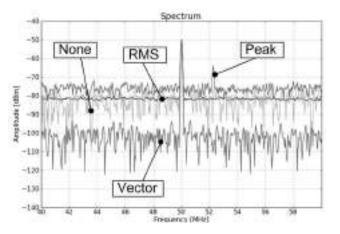


Figure 5-26. Averaging and Detection Modes

Video Bandwidth Filter Emulation

From the analog IF section of this document, the VBW filter uses a single pole lowpass structure. This filter operates on the video signal to reduce the variance of the random noise riding on the signal. The FFT-based signal analyzer does not implement a lowpass function, but rather maps the number of trace averages, using the RMS averaging mode to emulate the analog VBW's noise variance reduction capability [¹²].

Some measurement standards specify that the measurement be made with a particular VBW setting. Described here is the algorithm that maps the number of trace averages in a FFT-bases analyzer to the traditional analog IF RBW and VBW settings.

When the ratio of RBW to VBW is high, the approximation for the number of trace averages is proportional to the ratio of the equivalent noise bandwidths (ENBW) of the RBW to the VBW in the analog IF based signal analyzer. The analysis for this restriction is based on the demonstration that when detecting power using a nonlinear transformation, such as square-law detection, the spectral



transformation from IF to baseband results in the narrowband power spectrum at IF convolved with itself, except at DC [¹²].

$$S'(f) = c \int_{-\infty}^{\infty} S(g) S(f-g) dg$$

where S(f) is the normalized power spectrum of RBW filter and *c* is a constant.

Equation 5-3

The ENBW of the RBW filter is then shown in Equation 5-4.

$$ENBW_R = \frac{\int_0^\infty S'(f)df}{S'(0)}$$

Equation 5-4

For a synchronously-tuned RBW implementation, the magnitude response of the N-order filter is shown in **Equation 5-5**.

$$|H(f)| = \left[\frac{1}{\left[(2(f - f_c)/a]^2 + 1\right]}\right]^{(N/2)}$$

Equation 5-5

$$a = B/\sqrt{2^{1/N} - 1}$$

Equation 5-6

where, B = 3 dB BW of the filter f_c = the RBW filter center frequency

When white noise is passed through the filter, the normalized power spectrum is shown in Equation 5-7.

$$S(f) = \left[\frac{1}{\left[(2(f-f_c)/a]^2 + 1\right]}\right]^{(N^2/4)}$$

Equation 5-7

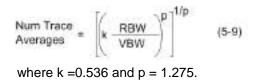
For a 4th order RBW filter, ENBW_R = 0.842 x B = 0.842 x RBW setting, the ENBW of a single pole lowpass filter is (π /2)VBW or 1.571VBW. So for VBW << RBW, the number of trace averages, N, is shown in **Equation 5-7**.

 $N \sim ENBW_R / ENBW_V = 0.536 x (RBW / VBW)$

Equation 5-8



However if the VBW bandwidth is not much less than RBW bandwidth, then the VBW has less affect on trace averaging in which case the number of trace averages should decline to 1. **Equation 5-9** approximates the transition from VBW << RBW to VBW > RBW:



Equation 5-9

Equation 5-8 is plotted in Figure 5-27.

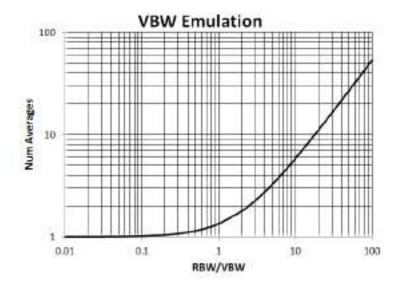


Figure 5-27. Number Trace Averages vs. RBW/VBW Ratio

Unit Conversion

The final block in the signal processing chain of **Figure 5-11** is the *unit conversion* stage. This allows the amplitude of the spectrum data to be recorded in several different formats including: Volts, $(Volts)^2$, Watts (computed using user controlled impedance setting), dBm **Equation 4-1**, and dBuV or dBmV **Equation 4-2**.

Log Mode Trace Averaging

In section 4 of this document, the thermal noise (sometimes called kTB noise) is present at the input of the signal analyzer. The signal analyzer then adds internally generated noise. For both cases, the noise over the measurement range is considered to be additive white Gaussian noise (AWGN). White refers to the noise having a flat frequency response, and Gaussian refers to the probability distribution function (pdf) of the noise voltage having normal distribution as shown in **Figure 5-28**.



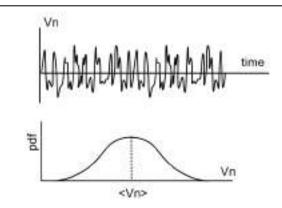


Figure 5-28. Thermal Noise Probability Distribution Function

With RMS averaging, the variance, or spread in the pdf curve, reduces without the average value of the noise voltage changing. This is depicted in **Figure 5-29**.

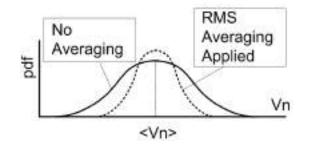


Figure 5-29. Variance Reduction with RMS Averaging

If one uses the ENBW setting for specifying the RBW bandwidth, then the displayed power after RMS averaging is the true representation of signals with any kind of statistics. If the 3 dB BW setting is used for specifying the RBW bandwidth, then the noise power gain corresponding to the selected windowing function must be added. For instance, if a 7-term Blackmann-Harris windowing function is used, then add 0.253 dB to the measured RMS averaged noise power.

The traditional spectrum analyzer with the analog IF **Figure 5-1** does not use RMS averaging. The signal is first sent through the log amplifier and then the envelope detector before the VBW filter effectively averages the signal. The act of envelope detection on a log scaled noise voltage alters the statistics of the noise probability distribution function as depicted in **Figure 5-30**.

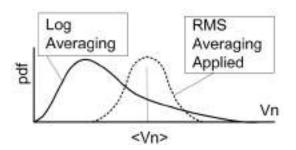


Figure 5-30. PDF with RMS and Log Averaging



The log averaging lowers the mean of the noise power by 2.51 dB [¹³]. The amplitude of a CW signal is not altered by either RMS averaging or log averaging, so for the measurement of CW signals, the SNR is a true 2.51 dB improvement using log averaging over RMS averaging.

In traditional spectrum analyzers with synchronously tuned RBW filters, the noise power gain is approximately 0.5 dB [¹⁴]. Thus the displayed average noise level, which effectively uses the 3 dB BW setting for the RBW shows an approximate -2.0 dB noise power from the true RMS noise level. **Figure 5-31** shows the SNR of a CW signal using RMS and Log averaging.

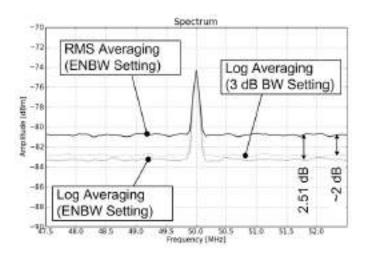


Figure 5-31. SNR with RMS Averaging and Log Averaging Modes.

The -2.51 dB noise power bias is a predictable offset for AWGN, so for the measurement of CW signals, log averaging is very appropriate. However, for modulated signals where the statistics of the signal is not predictable, then log averaging does not accurately report the signal's power level. In the case of modulated signals, the VBW filter setting of the traditional spectrum analyzer needs to be set much wider than the RBW setting. This, however, offers no variance reduction in the signal. With the advent of digitizing the video signal, trace averaging using digital techniques made variance reduction possible without altering the mean value of the signals power. However, by using RMS averaging mode, the statistics of both the broadband thermal noise and the digitally modulated signal are unaltered, making RMS averaging the appropriate averaging mode for this class of signals.

Care should be used when comparing average noise level performance of signal analyzers. Even though modern signal analyzers with all digital IFs do not need to average on the envelope of the log data, some do this anyway using DSP means in order to add the apparent 2.51 dB improvement in noise level specifications. Unless the specification calls out the average noise level with RMS notation, one can assume that the noise is measured using log averaging. For a true comparison of the signal analyzer's noise figure or a true comparison of the SNR when measuring modulated signals, the noise specification should be converted to RMS.



6. Dynamic Range

Dynamic range is arguably one of the more import figures of merit for the signal analyzer. In section 2, the mechanisms that cause distortion in the signal analyzer were discussed. In section 4, the amplitude control of the signal analyzer was introduced. In this section, the discussion turns to how to optimize the signal analyzer to minimize both its distortion and noise. In the next section, the accuracy of the measurement will be presented.

We first need to define what is meant by dynamic range and what some of the more common measurements where dynamic range is important are. The dynamic range versus mixer level chart allows the user to determine how to optimize the signal analyzer settings for best dynamic range performance. Construction of the dynamic range vs. mixer level chart is first described, and then the use this chart for various measurements is discussed.

In section 1, the idea of signal analysis versus vector signal analysis was introduced. Recall that signal analysis uses narrowband analog IF filtering and vector signal analysis requires relatively wide analog IF bandwidth (bandwidth must be greater than the modulation bandwidth of the signal being measured). The signal analyzer optimization is different depending on whether the measurement is narrowband versus wideband. Optimizing the signal analyzer for each type of measurement is discussed next.

Dynamic Range Definitions

One of the most overloaded terms in RF is the term *dynamic range* itself. Without qualification dynamic range can be used to describe spurious-free dynamic range, harmonic distortion, signal-to-noise ratio, intermodulation distortion, gain compression, and more. But, in general, dynamic range usually refers to the ability to simultaneously display large and small amplitude signals in a single measurement.

How the signal analyzer corrupts the dynamic range of a general distortion measurement is shown in **Figure 6-1**.

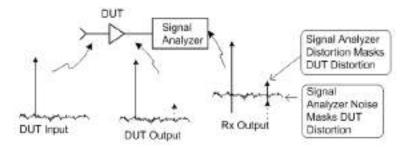


Figure 6-1. Signal Analyzer's Contribution to Distortion and Noise.

Consider a pure, distortion-free signal input to a DUT. The DUT will produce some distortion -- harmonic distortion in this case. Next the signal analyzer attempts to measure the DUT's distortion. The signal analyzer will create distortion of its own and these distortion products fall at the same frequencies as those of the DUT. Potentially, the amplitude of the signal analyzer's distortion could be larger than the DUT's. Additionally, noise generated by the signal analyzer could mask the DUT's distortion. With either signal analyzer generated distortion and/or noise, the effect is the same: the signal analyzer does not have enough dynamic range to measure the DUT's distortion.



For the context of this discussion, only the dynamic range of the signal analyzer is considered. The signal analyzer's dynamic range is what is trying to be maximized. The general definition of dynamic range for the signal analyzer is:

Signal analyzer dynamic range is the ratio of the input signal amplitude to the signal analyzer generated distortion or signal analyzer generated noise.

This definition does not state anything about the amplitude display range, or about the maximum signal to the noise floor. This definition also does not state that the measurement of the signal and distortion be made with the same signal analyzer settings.

Maximizing the signal analyzer's dynamic range involves changing gain settings to trade off signal analyzer generated distortion and noise. When the distortion amplitude equals the noise floor, *maximum dynamic range* is achieved. **Figure 6-2** shows the conditions for maximum dynamic range.

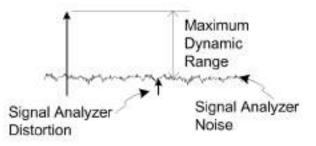


Figure 6-2. Maximum Dynamic Range

Now we turn to a listing of several distortion measurement examples.

Gain Compression

A linear device is normally a passive device which when driven with an input signal, its output signal has an amplitude that is a scaled version of the input signal's amplitude. This scaling is constant with applied power. A nonlinear device will have a scaling factor that can have different values depending on the applied signal level.

Many nonlinear devices, such as amplifiers and mixers, behave nearly linear when lower amplitude signals are applied. That is, these devices have near constant gain versus input power. **Figure 6-3** shows the graph output power vs. input power, both on a log scale, for an amplifier.



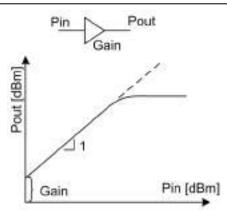


Figure 6-3. Pout versus Pin for Two Port Device

For lower input power levels, the slope of the curve is one, meaning that for every one dB increase in the power level of the signal at the amplifier's input port, the output power level increases by approximately one dB. The offset in this curve is the small-signal gain of the device.

Plotting the gain versus power at either the input or output, at some power level, the gain will begin to fall, or compress. The common figure of merit is the one dB gain compression level, which is the power level at which the small-signal gain drops by one dB. This level is termed P1dBm. The plot of gain versus power level is shown in **Figure 6-4**.

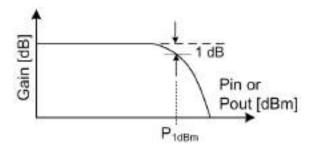


Figure 6-4. Gain Compression in a Two Port Device

Amplifiers normally specify P1dBm at the output and mixers normally specify P1dBm at the input.

The signal analyzer, largely comprised of amplifiers and mixers, also has a gain compression specification. Most often the specification conditions are with 0 dB RF input attenuation, which is a way of describing the mixer level. However, for default gain settings, one cannot directly measure the gain compression of the signal analyzer. Recall from section 4 the signal analyzer's internal gain is such that for an input signal whose power level is positioned at the reference level setting, the signal level at the final IF is close to the ADC's full scale input level. Measuring a signal whose amplitude is above the reference level runs the risk of over-driving the ADC. The maximum mixer level constrains such that the signal level at the first mixer is at least 10 dB to 15 dB below the signal analyzer's gain compression level. Measuring a signal with amplitudes nearing the signal analyzer's gain compression level results in an IF overload error, indicating that the ADC full scale input level is being exceeded.

For signal analysis, the danger exists when large amplitude signals fall outside the frequency span as depicted in **Figure 6-7**.



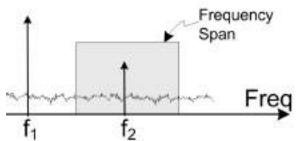


Figure 6-7. Gain Compression in a Signal Analyzer.

A large amplitude signal at frequency f_1 outside of the frequency span can drive the signal analyzer's frontend into gain compression without over-driving the ADC. Narrowband filters in the signal analyzer's IF chain can attenuate out of band signals. When driven into gain compression by out of band signals, the lower level signals that fall inside the viewing range, such as the signal at frequency f_2 in **Figure 6-7**, have an unpredictable amplitude error.

Another overlooked issue associated with signal analyzer gain compression is the measurement of digitally modulated signals. To a rough approximation, the displayed power spectral density of the modulated signal, whose modulation bandwidth equals BW_m , is $10Log(BW_m)$ below a CW signal whose power is the same at the average power of the digitally modulated signal. For instance, for a digitally modulated signal whose modulation bandwidth is 1 MHz, the signal level appears 60 dB below a CW tone whose power equals the average power of the modulated signal. However, this is only partly true since the displayed amplitude is also a function of RBW setting. **Figure 6-8** shows the relationship between the displayed digitally modulated signal and a CW signal of equal power levels. **Figure 6-8** is with a RBW setting of 1 Hz.

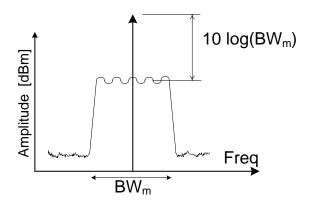


Figure 6-8. Power of a Modulated Signal

The signal analyzer gain compresses according to the total power at the first mixer. By measuring the power in band, a feature within the NI Spectral Measurement Toolkit, the total power of the modulated signal can be assessed. It is quite easy to have the modulated signal appear below the signal analyzer's reference level without registering an IF overload error, yet its total power exceeds the gain compression specification.

Many signal analyzers specify the gain compression using the *two-tone desensitization technique*. The two-tone desensitization measurement technique is a requirement for the signal analyzer as a signal tone



technique as depicted for the amplifier example in **Figure 6-3** would cause an IF overload condition. **Figure 6-9** shows the spectrum for the two-tone desensitization technique.

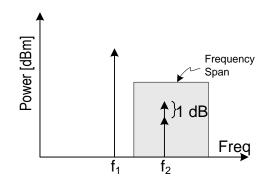


Figure 6-9. Two-tone Desensitization Measurement for Gain Compression

First a low level signal at frequency f_2 is placed within the frequency span of the signal analyzer. An outof-band signal at frequency f_1 , where the separation frequency is wider than the analog IF bandwidth of the signal analyzer, is configured such that its amplitude is well below the gain compression level. As the out-of-band signal amplitude is increased, the amplitude of the in-band signal at f_2 will fall by 1 dB at a certain power level. The power level of the out-of-band signal is considered the gain compression level of the signal analyzer. Because of attenuation due to IF filtering within the signal analyzer, the large amplitude tone will not overload the ADC.

Harmonic Distortion

Consider the signal analyzer when operating far below its gain compression level. Under this condition, the signal analyzer can be treated as a *weakly nonlinear device*. The output voltage as a function of input voltage for a weakly nonlinear device can be described with a power series expansion as shown in **Equation 6-1**.

$$V_{out}(t) = a_0 + a_1 V_{in} + a_2 V_{in}^2 + a_3 V_{in}^3 + \dots$$

Equation 6-1

For the signal analyzer, V_{out} is used to form the power recorded for display and V_{in} is the voltage corresponding to the signal at the RF input port.

Consider the a sinusoidal input signal given by $V_{in} = Acos(\omega t)$. Inserting V_{in} into **Equation 6-1** and expanding terms results in **Equation 6-2**.

$$V_{out}(t) = a_1 A \cos(\omega t) + a_2 A^2 \cos(2\omega t) + a_3 A^3 \cos(3\omega t) + \dots$$

Equation 6-2

Plotting **Equation 6-2** on a spectrum with $\omega = 2\pi f_{o}$, the result is the fundamental tone at frequency f_{o} , second harmonic at frequency $2f_{o}$, third harmonic at frequency $3f_{o}$, and so on.



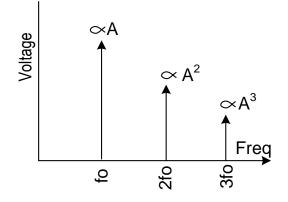


Figure 6-10. Spectrum Showing Harmonic Distortion.

The order of the distortion product stems from its amplitude dependence. The second harmonic has second order distortion because its amplitude is proportional to the square of the fundamental signal amplitude. Likewise, the third harmonic has third order distortion due to its amplitude being proportional to the cube of the fundamental signal's amplitude.

When plotted on the log scale, for every 1 dB change in the fundamental signal amplitude, the second order distortion product amplitude drops by 2 dB, the third order distortion product amplitude drops by 3 dB, and so on as shown in **Figure 6-11**.

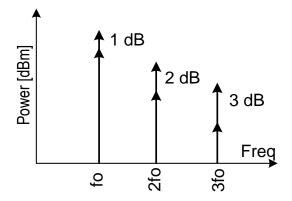


Figure 6-11. Harmonic Distortion on a Log Scale.

Second Harmonic Intercept

Second harmonic intercept (SHI) is another figure of merit for the signal analyzer. SHI is a single value with power units in dBm. With this single value, the user can determine the second harmonic distortion product amplitude for any given fundamental power level. **Figure 6-12** shows graphically the concept of SHI.



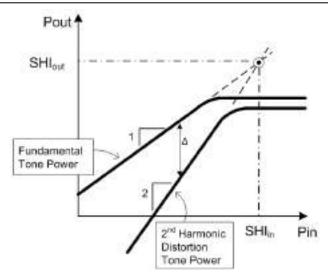


Figure 6-12. Second Harmonic Intercept

Both the fundamental tone power and 2nd harmonic distortion product power are plotted on the Pout vs. Pin graph. Below gain compression, the fundamental tone curve has a slope of one and the 2nd harmonic tone curve has a slope of two. If these two curves are extrapolated, they cross. This crossover point corresponds to the SHI value. Input SHI is the power at the input at the device where the lines cross and output SHI is the device's output power level where these lines cross. The difference between input and output SHI is the gain of the device. Of course, in reality these lines do not cross as the device goes into gain compression before crossover can occur.

Even though fictitious, SHI is still a very useful parameter for both comparing one signal analyzer's performance against another and for predicting distortion when operating in the linear region. By using simple geometry on the graph shown in **Figure 6-12**, an equation for SHI can be deduced as shown in the following:

 $SHI = Psig + \Delta$ (*dBm*) where *Psig* is the power of the fundamental tone and $\Delta = Psig - Pharm$ *Pharm* is the power of the 2nd harmonic signal.

Equation 6-3

For output SHI, *Psig* is measured at the output of the device and for input SHI, *Psig* is measured at the input of the device. SHIout (dBm) =is equal to SHIin (dBm) plus Gain (dB). The signal analyzer has a gain of zero dB, and so SHIout is equal to SHIin. Furthermore, SHI for the signal analyzer is normally specified with 0 dB RF input attenuation.

For the signal analyzer, SHI is a given value. One can use this value to determine the harmonic level. For example, suppose the signal analyzer SHI specification is +40 dBm and the power at the input port with 0 dB RF input attenuation is -10 dBm. Plugging these values into **Equation 6-3** results in the following:

SHI = $Psig + \Delta$ $\Delta = SHI - Psig$ $\Delta = 40 - (-10) = 50 dB$ $\Delta = Psig - Pharm$ $Pharm = Psig - \Delta$



Pharm = -10 - 50 = -60 dBm

This states that the harmonic level is -50 dBc, or -60 dBm.

Intermodulation Distortion

Intermodulation distortion (IMD) occurs when two or more signals are present at the input of a nonlinear device. Frequency mixing between these fundamental components will create distortion tones. The most common (IMD) is with two input tones of equal amplitude. Back to **Equation 6-1**, which shows the V_{out} vs. V_{in} relationship of a weakly nonlinear device, suppose V_{in} is the summation of two sinusoids shown in the following:

 $V_{in} = Acos(\omega_1 t) + Bcos(\omega_2 t)$

the resulting output voltage is the following:

 $V_{out} = a_1 \{Acos(\omega_1 t) + Bcos(\omega_2 t)\} + a_2 \{A^2 Bcos[(2\omega_1 - \omega_2)t]\} + a_2 \{AB^2 cos[(2\omega_2 - \omega_1)t]\} + \text{many other terms}$

Equation 6-4

With $\omega_1 = 2\pi f_1$ and $\omega_2 = 2\pi f_2$, the two-tone IMD products closest in frequency to the fundamental tones are shown in **Figure 6-13**.

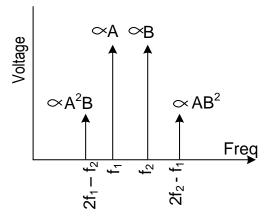


Figure 6-13. Third Order Intermodulation Distortion

These IMD products are particularly troublesome due to their proximity to the fundamental tones. If f_1 and f_2 are separated in frequency by Δf , the distortion products are also separated by Δf from their closest fundamental tones. Filtering these distortion products is nearly impossible when the fundaments tones are at high RF and the separation frequency is small.

If the fundamental tone powers both have amplitude of 'A', then it is easy to see that the distortion product amplitudes are proportional to A^3 making these third order. To reiterate, the nature of third order distortion is, as shown in **Figure 6-14** when the fundamental tone powers both change by 1 dB, that the IMD power levels change by 3 dB.



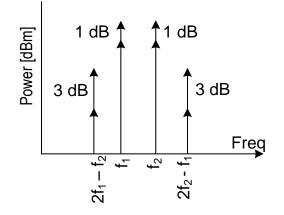


Figure 6-14. Third Order Intermodulation Distortion on a Log Scale

As with second harmonic distortion, the two-tone IMD also uses the concept of the intercept point as a figure of merit. The intercept point for two-tone intermodulation distortion goes by either third order intercept (TOI) or IP3. **Figure 6-15** shows the Pout vs. Pin relationship for third order IMD.

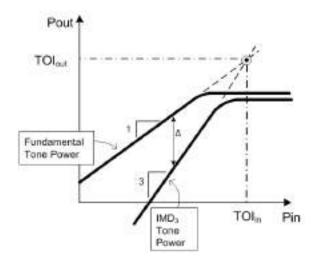


Figure 6-15. Third Order Intercept Point

For third order distortion, the IMD curve has a slope of three when operating in the linear region, which is below the gain compression level. It is assumed that the fundamental tone powers are both being adjusted to create the curve with the slope of one. Extrapolating these two curves beyond gain compression shows a similar crossover as with the second harmonic distortion graph of **Figure 6-11**. Input TOI or IP3 is the input power at the crossover and output TOI or IP3 is the output power at the crossover.

Amplifiers tend to specify output TOI or IP3 and mixers normally specify input TOI or IP3. As with SHI, TOIout (dBm) is equal to TOIin (dBm) plus Gain (dB). Again, signal analyzers have a gain of 0 dB, which makes TOIout equal to TOIin for the signal analyzer.

With some simple geometry on the fundamental and IMD curves in Figure 6-15, TOI equates to Equation 6-5.



 $TOI = Psig + \frac{\Delta}{2} (dBm)$ where *Psig* is the power of each fundamental tone and $\Delta = Psig - P_{IDM}$ P_{IMD} is the power of each IMD product.

Equation 6-5

For the signal analyzer, TOI is usually specified with the RF input attenuator set to 0 dB. To calculate the IMD power level, **Equation 6-5** can be rearranged into **Equation 6-6**.

 $P_{IMD} = 3P_{sig} - 2TOI$

Equation 6-6

As an example, suppose the TOI specification for the signal analyzer is +20 dBm and the power of each fundamental tone is -10 dBm. The IMD power level is then as follows:

$$P_{IMD} = 3x(-10 \ dBm) - 2x(20 \ dBm)$$

= -50 dBm

Optimizing the Signal Analyzer's Dynamic Range Performance

In section 4, we presented a qualitative discussion on setting the signal analyzer's RF input attenuation and reference level settings in regards to the signal analyzer's noise and distortion. Now a more detailed analysis on how to optimize the signal analyzer's setting to maximize its dynamic range performance is presented. We will first analyze the optimization process for the measurement of CW signals for both narrowband and wideband IF modes. This will then help lead into the more challenging optimization for the measurement of digitally modulated signals.

Dynamic Range versus Mixer Level Chart

The *dynamic range vs. mixer level chart* is an often overlooked piece of information found in some signal analyzer data sheets. However, this tool's importance cannot be stressed enough for setting the signal analyzer's gain parameters for highest dynamic range. **Figure 6-16** shows an example dynamic range vs. mixer level chart taken from a signal analyzer's data sheet [¹⁵].



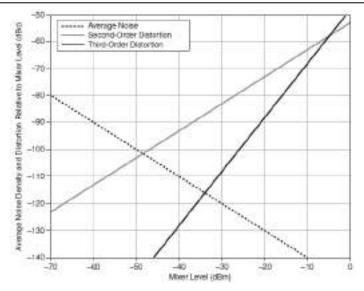


Figure 6-16. Example Dynamic Range Chart

Admittedly the axis labeling lacks clarity and most data sheets do not offer information on how to use this chart for dynamic range optimization. The next part of the discussion is centered on the construction of this chart. By showing how this chart is constructed, more insight in its use will be apparent. More importantly, some signal analyzer data sheets do not contain this chart. Being able to construct this chart with measured data will allow the user to know how to optimize the signal analyzer's dynamic range.

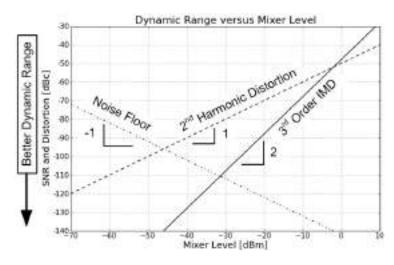


Figure 6-17 shows another view of the dynamic range chart with some clarity added.

Figure 6-17. Dynamic Range Chart Definition

The three lines drawn are the *noise floor*, 2nd *harmonic distortion*, and 3rd *order intermodulation distortion*. When considering the noise floor curve, the y-axis serves as the noise power relative to the signal amplitude with units of dB relative to the carrier (dBc). This is signal-to-noise ratio with the sign inverted. When considering the distortion curves, now the y-axis represents the distortion amplitude relative to the fundamental signal level (the units are dBc). For both noise and distortions, the y-axis represents dynamic range with better dynamic range performance towards the bottom of the chart.



The x-axis is the mixer level. Why not use the signal analyzer's RF input power level for the x-axis? This is certainly possible; however, a separate chart for each RF input attenuation value would need to be constructed. By normalizing x-axis power to the input of the first mixer, one chart only needs consideration. Since changing the test signal's power level or the RF input attenuation setting affects the mixer level, using mixer level as the x-axis adds further flexibility in that the dynamic range chart is valid for any combination of power settings.

Noise Floor Curve on the Dynamic Range Chart

Assume there is no excess noise from the signal being measured. With this stipulation, the noise power at the input of the signal analyzer is kTB or -174 dBm/Hz for a 50 Ohm system. For passive components whose characteristic impedance is the same as the system (50 Ohm in this case), the noise floor cannot drop below kTB. Considering that only passive devices are present before the first mixer in the signal analyzer with the preamplifier bypassed, the noise floor at this point is kTB. Viewing the signals in **Figure 6-18**, with constant noise floor, for every one dB drop in the signal level, the SNR drops one dB.

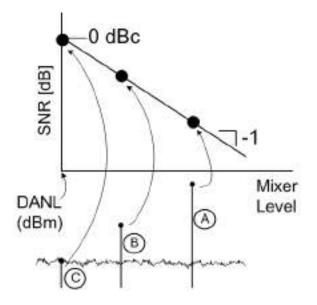


Figure 6-18. Mapping SNR to Dynamic Range Chart

Mapping the signals onto the dynamic range chart in **Figure 6-16**, signal A with its higher SNR relative to signal B falls lower on the y-axis. The dB per dB relationship with SNR and signal level gives the noise curve slope of -1. The signal analyzer will add noise to the measured results; however, the SNR to mixer level slope still stays the same value of -1.

This straight line curve is a y = mx + b equation to solve. The missing element is the offset, b. Consider when the signal level equals the noise floor, depicted by signal C in **Figure 6-18**. The SNR is by definition, 0 dBc. Solving for offset in the straight line equation shows b to equate to the value of the signal analyzer's DANL.

In Figure 6-18, the DANL in a 1 Hz RBW setting is -150 dBm. The equation of line 1 in Figure 6-18 is:

y = -x - 150.

Figure 6-19 also shows the noise floor curve as a function of RBW setting. For the measurement of CW signals, the signal level is independent of the RBW setting, but from **Equation 5-2**, the noise varies by



10log(RBW ratio). In **Figure 6-19**, curve B is with an RBW setting of 1 kHz, giving a 30 dB offset in SNR over curve A which uses a 1 Hz RBW setting.

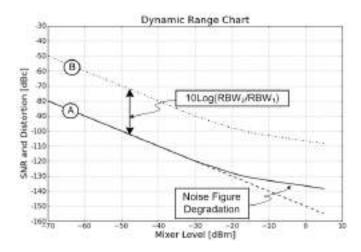


Figure 6-19. RBW Effect on Noise Curve

Figure 6-19 also shows some more realistic behavior of most signal analyzers in that as the mixer level increases, the system gain drops. This usually has an effect on noise figure degradation at these higher mixer levels.

Distortion Curves on the Dynamic Range Chart

In **Figure 6-11** we demonstrated that for a second order distortion product, every one dB change in the fundamental tone power, the distortion level changes two dB. The dynamic range chart plots relative distortion on the y-axis and **Figure 6-20** shows how to map absolute distortion level to relative distortion level.

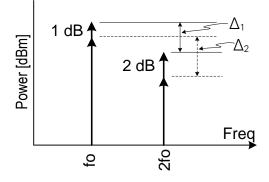


Figure 6-20. Relative 2nd Harmonic Distortion Level Change

With the fundamental a lower power setting, the relative 2^{nd} harmonic power is Δ_2 dBc. Increasing the fundamental power by one dB, the relative 2^{nd} harmonic power becomes Δ_1 dBc. The difference between Δ_2 and Δ_1 is +1 dB. The slope of the line in the dynamic range chart is the change in the relative harmonic level divided by the change in mixer level; hence, the slope is +1. In fact, the slope of any distortion line on the dynamic range chart is always one less than the order of the distortion.



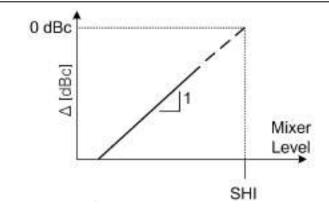


Figure 6-21. Mapping Second Harmonic Distortion to the Dynamic Range Chart

Again, there is a y = mx + b straight-line equation to solve where the offset, b, is unknown. From the definition of second harmonic intercept, when the mixer level is at the SHI value in dBm, the second harmonic is 0 dBc, yielding a value of –SHI for the offset, b. Even though this mixer level is beyond the gain compression level, it can still be used to draw the second harmonic distortion curve.

For third order IMD, a similar analysis reveals that for every one dB change in mixer level, the difference in relative distortion levels is two dB as shown in **Figure 6-22**.

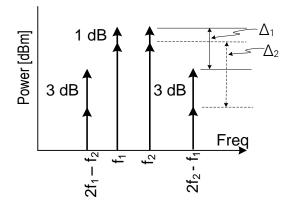


Figure 6-22. Relative 3rd Order IMD Level Change

The resulting slope of the IMD curve on the dynamic range is then +2. Like for the case of SHI, when the mixer level is at the TOI value, the IMD level is 0 dBc. Solving the y = mx + b strait-line equation shows that offset, b, equates to a value of -2xTOI.



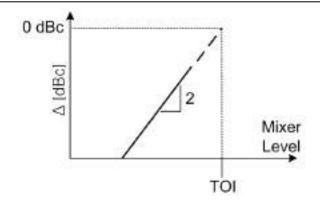


Figure 6-23. Mapping 3rd Order IMD to the Dynamic Range Chart

Figure 6-24 shows the second harmonic distortion and the third order IMD curves. In this example, SHI equals +30 dBm and TOI equals +20 dBm. The shaded region corresponds to power levels above the maximum mixer level where the IF overload error may occur. However, this graph shows how the two curves anchored at 0 dBc with fictitious mixer level values equal to either SHI or TOI

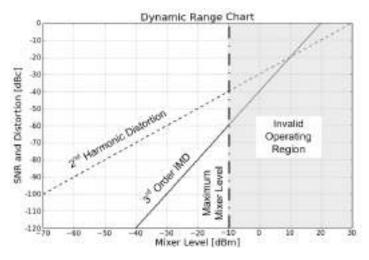


Figure 6-24. Distortion Curves on the Dynamic Range Chart

Phase Noise Curve on the Dynamic Range Chart

Phase noise when plotted on the dynamic range chart is the value at one particular offset frequency from the carrier. As shown in **Figure 6-25**, the single sideband phase noise value is a function of offset frequency,

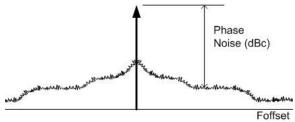


Figure 6-25. Phase Noise Pedestal



However, the relative phase noise power is constant as a function of carrier power level. Phase noise specification most often is normalized to a 1 Hz RBW value. Add 10Log(RBW) to the phase noise specified in dBc/Hz to arrive at the actual dBc value of phase noise.

On the dynamic range chart, phase noise appears as a horizontal line as shown in Figure 6-26.

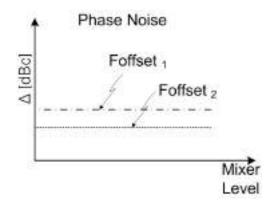


Figure 6-26. Phase Noise on the Dynamic Range Chart

Phase noise values at two different frequency offsets are shown.

Complete Dynamic Range Chart

Including the thermal noise, 2nd harmonic distortion, 3rd order intermodulation distortion and phase noise completes the dynamic range chart as shown in **Figure 6-27**.

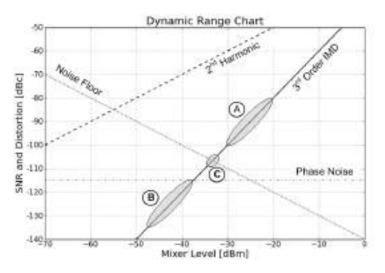


Figure 6-27. Complete Dynamic Range vs. Mixer Level Chart

The *noise floor* sets a hard line for this chart. Above the noise floor line, the signal can be viewed, but below the line, the signals are buried in noise. **Figure 6-28** demonstrates this concept. The letters correspond to the regions shown in **Figure 6-27**.



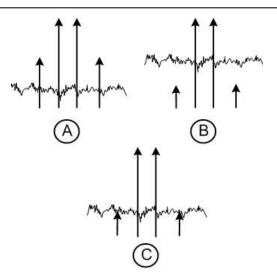


Figure 6-28. 3rd Order IMD at Various Mixer Levels.

When the mixer level is set for operating in region A, the distortion products are above the noise floor. The dynamic range of the signal analyzer is distortion limited. When the mixer level is set for operating in region B, the distortion products fall below the noise floor. In this case the dynamic range is noise limited. If the mixer level is set for region C, the distortion and noise levels are the same. The dynamic range is at its maximum when operating in region C.

When the mixer level is adjusted for maximum dynamic range, this value is termed the optimum mixer level. As shown in **Figure 6-29**, there is an optimum mixer level for 2nd harmonic distortion and one for 3rd order IMD.

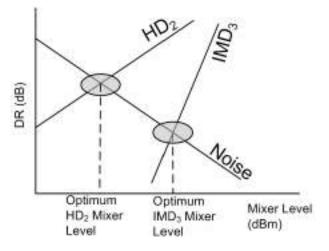


Figure 6-29. Optimum Mixer Levels

Some geometry on the curves shown in **Figure 6-29** yields the following equations for maximum dynamic range and the corresponding optimum mixer level.

3rd Order IMD:

Maximum dynamic range: $DR_{max} = (2/3)[TOI - DANL]$



Equation 6-7

Optimum mixer level: MLopt = (1/3)[2TOI + DANL]

Equation 6-8

2nd Order Harmonic Distortion:

Maximum dynamic range: $DR_{max} = (1/2)[SHI - DANL]$

Equation 6-9

Optimum mixer level: MLopt = (1/2)[SHI + DANL]

Equation 6-10

The following is a worked example for Figure 6-27: RBW = 10 Hz DANL = -150 dBm/Hz DANL = -140 dBm TOI = +20 dBm $DR_{max} = (2/3)[20 - (-140)]$ = 106.7 dB MLopt = (1/3)[2(20) + (-140)] = -33.3 dBm

A Few Observations Regarding the Dynamic Range Chart

One of the first mental barriers to overcome is the seemingly low mixer level for best dynamic range performance. Users who are familiar with digitizers know that for best dynamic range, the signal should be set as high as possible (see **Figure 4-6** for more information on this topic). But, the dynamic range chart as we have defined it so far, does not consider the ADC. It considers only the frontend.

From **Figure 6-27**, the RF input attenuator must be set quite high to achieve around -35 dBm for maximum 3rd order IMD dynamic range and -55 dBm for maximum 2nd harmonic distortion dynamic range. During these measurements, it is quite easy to assume that the excessive noise with such low mixer levels impedes good dynamic range performance. But, the chart clearly demonstrates that the signal analyzer's best dynamic range performance is at these seemingly low mixer levels.

Figure 6-30 shows that a 10 dB improvement in noise floor equates to a 5 dB improvement in 2nd harmonic dynamic range and 6.67 dB in 3rd order IMD dynamic range performance.



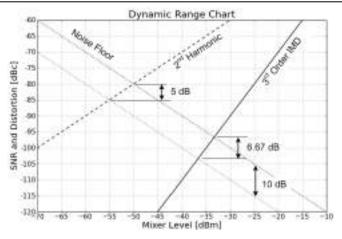


Figure 6-30. Noise Floor and the Dynamic Range Chart

Figure 6-31 shows two things: that higher TOI results in an IMD line that is displayed lower on the dynamic range chart and that a 10 dBm improvement in TOI lowers the IMD line 20 dB, but all this improvement still only yields a 6.67 dB improvement in dynamic range performance.

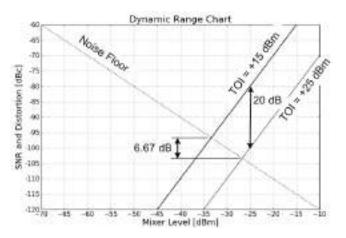


Figure 6-31. 3rd Order IMD and the Dynamic Range Chart

Preamplifier Dynamic Range

Does cascading a preamplifier in front of the signal analyzer help with dynamic range? From **Equation 4-6**, the preamplifier can certainly improve system noise figure. To answer the question of dynamic range, a numeric example is worked out. **Figure 6-32** shows the signal analyzer and preamplifier performance parameters for this example.

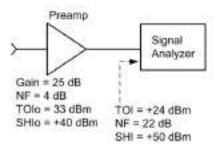




Figure 6-32. Preamp Cascaded with the Signal Analyzer

Equation 4-6 allows the cascaded preamplifier plus signal analyzer noise figure to be calculated. The following equations can be used to compute cascaded SHI and TOI:

$$SHI_{sys} = -20 \log \left\{ 10^{(-SHI_1^{'}/20)} + 10^{(-SHI_2^{'}/20)} \right\}$$

Equation 6-11

$$TOI_{sys} = -10 \log \left\{ 10^{(-TOI_{1}^{'}/10)} + 10^{(-TOI_{2}^{'}/10)} \right\}$$

Equation 6-12

The *SHI*['] and *TOI*['] are values referred to the system input by subtracting the system gain that precedes the device. For instance, *SHI*['] for the signal analyzer is SHI of the signal analyzer minus the preamplifier gain or +50 – 25, or +25 dBm.

The results of the signal analyzer dynamic range and the cascaded preamplifier plus the signal analyzer dynamic range are summarized in **Table 6-1**.

| | Signal Analyzer | Preamp + Signal Analyzer | Change |
|---|--------------------|--------------------------------|----------|
| TOI | +24 dBm | -1.5 dBm | -25.5 dB |
| SHI | +50 dBm | +12.6 dBm | -37.4 dB |
| Noise Figure | 22 dB | 4.8 dB | -17.2 dB |
| DANL | -152 dBm/Hz | -169.2 dBm/Hz | -17.2 dB |
| Max 3 rd order dynamic range | 117.3 dB | 111.8 dB | -5.5 dB |
| Max 2 nd harmonic dynamic range | 101 dB | 90.9 dB | -10.1 dB |

Table 6-1. Preamp Dynamic Range Results

Indeed, the system noise figure improves by 17.2 dB, but the system TOI degrades by 25.5 dB and system SHI degrades by 37.4 dB. The resulting dynamic range degradations for 3rd order IMD and 2nd harmonic are 5.5 dB and 10.1 dB respectively.

Figures 6-33 and 6-34 show the dynamic range charts for 3rd order IMD and 2nd harmonic distortion for the signal analyzer with and without the cascaded preamplifier. The x-axis requires some explanation. With the preamplifier in the system, the x-axis represents the power level at the input of the preamplifier. Note the degradation in dynamic range with the preamplifier on. Also note how much lower the power level must be with the preamplifier in the system for best dynamic range.



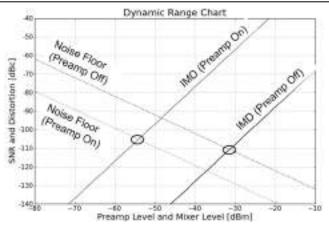


Figure 6-33. Preamp 3rd Order IMD and the Dynamic Range Chart

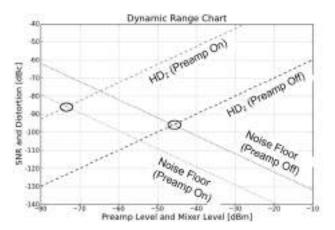


Figure 6-34. Preamp 2nd Harmonic Distortion and the Dynamic Range Chart

So the answer to the above question is no, the preamplifier does not improve dynamic range. The degradation in distortion is worse than the improvement in noise figure. Use the preamplifier for its intended purpose, which is to improve system sensitivity for the measurement of near noise signals. Do not use it in the presence of large amplitude signals that may produce distortion.

Near Noise Distortion Measurements

In reality, as the CW distortion signal approaches the noise floor the signal and noise combine. In the dynamic range chart, a true intersection to the distortion and noise curves does not actually occur. **Figure 6-35** demonstrates the effect of near-noise measurements.



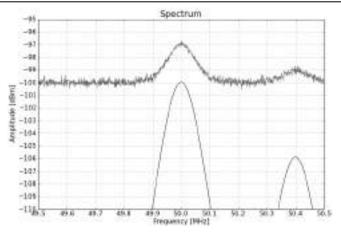


Figure 6-35. Near Noise Signals

The signal at 50 MHz in **Figure 6-35** has an amplitude equal to the noise floor. However, the noise floor shows a 3 dB addition. The dynamic range charts as drawn so far would not show this 3 dB noise bump. Even when the signal is below the noise floor, the noise floor shows a bump. At 50.4 MHz in **Figure 6-35**, the signal is 6 dB below the noise floor, yet the noise floor shows a 1 dB bump above this signal.

Figure 6-36 shows the combined signal plus noise near the intersection of the distortion and noise curves.

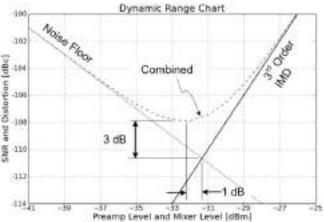


Figure 6-36. Near Noise Signals on the Dynamic Range Chart

Note that the signal plus noise adding in the manner described is only valid for RMS averaged noise. For 3rd order IMD and noise, the maximum dynamic range occurs one dB lower in mixer level than the optimum mixer level described by **Equation 6-8**. For second order distortion, the optimum mixer level described by **Equation 6-10** is still valid when considering the distortion plus noise curve.

RF Input Attenuator Effect on Dynamic Range

The RF input attenuator affects dynamic range in two different ways, the first way being the range of RF input attenuation values. The highest performance signal analyzers have 70 dB RF input attenuation range. With +30 dBm maximum measureable input power, the 70 dB allows for a -40 dBm mixer level. Using the dynamic range chart of **Figure 6-27** as a guide, optimum mixer levels for 3rd order IMD falls in the -30 dBm to -40 dBm mixer level range. RF input attenuator range is set in order to measure the highest measureable signals with best dynamic range.



The other way the RF input attenuator affects dynamic range is with its step resolution. The maximum dynamic range only occurs if the mixer level can hit the optimum value. With coarse RF input attenuator resolution, the optimum mixer level may never be achieved. **Figure 6-37** demonstrates the potential unachievable dynamic range as a function of RF input attenuator resolution.

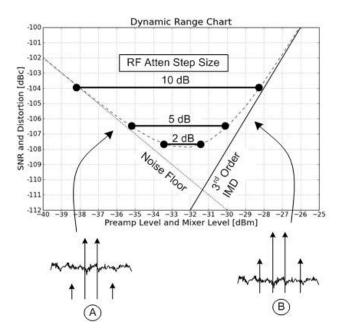


Figure 6-37. RF Input Attenuator Step Size vs. Dynamic Range

In worst case scenarios, the mixer level could bounce back and forth between noise limited dynamic range (graphic A in **Figure 6-37**) and distortion limited dynamic range (graphic B in **Figure 6-37**).

Table 6-2 summarizes the potential dynamic range that may be given up versus RF input attenuator step size.

| | Potential | Potential |
|-----------|-------------|-------------|
| | Unavailable | Unavailable |
| | Dynamic | Dynamic |
| RF Atten. | Range Range | |
| Step Size | (Ideal) | (Actual) |
| 10 dB | 6.7 dB | 4.08 dB |
| 5 dB | 3.3 dB | 1.29 dB |
| 2 dB | 1.3 dB | 0.23 dB |
| 1 dB | 0.7 dB | 0.06 dB |

Table 6-2. Potential Unavailable Dynamic Range vs. RF Input Attenuator Step Size

In **Table 6-**2 the *ideal* column uses the unrealistic noise and 3rd order IMD lines intersecting with no nearnoise addition and the *Actual* column is with the noise and distortion curves adding, which is the true behavior of the signal analyzer. The *Ideal* column could also be used to determine the dynamic range change when operating in region A of **Figure 6-27**.



One can see that there is a diminishing return on dynamic range with smaller RF input attenuator step size. When considering optimization at the maximum dynamic range location, a two dB RF input attenuator resolution may be sufficient.

ADC Contribution to Dynamic Range

As briefly mentioned in section 4, the ADC has a very real impact on the signal analyzer's dynamic range. How the ADC contributes to the dynamic range performance of the signal analyzer is now presented.

Representative spurious-free dynamic range (SFDR) data for an ADC is shown in Figure 6-38 [¹⁶].

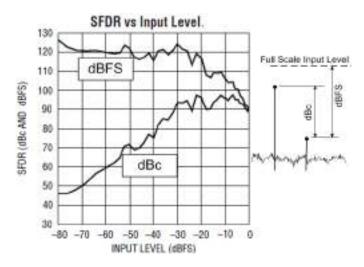


Figure 6-38. ADC Spurious-Free Dynamic Range

The two lines are different viewpoints of the same data. The spectrum graphic in **Figure 6-38** helps explain the meaning of the two curves. The *dBFS* curve is the distortion amplitude relative to the maximum ADC input amplitude (full scale input level). The *dBc* curve is the distortion amplitude relative to the fundamental signal. The *dBc* curve is more relevant to the mapping of the ADC distortion onto the dynamic range chart.

Studying this ADC SFDR data reveals that for the fundament signal amplitude near the full scale level that the distortion improves as the fundamental amplitude decreases. However, below -15 dBFS, the distortion amplitudes do not decrease with decreasing fundamental signal amplitude. Reference [⁶] describes the reasons for this behavior in the pipelined ADC topology which is primarily used for IF sampling in today's signal analyzer structures.

Figure 6-39 demonstrates how the two different behavioral regions of the ADC SFDR data relate to the dynamic range chart.



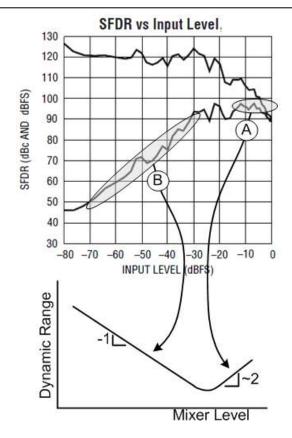


Figure 6-39. ADC Spurious-Free Dynamic Range Mapped to the Dynamic Range Chart

In region A, the dynamic range of the ADC improves with decreasing mixer level. Mixer level in this case is directly related to the ADC input power through the IF gain of the signal analyzer. It is assumed that there is enough IF gain to drive the signal to the ADC's full scale input level. The slope of the dynamic range curve in region A is approximately two for two-tone IMD. In region B, the distortion amplitudes stay fixed for any given fundamental signal level. The slope for the dynamic range curve is -1. This is the same slope as the noise curve, but keep in mind that it is ADC distortion that is being discussed here.

The ADC distortion looks pretty damaging to the dynamic range performance of the signal analyzer, and it can be if the signal analyzer settings are not controlled properly. It is here that we must consider vector signal analysis separately from spectrum analysis.

ADC Distortion Associated with Vector Signal Analysis

Recall that the term vector signal analysis refers to a relatively wide analog IF bandwidth where the signal and distortion components fall inside this bandwidth. In this case the SFDR of the ADC dominates distortion performance for multi-tone intermodulation distortion.

In **Figure 6-40**, the ADC distortion curve is shown with the frontend distortion. These curves are for twotone IMD.



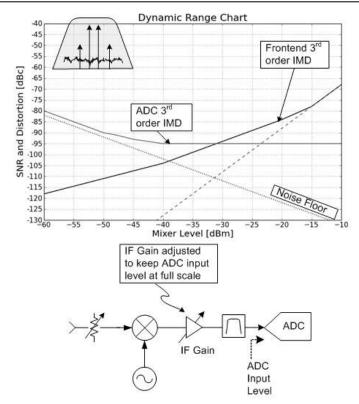


Figure 6-40. Dynamic Range Chart in VSA Mode. IF Gain Constantly Adjusted.

The overall system distortion will be the worst of the ADC and frontend distortion. It is broken out in **Figure 6-40** for discussion purposes. In this particular case, the IF gain is adjusted at each mixer level setting to ensure that the fundamental signal amplitudes are brought close to the ADC's full scale level. The IF gain is indirectly controlled with the reference level. When operating near full scale, the ADC's SFDR is maximized. At a certain mixer level, approximately -40 dBm in **Figure 6-40**, the IF gain maxes out. Once there is no more IF gain to offer, the ADC can no longer operate at full scale and the SFDR performance begins to drop.

Notice the frontend IMD performance in **Figure 6-40**. As the mixer level drops and more IF gain is requested, the final analog amplifiers now begin to dominate the frontend TOI performance. Each device contributes to the system TOI as shown in **Equation 6-13**.

$TOI_{sys} = TOI_{device} - Gain_{before device}$

Equation 6-13

As the system gain in front of the device increases, which is the case for low reference level settings, the device's TOI becomes a larger contributor to the overall system TOI.

Figure 6-41 shows what can happen when the user is not careful in adjusting IF gain (via the reference level setting) to ensure near full scale ADC operating level.



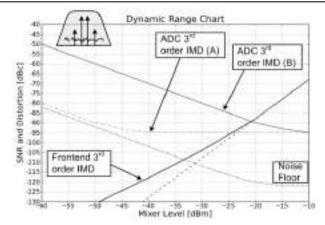


Figure 6-41. Dynamic Range Chart in VSA Mode. IF Gain not Adjusted.

The line labeled *ADC* 3^{rd} order *IMD* (*A*) is the case from **Figure 6-40** where IF gain adjustment takes place. However, the line labeled *ADC* 3^{rd} order *IMD* (*B*) is where the reference level setting is fixed at the highest mixer level value shown (-10 dBm in this case). Now the ADC SFDR completely dominates the signal analyzer's dynamic range performance.

Careful comparison between **Figure 6-40** and **Figure 6-41** shows that the frontend IMD degradation is not as severe when the IF gain is not constantly being adjusted. From **Equation 6-13**, less gain in front of the device equates to better system TOI performance. The noise floor at high mixer levels does start to degrade. Noise is opposite from distortion in that more system gain is required in front of any given device to make that device's noise contribution to the overall system noise as small as possible. ADCs have very large noise figures (> 30 dB) requiring a great deal of up front system gain to ensure the ADC's noise is not dominating the system noise performance. At higher mixer levels, the IF gain is lower and the ADC's noise starts to dominate. This, however, is not really a problem for system dynamic range performance as the distortion at high mixer levels dominates over noise.

To summarize in vector signal analysis mode the wide analog IF bandwidth exposes the ADC to the fundamental signals during the measurement of the distortion components. Because of this, the dynamic range of the signal analyzer is limited to the SFDR performance of the ADC. However, in many actual applications, the ADC cannot be made to operate at full scale input level, and then the true SFDR degrades.

ADC Distortion Associated with Spectrum Analysis

For spectrum analysis, the analog IF bandwidth is narrow enough to select only the fundamental signal or the distortion signal at any given time. **Figure 6-42** demonstrates the effect of narrowband filtering in the final IF of the signal analyzer.

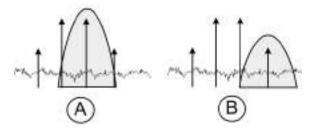




Figure 6-42. IMD Measurements using Spectrum Analysis Mode.

At point A in **Figure 6-42**, one of the fundamental signals in this two-tone IMD measurement is selected. At point B, the distortion product is being measured. The IF filter bandwidth is narrow enough that all signals, except the distortion tone being measured, are attenuated.

Figure 6-43 shows that the location of the analog IF filters should be placed before both the ADC and the final analog IF gain stages. In fact, placing the IF filters as far up the signal chain as possible is a goal. However, practical reasons of ensuring high enough filter quality factor, Q, for the narrowest bandwidths usually restricts placement of the IF filter in the final IF.

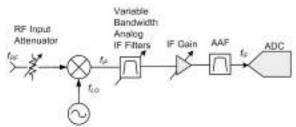


Figure 6-43. Final IF Amplifier and ADC Affected by Analog IF Filters

By not allowing the relatively large amplitude fundamental signal to reach both analog devices and the ADC, which sit subsequent to the narrowband IF filter, gives the signal analyzer several advantages over the wideband IF used in vector signal analysis mode:

- The chances of gain compression and worse, delivering signals whose amplitude are greater than the ADC's full scale input level, are diminished
- The effective TOI of the devices following the IF filter is dramatically improved
- The IF gain can be increased which makes the ADC noise figure contribution to the overall system noise much less

Of the advantages listed, the effective TOI improvement is the one that renders itself to a quantitative analysis.

Effective TOI Improvement for Analog Devices

Figure 6-44 is used to illustrate how the analog narrowband IF filters improve the effective TOI of an analog active device. A two-tone signal drives the input of a bandpass filter with 'R' dB of stopband rejection placed in front of the analog gain device.

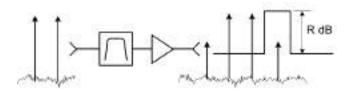


Figure 6-44. Effective TOI Enhancement with Filtering in Front of Analog Device

From **Equation 6-4**, the third order IMD products resulting from fundamental signals with amplitude, A, are described by **Equation 6-14**.

$$IMD_{lower} = kA^{3}cos[(2\omega_{1} - \omega_{2})t]$$
$$IMD_{upper} = kA^{3}cos[(2\omega_{2} - \omega_{1})t]$$



where k is a proportionality constant let r be the linear representation of the filter rejection.

Equation 6-14

$$r = 10^{-R/20}$$

Equation 6-15

With the filter, the IMD signal equations become Equation 6-16.

 $IMD_{lower} = k(rA)^3 cos[(2\omega_1 - \omega_2)t]$ $IMD_{upper} = k(rA)^3 cos[(2\omega_2 - \omega_1)t]$

Equation 6-16

The ratio of the amplitudes of the IMD products with the filter and without the filter is shown in **Equation 6-17**.

$$\frac{IMD \text{ amplitude with filter}}{IMD \text{ amplitude without filter}} = \frac{kr^3A^3}{kA^3} = r^3$$

Equation 6-17

On a log scale, the IMD amplitudes are reduced by *R dB*. Without filtering, TOI of the active device is given by **Equation 6-18**.

 $TOI_{device} = P_{fund} + \frac{P_{fund} - P_{IMD}}{2}$

where P_{fund} is the power level of the fundamental signal in dBm and P_{IMD} is the power level of the distortion product.

Equation 6-18

Adding the filter gives an effective TOI shown in Equation 6-19.

$$TOI_{effective} = P_{fund} + \frac{P_{fund} - (P_{IMD} - 3R)}{2}$$
$$TOI_{effective} = TOI_{device} + \frac{3}{2}R$$

Equation 6-19

Equation 6-19 demonstrates that by placing a filter in front of an analog device the effective TOI of the combination increases by 1.5 times the filter rejection in dB. This assumes both fundamental tones are attenuated by the same amount R. Actual filters normally do not have constant stopband rejection, so the results in equation may be conservative.

The effective TOI improvement given by **Equation 6-19** is quite startling. Designers of active devices go through heroic efforts to improve the TOI by a modest 2 or 3 dB. To yield a 30 dB improvement using very moderate 20 dB filter rejection is quite profound.



Effective TOI Improvement for ADCs

The act of placing a filter in front of the ADC also improves effective TOI. However, a closed form equation cannot be rendered. The analysis is ADC device dependent and involves careful study of the particular device's SFDR chart similar to that shown in **Figure 6-39**.

To illustrate the steps in determining the effective TOI of an ADC with a bandpass filter placed in front of it, **Figure 6-45a** and **Figure 6-45b** is used.

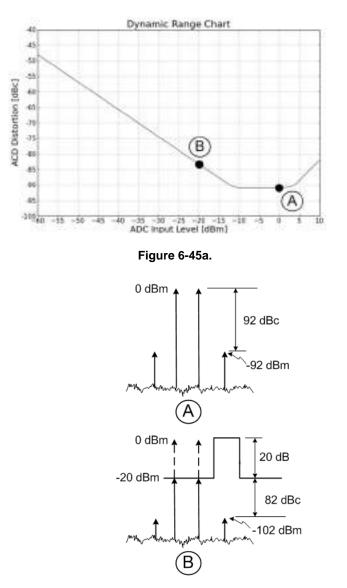


Figure 6-45b. Effective TOI Enhancement with Filtering in Front of an ADC

In **Figure 6-45b**, the graphic labeled A, shows the fundamental signals placed at 0 dBm at the ADC input. **Figure 6-45a** shows the position of A on the dynamic range chart. **Figure 6-45a** is the ADC's SFDR data mapped to the dynamic range chart. The x-axis is the ADC input level in dBm with a full scale input level of +10 dBm. So 0 dBm in graphic A corresponds to -10 dBFS. The distortion product amplitudes are -92 dBc, or -92 dBm with a 0 dBm fundamental signal level.



In graphic B, the bandpass filter with constant stopband rejection of 20 dB is placed in front of the ADC. During the portion of the measurement where the IMD products are being measured, the fundamental tones are attenuated be the 20 dB filter stopband rejection. **Figure 6-45a** shows that at point B, the ADC distortion is approximately 82 dB, which places the distortion amplitudes at -102 dBm.

The distortion measurement is a two step process: first, measure the fundamental signal amplitude and then measure the IMD product amplitude. From a measurement prospective, the IMD amplitude is relative to the unfiltered amplitude of the fundamental signal. Without the filter, the dynamic range of the IMD measurement is 92 dB, which is the limit of the ADC's SFDR performance. With the filter, the dynamic range is roughly 10 dB better, which is an effective TOI improvement of 5 dB.

Empirical measurements bear out this roughly 10 dB improvement in dynamic range performance when using narrowband IF filtering compared to wideband filtering used for vector signal analysis measurements.

Reference Level Re-Ranging

Narrowband filtering certainly helps improve distortion performance of the ADC. However, using the filter alone still limits the overall signal analyzer's dynamic range. The dynamic range charts provided in this chapter show the frontend capable of greater than 110 dB of maximum third order dynamic range. Modest TOI of +18 dBm and DANL = -150 dBm/Hz (rms) yields 112 dB dynamic range using **Equation 6-7**. The filtered ADC still has 10 dB to 15 dB worse dynamic range than the frontend. To make the ADC's contribution to system level dynamic range small compared to the frontend, the extra step of so-called *reference level re-ranging* needs to take place.

Reference level re-ranging refers to using two different reference level settings, one for the measurement of the fundamental signals and one for the measurement of the distortion signals. The principle behind this relies on the fact that the ADC's distortion level measured in dB relative to full scale (dBFS) stays constant when operating 15 dB or so below full scale amplitude, shown in **Figure 6-38**.

Figure 6-46 illustrates an example of using reference level re-ranging to improve the effective dynamic range of the signal analyzer.

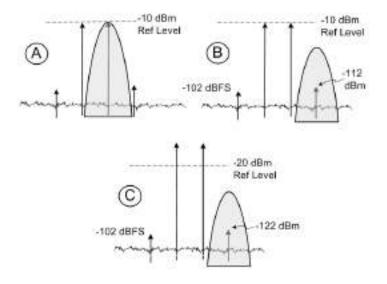




Figure 6-46. Re-Ranging the Reference Level to Improve ADC Effective TOI

Keep in mind that reference level is an indirect control of the IF gain. To be effective, most of the IF gain variation needs to take place in the stages that follow the narrowband analog IF filters.

In graphic A, the fundamental signal is being measured. The reference level is set so as to not overdrive the ADC's full scale input level. In graphic B, the distortion product is being measured without reference level re-ranging. In graphic C, the reference level has been re-ranged by 10 dB in this case. Because the large fundamental tones are attenuated by the filter's stopband rejection, the ADC does not overload, nor do the analog gain stages following the filter add distortion. Between graphics B and C, the ADC's distortion amplitude stays fixed at the same value relative to the ADC's full scale input level. However, the displayed amplitude of the distortion product is offset lower in value by the reference level change; 10 dB lower in this example.

Figure 6-47 summarizes the ADC performance for the various IF filtering modes. Wideband filtering, assuming that the reference level is constantly adjusted to bring the input signal at the ADC to the ADCs full scale input level results in system dynamic range limited to the performance of the ADC. This is usually in the 80 dB to 90 dB range. Using a narrowband filter for spectrum analysis mode results in approximately 10 dB improvement in dynamic range. As discussed, the best performance requires reference level re-ranging.

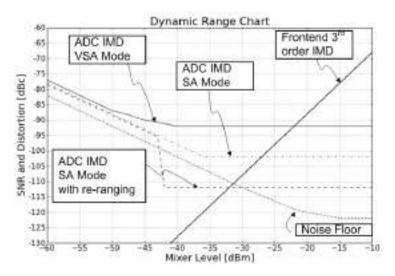


Figure 6-47. Summary of IF Filtering Effect on ADC IMD

The case for reference level re-ranging reverts back to the case of standard non re-ranging when there is no longer enough analog gain to support this mode. If the mixer level at the system, runs out of extra analog gain, falls below the optimum mixer level for maximum frontend dynamic range, then no loss in dynamic range performance is incurred.

Dynamic Range Considerations for Digitally Modulated Signals.

Nearly all of the discussion up to this point has involved only the analysis of CW signals. However, modern communications signals use digital modulation, which is relatively wideband. So much attention has been given to CW signals in this chapter that a question of relevance comes to mind. Several reasons for the detailed consideration of CW signals for the signal analyzer are as follows:



- 1. The signal analyzer, by tradition, is calibrated for the amplitude of a CW signal. This tradition will be hard to break.
- 2. For distortion and noise analysis, the digitally modulated signal can be broken down into CW sub-components.
- 3. Much of the design work for transmitters and receivers, even when their intended purpose is for digitally modulated signals, rely on CW characterization. Most often solid state devices such as mixer and amplifiers have only CW distortion performance specifications.

The dynamic range measurement considered in the following discussion is the measurement of spectral leakage into adjacent channels. **Figure 6-48** illustrates spectral re-growth generated by the signal analyzer. Suppose a distortion free digitally modulated signal is introduced at the input of the signal analyzer. Intermodulation distortion in the signal analyzer causes the shoulders seen in the adjacent channels. Primarily, this is third order distortion, but 5th and sometimes 7th order distortion components contribute to the spectral re-growth. Additionally, broadband noise can degrade SNR. The signal analyzer's phase noise is one more component that can become a limitation.

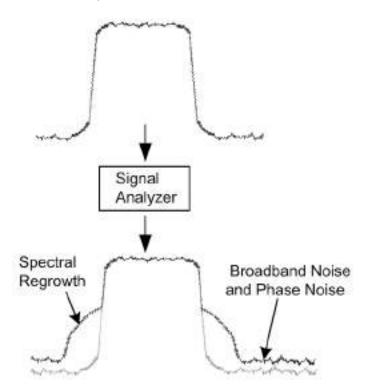


Figure 6-48. Spectral Re-growth Generation in the Signal Analyzer.

The measurement at hand is the characterization of a DUT. To ensure that the measured spectral regrowth is due to the DUT and not the signal analyzer, the signal analyzer's spectral re-growth must be made as small as possible.

One common metric for spectral re-growth is *adjacent channel power ratio* (ACPR) and *adjacent channel leakage ratio* (ACLR). In general, this measurement is the power in the main channel divided by the power in either the lower or upper adjacent channels. **Figure 6-49** shows the spectrum of this kind of measurement.



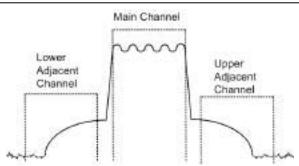


Figure 6-49. Defining Terms for the ACLR Measurement.

Different communication standards define their own unique methodologies in terms of the bandwidths used in integrated power in the adjacent channels. Some standards do not require integration for the power measurement, but rather define spot frequency measurements in clearly defined RBW and VBW filter settings. No matter the exact power ratio measurement methodology, the general ideas on signal analyzer generated noise and distortion associated with digitally modulated signal outlined in the following discussion are relevant.

Once again, whether the measurement uses wideband IF or narrowband IF, the dynamic range results and the measurement technique are different. Wideband in this case means that the IF bandwidth is wide enough to capture both the main channel and at least one of the adjacent channels. With wideband analysis (vector signal analysis) the intent is to rapidly measure both the main and the adjacent channels simultaneously. Narrowband IF in this case refers to the analog IF bandwidth significantly narrower than the main channel bandwidth. The intent for narrowband analysis (spectrum analysis) is that the main channel signal power is attenuated when measuring the adjacent channels. The concepts of wide and narrow band analysis are depicted in **Figure 6-50a and Figure 6-50b**.

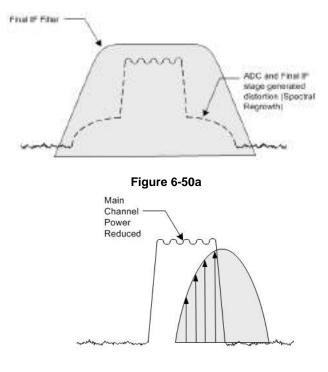


Figure 6-50b



Figure 6-50. Wide and Narrow IF Bandwidth Analysis on Modulated Signals

The goal here is to be able to represent the ACLR problem on the dynamic range chart for the same reasons as was done for the CW signal case, which is to determine maximum dynamic range and the associated optimum mixer level. For signal analyzers operating below gain compression, the analysis of internally generated distortion and noise can be approximated to a high degree of accuracy by breaking the digital signal down into a series of CW signal equivalents. First, the dynamic range chart's signal to noise ratio curve is analyzed.

SNR for Digitally Modulated Signal

In **Figure 6-8**, the idea behind the total integrated power of a signal with modulation was first presented. This partially describes the signal-to-noise ratio of the digitally modulated signal. **Figure 6-51** graphically describes SNR associated with digitally modulated signals.

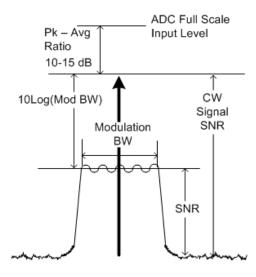


Figure 6-51. SNR of the Digitally Modulated Signal

As discussed earlier, the frontend components react according to total integrated power. The difference between a CW signal's power level and a digitally modulated signal's power level is approximately $10\log(BW_m)$, where BW_m is the modulation bandwidth. This can be significant: for $BW_m = 1$ MHz, this is a 60 dB difference; for a $BW_m = 20$ MHz, this is a 73 dB difference. For wide IF bandwidth, the crest factor or peak-to-average power ratio (PAPR) must be taken into account. Analog filtering will reduce the PAPR by a few dB, but to be safe it is wise to back off the mixer level by the PAPR of the signal.

After backing off by the PAPR and accounting for the modulation bandwidth, the SNR, as shown in **Figure 6-51**, is 75 dB to 90 dB lower than the SNR for an equivalent power CW signal.

For the CW signal, the SNR is a function of the resolution bandwidth filter setting. Not so for the digitally modulated signal. The signal itself is noise-like and of course the noise floor is noise-like. Both the signal's displayed power and the noise floor displayed power vary by 10log(RBW) as shown in **Figure 6-52**. SNR for the digitally modulated signal is independent of RBW setting.



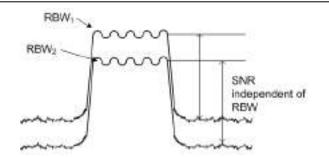


Figure 6-52. SNR for the Digitally Modulated Signal is Independent of RBW Setting.

As an example for the calculation of SNR, use the parameters of a W-CDMA signal: $BW_m = 3.84$ MHz and PAPR is approximately 12 dB. Also suppose the signal analyzer's maximum mixer level is -10 dBm and its displayed average noise level is -155 dBm/Hz. For a CW tone, the maximum SNR is as follows:

 $SNR_{CW,max} = Mixer \ Level_{max} - DANL$ $SNR_{CW,max} = -10 - (-155) = 145 \ dB$

The maximum SNR for the digitally modulated signal is as follows:

$$SNR_{dig,max} = SNR_{CW,max} - 10 \log(BW_m) - PAPR$$

$$SNR_{dig,max} = 145 - 65.8 - 12 = 67.2 dB$$

This is somewhat conservative as normally there is some headroom already built in to the system gain such that at maximum mixer level, the ADC is several dB away from being overdriven. What this means is that in many signal analyzers, the SNR of the digitally modulated signal does not need to be backed off by the full PAPR of the signal.

IMD for Digitally Modulated Signal

Normal operating mode for the signal analyzer places the maximum signal power below the signal analyzer's gain compression level. As in the case of CW signals, the signal analyzer can be considered a weakly nonlinear system. Give this constraint, the spectral re-growth stemming from intermodulation distortion for the digitally modulated signal can be approximated using the following analysis:

First divide the spectrum of the digitally modulated signal into a series of equal frequency slices as shown in **Figure 6-53**.

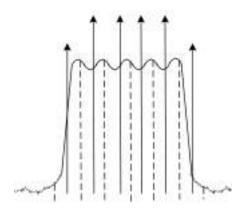




Figure 6-53. Digitally Modulated Signal Represented by CW Signal Equivalents.

For each frequency slice, compute the integrated power using as shown in Equation 6-20.

$$P_{slice} = 10\log\left(\sum_{i=f_a}^{f_b} 10^{X(f_i/10)}\Delta f\right)$$

where X(f) is the frequency domain representation of the signal, f_a and f_b are the frequency limits of the slice, and Δf is the frequency width of adjacent sample points.

Equation 6-20

Next, represent each frequency slice as a CW signal whose power level is P_{slice} . Place the CW signals frequency in the center of each associated frequency slice.

Using **Equation 6-6**, calculate the amplitude of the third order intermodulation distortion products. Use every possible combination of pairs the CW signals that represent the digitally modulated signal. The frequencies of each distortion product can be calculated by manipulating **Equation 6-4**. The distortion products will start falling at the same frequencies as other products. **Figure 6-54** shows this process.

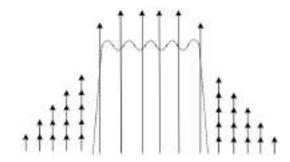


Figure 6-54. Spectral Regrowth from Two-tone IMD

Empirical evidence shows that the amplitudes of the multiple distortion products add as in-phase vectors. At each distortion component frequency the summed distortion power is calculated using **Equation 6-21**.

$$P_{IMD} = 20 \log \left(\sum_{i} 10^{P_i/20} \right)$$

where P_i is the power in dBm of each individual distortion component.

Equation 6-21

For the case of W-CDMA, experimentally it has been found that the spectral re-growth power integrated over the adjacent channel can be approximated using **Equation 6-6**. The TOI value used is the signal analyzer's TOI with an experimentally determined offset. This offset is a function of PAPR, but is roughly 2 dB to 5 dB.

$$P_{IMD,adjacent} = 3P_{Mix \ Lvl} - 2(TOI_{CW} + TOI_{offset})$$



where $P_{Mix Lvl}$ is the power of the digitally modulated signal at the signal analyzer's first mixer.

Equation 6-22

Suppose the TOI of the signal analyzer is +20 dBm, the signal power is 0 dBm, the RF input attenuator is 10 dB, and the TOI offset value is approximately 4 dB. The mixer level is -10 dBm. From **Equation 6-22**, the integrated spectral re-growth in the adjacent channel is -78 dBm, giving a relative value of -68 dB.

Phase Noise for Digitally Modulated Signal

Phase noise cannot be ignored for the measurement of adjacent channel spectral re-growth. **Figure 6-55** shows the process of determining the phase noise contribution of the digitally modulated signal in the signal analyzer.

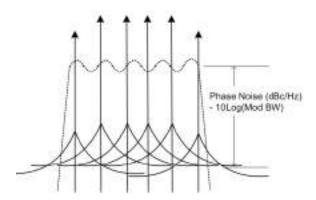


Figure 6-55. Signal Analyzer Phase Noise for the Digitally Modulated Signal.

Using the model of the digitally modulated signal, which uses CW signal representation, superimpose the signal analyzer's phase noise skirt onto each of the CW signals. At a single frequency in the adjacent channel, add the phase noise powers using **Equation 6-22**.

$$P_{Phase \ Noise} = 10 \log \left(\sum_{i} 10^{P_i/10} \right)$$

where $P_{Phase Noise}$ is the summed phase noise power in a 1 Hz bandwidth and P_i is the phase noise power in dBm/Hz associated with each CW signal.

Equation 6-22

Quite often the phase noise pedestal of the signal analyzer is flat with frequency for offset frequencies beyond 1 MHz. For most wide bandwidth (> 1 MHz) modulation formats, this wider offset phase noise is the main contributor. Given a flat phase noise pedestal, the phase noise contribution to spectral regrowth is approximately the signal analyzer's far offset phase noise in dBc/Hz – $10\log(BW_m)$.

Suppose the far offset phase noise of the signal analyzer is -150 dBc/Hz and, using W-CDMA as an example, the modulation bandwidth is 3.84 MHz. The relative spectral re-growth due to phase noise is roughly -84 dB.



ACLR using Vector Signal Analysis Mode

When using vector signal analysis mode, the mixer level and reference level need to be the same between the measurement of the main channel power and the adjacent channel power. The techniques outlined so far can be used for optimizing the RF input attenuator and reference level for best ACLR performance.

The dynamic range chart in **Figure 6-56** shows ACLR performance when the signal analyzer is configured for vector signal analysis mode.

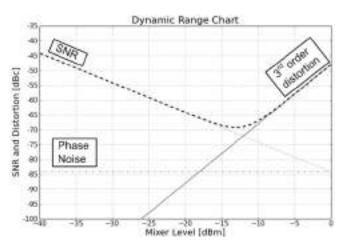


Figure 6-56. ACLR VSA Mode

In this example, the mixer level that yields best ACLR is around -13 dBm, yielding W-CDMA ACLR performance of around 70 dB.

ACLR using Spectrum Analysis Mode

In spectrum analysis mode, the IF bandwidth is set narrower than the signal's modulation bandwidth. During the measurement of the adjacent channel, the main channel signal is greatly attenuated before reaching the final IF amplifier stages and the ADC. This allows the reference level to be set lower, corresponding to higher IF gain. Not only is the distortion performance improved, but the overall SNR is better. The result is a 5 dB or more improvement in ACLR performance. **Figure 6-57** shows the dynamic range chart for W-CDMA ACLR when narrowband IF filters are used.

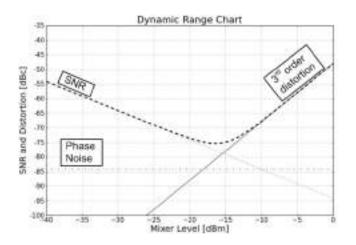




Figure 6-57. ACLR Spectrum Analysis Mode

Notice that the phase noise is no longer invisible to the measurement. The phase noise falls 5 dB below the intersection of the SNR and IMD curves. Phase noise in this example contributes approximately one dB to the ACLR degradation.

Summarizing the Dynamic Range Information

Generating these dynamic range graphs requires knowledge of the internal workings of the signal analyzer. At the very least it may require measurement of the signal analyzer's DANL and TOI beyond what is posted in the specification sheets. This section has mostly shown that optimizing the signal analyzer is not a trivial task. To extract the best dynamic range out of the signal analyzer, some work is involved. Simply letting the signal analyzer stay in its default attenuator and gain setting will result in only modest performance.

The major points of this dynamic range section are as follows:

- Mixer level is the most important item for achieving best dynamic range
- Wide bandwidth vector signal analysis mode dynamic range is limited by the ADC SFDR
- When in vector signal analysis mode, the reference level must be adjusted to the test signals power level. This sets the signal at the ADC to full scale, which results in the best SNR and SFDR performance.
- Narrow bandwidth spectrum analysis mode allows dynamic range beyond that of the ADC
- Re-ranging the reference level when measuring the distortion amplitude allows the best dynamic range
- ACLR measurements for digitally modulated signals follow the same techniques for optimizing dynamic range as for the measurement of CW signal IMD.

Follow the key concepts of adjusting mixer level using either the RF input attenuator or by adjusting the test signal amplitude. Then set the reference level to the power of the test signal. Those are guidelines; the rest is trial and error.

Attenuator Test

In the course of optimizing the signal analyzer for dynamic range, the attenuator test can be used as a check to determine whether the measured distortion is due to the test signal or due to the signal analyzer. When viewing the distortion amplitude, toggle the RF input attenuator setting. If the measured amplitude of the distortion amplitude does not change, then there is a high assurance that the distortion is entirely due to the test signal. However, if the distortion amplitude changes, then the signal analyzer is contributing to the measurement.



7. Amplitude Accuracy

In addition to possessing high dynamic range for the measurement of distortion, the signal analyzer also serves to accurately measure the signal amplitude. Attention now focuses on the uncertainty associated with the measurement of signal power.

Absolute power measurement refers to the single measurement of signal power. This entails recording the signal amplitude reported by the signal analyzer. Uncertainty is traceable back to some power standard, most commonly the National Institute of Standards and Technology (NIST).

Relative power measurement refers to two or more power measurements. For example, the harmonic measurement is a good example of a relative power measurement. Many of the absolute amplitude measurement uncertainty terms are common between the measurements and do not contribute to the relative measurement uncertainty. However, there are multiple measurements, so some of the uncertainties add up.

Absolute Amplitude Accuracy at the Calibration Reference Frequency

Many signal analyzers break up frequency response into two parts: one is sometimes termed *absolute amplitude accuracy at the reference frequency* and the other is termed the *frequency response*. Signal analyzers whose amplitude uncertainty is specified in this manner usually have an internal calibration signal. At the single point calibration frequency, an amplitude uncertainty is given.

Frequency Response

Often the frequency response is relative to the amplitude at the calibration reference frequency. **Figure 7-1** shows how frequency response is related to the absolute amplitude accuracy at the calibration reference frequency.

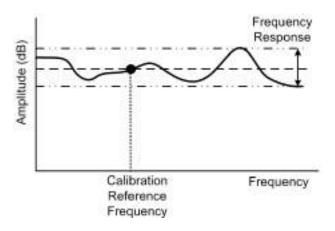


Figure 7-1. Frequency Response Relative to the Calibration Reference.

For single point power measurements at any given frequency, the frequency response uncertainty must be added to the absolute amplitude uncertainty. For relative power measurements involving the power ratio of multiple measurements at different frequencies, only the frequency response uncertainty applies.

IF Amplitude Response

Frequency response is measured at the center of the IF passband. Think of the IF passband being superimposed on the frequency response curve as suggested in **Figure 7-2**.



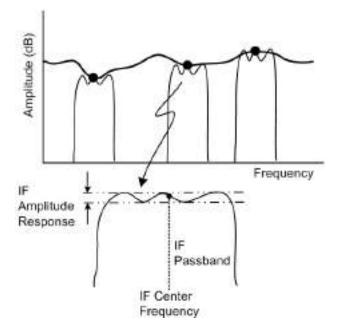


Figure 7-2. IF Amplitude Response

For signal analyzers that use FFT processing to create a spectrum display, the spectrum consists of multiple FFT segments cascaded together as shown in **Figure 7-3**.

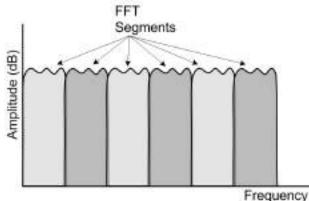


Figure 7-3. Spectrum Created from Multiple FFT Segments

When the measurement frequency is offset from the center of a particular FFT segment, the *IF amplitude response uncertainty* must be included. IF amplitude response is relative to the center of the IF passband as shown in **Figure 7-2**.

To eliminate the uncertainty associated with the IF amplitude response, set the frequency span to be narrow enough to ensure only one FFT segment is being displayed. This usually means that the span is narrower than the analog IF bandwidth. Then, adjust the center frequency setting such that the signal of interest is in the center of the display.

Signal Processing Amplitude Uncertainties



For signal analyzers that use FFT processing for displaying the spectrum, the scalloping loss mentioned in section 5 and summarized in **Table 5-1** must be added to the uncertainty. The amount of scalloping loss depends on the windowing function selected.

RF Input Attenuator Switching Uncertainty

The specification sheet for the signal analyzer may or may not include information for RF input attenuator switching uncertainty. In either case, the absolute amplitude accuracy at the calibration reference frequency and the frequency response are qualified for a particular RF input attenuation value. For the single frequency absolute amplitude measurement, it is best to use the RF input attenuation specified. However, for relative amplitude measurements, this uncertainty does not occur if the same RF input attenuator setting is used for all measurements.

The absolute measurement may need a different RF input attenuation value than what is used for the amplitude accuracy specifications. Likewise, the RF input attenuation value may need changing between measurements. If this is the case and the user wants to eliminate this error, then an RF substitution technique can be used. At the measurement frequency, set the RF input attenuator to value 1 and make an amplitude measurement. Then switch to RF input attenuator value 2 and measure the amplitude. The difference in amplitudes between these two measurements is the RF input attenuator switching error. The error can be subtracted from all subsequent measurements.

YIG Tuned Filter Amplitude Response

As mentioned in Section 3 of this document, the highband path in the signal analyzer may include a YIG tuned filter (YTF) as a preselector. Unfortunately, the frequency response when selecting the YTF can be compromised.

The YTF is an open loop device meaning that there is no feedback mechanism to ensure that the filter center frequency is precisely deterministic. When the signal analyzer is tuned to a particular highband frequency, the YTF more than likely will not be centered where it was during instrument calibration. Furthermore, within a measurement session the YTF's center frequency is not repeatable. Tuning the signal analyzer from frequency A to frequency B and then back to frequency A again most likely results in a YTF center frequency offset.

The lack of YTF repeatability stems from several mechanisms. One is tuning hysteresis. The YTF center frequency is controlled by an electro-magnet which has its associated hysteresis: the magnetic flux strength at a particular current value differs depending on whether the current is being ramped up or being ramped down. The different magnetic flux translates to a YTF center frequency offset depending on which way the YTF is being tuned. Take for example the signal analyzer tuned to 10 GHz. If the previous center frequency was 5 GHz the YTF center frequency will be at one frequency. If the previous center frequency was at 15 GHz, then the YTF center frequency may be different. Efforts are made to reduce the hysteresis using a de-hysteresis pulse between each center frequency change.

Post tune drift is another repeatability mechanism. A fair amount of heat is generated in the electromagnet in the YTF. The heat causes center frequency drift (the YTF metal housing expands with heat and this tends to concentrate the magnetic flux). The higher the tune frequency, the more magnet current and hence the more heat generated. In addition to the YTF center frequency drifting over time at elevated frequencies, the heat is retained when tuning to lower frequencies. At these lower frequencies, the YTF cooling off will cause center frequency drift.



Figure 7-4 shows the consequence of YTF center frequency drift and repeatability.

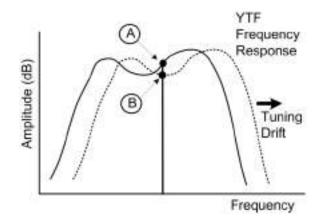


Figure 7-4. Frequency Response of the YTF

The frequency response over the bandwidth of the YTF is not flat. As the frequency changes due to either drift or repeatability, the measured amplitude will fluctuate.

The signal analyzer frequency response specifications reflect the extra amplitude uncertainty due to the YTF – there is usually a clear breakpoint in frequency response between lowband path and highband path.

Many signal analyzers have a path to bypass the YTF. **Figure 3-4** shows such a structure. Bypassing the YTF can lead to undesired spurious performance issues. However an RF substitution measurement can be used to reduce the YTF frequency response uncertainty. First use the YTF path to locate and measure the test signal's amplitude. With the YTF selected, the user can be assured that the signal displayed is the true signal and not an image. Next, select the YTF bypass path and measure the test signal's amplitude. When the YTF path is once again selected, the amplitude difference between the YTF and the YTF bypass path can be subtracted from subsequent measurements.

Near Noise Amplitude Error

Noise voltage is random in both amplitude and phase. Vector addition of the random noise voltage and a CW signal results in a combined amplitude that is still random. The method to combine amplitudes is through a power addition as shown in **Equation 7-1**.

$$P_{Signal+Noise} = 10\log(10^{P_{signal}/10} + 10^{P_{noise}/10})$$

Equation 7-1

As demonstrated in **Figure 6-35**, as the signal amplitude approaches the noise floor of the signal analyzer, the signal power and the noise powers add. With sufficient trace averaging, this offset can be treated as a deterministic error, not an uncertainty. One can measure the noise floor adjacent to the signal, or if possible, by removing the signal. Then measure the displayed signal plus noise amplitude. The true signal power then can be computed.

Figure 7-5 shows the terms used in the measurement of near noise signal amplitudes.



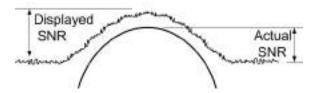


Figure 7-5. Near Noise Displayed SNR and Actual SNR.

Use **Equation 7-1** to compute the *actual SNR* from the measurement of the *displayed SNR* as shown in **Equation 7-2**.

Actual SNR = $10\log(10^{Displayed SNR/10} - 1)$

Equation 7-2

The chart in Figure 7-6 shows the curve generated by Equation 7-2.

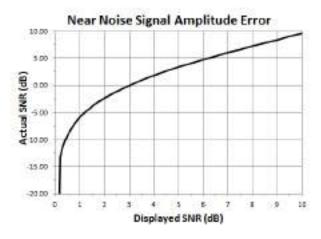


Figure 7-6. Signal + Noise versus Signal Level Relative to the Noise Floor.

Suppose the displayed SNR is 1.0 dB, then the actual SNR is -5.9 dB. This result implies that the signal is really almost 6 dB *below* the noise floor even though the apparent SNR is 1 dB. With knowledge of the noise floor power level and the actual SNR, the signal amplitude is readily apparent.

Of course, accounting for this error is time consuming. If the SNR is kept greater than 20 dB, then the error between the displayed SNR and the actual SNR is less than 0.05 dB.

Coherent Signal Addition

Coherent signal addition occurs when two or more non-random signals whose frequencies are the same combine. In the signal analyzer, this is most common in the measurement of distortion. The distortion component from the DUT can fall at the same frequency as the distortion from the signal analyzer. These two signals combine as voltages using vector addition. **Figure 7-7** shows that the resultant is the result of vector addition of the two signals.



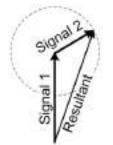


Figure 7-7. Coherent Addition of Two Signals

When the relative phases are the same, the two signals add constructively. At the other extreme, the signals with relative phases 180 degrees apart add destructively. The amplitude uncertainty due to two signals combining destructively is given by **Equation 7-3**.

Uncert = $20 \log(1 \pm 10^{-\Delta/20})$ [dB] where Δ is the relative amplitudes of the two signals in dB

Equation 7-3

Equation 7-3 is plotted in Figure 7-8.

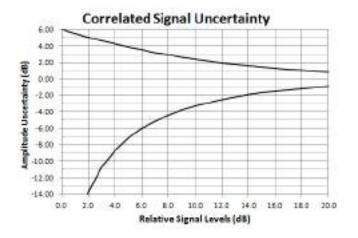


Figure 7-8. Amplitude Uncertainty versus Relative Signal Level for Coherent Signal Addition.

For example, if two signals at the same frequency have the same amplitude, the resulting combined signal could either range between 6 dB addition or completely cancel. For an uncertainty of less than ± 1 dB, the relative signal amplitudes need to be at least 19 dB apart.

Figure 7-9 shows that for the dynamic range chart, when considering amplitude uncertainty, the range of available mixer levels is limited.



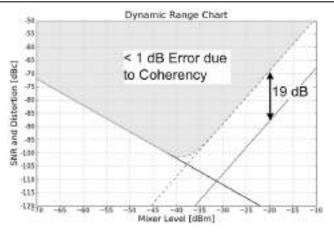


Figure 7-9. Dynamic Range Chart Show Range of Measurements for Uncertainty Due to Coherent Signal Addition is Less than 1 dB.

Mismatch Uncertainty

Signal analyzers can operate at high enough frequencies where only the average power is relevant. At very low frequencies, the phase of the voltage signal does not appreciably change as a function of cable length between the DUT and the signal analyzer. However, as the frequency increases, the assumption of constant phase does not hold true, in which case the amplitude of the voltage at the signal analyzer input does vary even for minute changes in cable length.

Average power, on the other hand, stays independent of cable length. Average power flow into the signal analyzer can be depicted as shown in **Figure 7-10**.

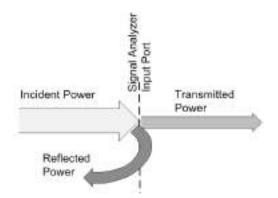


Figure 7-10. Power Flow into the Signal Analyzer.

Power flow is incident on the input of the signal analyzer. A portion of the signal is transmitted thorough to the internal path of the signal analyzer while the other portion gets reflected back to the DUT. Only the transmitted power is measured by the signal analyzer.

The term given to the ratio of the transmitted power to the incident power is *match loss*. Match loss is accounted for during the calibration process. However, what cannot be accounted for during the calibrations process is the re-reflected signal. **Figure 7-11** shows the flow graph of the re-reflected signal.



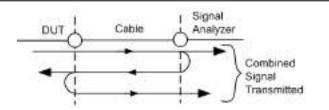


Figure 7-11. Reflected Signal Re-reflected Back to the Signal Analyzer.

The reflected signal exiting the signal analyzer's RF input port travels back to the DUT. At the DUT output, a portion of this signal re-reflects and travels once again to the signal analyzer. At the signal analyzer, the original transmitted signal and the re-reflected signals combine. This signal combination is a coherent addition of two signals. Unfortunately, the phases of the two signals are unknown, which creates an uncertainty. This uncertainty is termed *mismatch uncertainty*.

There are a few related terms used to qualify the input impedance of the signal analyzer. The majority of signal analyzers have a nominal impedance of 50 ohms. To calculate mismatch uncertainty, the *reflection coefficient* of both the signal analyzer input and DUT output must be known. *Voltage standing wave ratio* (VSWR) is normally the characteristic used to describe the signal analyzer's impedance. Most often *return loss* in dB is used to describe the DUT's output impedance. Use the following equations to arrive at reflection coefficients:

 $|\rho| = \frac{VSWR - 1}{VSWR + 1}$ where ρ is the reflection coefficient

Equation 7-4

$$|\rho| = 10^{Ret \ Loss/20}$$

Equation 7-4a

Return Loss = $20 \log(\rho)$

Equation 7-4b

Once the signal analyzer and DUT terms are converted to reflection coefficients, the bounds of the mismatch uncertainty are shown in **Equation 7-6**.

$$Uncert = 20 \log(1 \pm |\rho_{signal \, analyzer}||\rho_{DUT}|)$$

Equation 7-6

Figure 7-12 shows the relationship between VSRW and return loss.



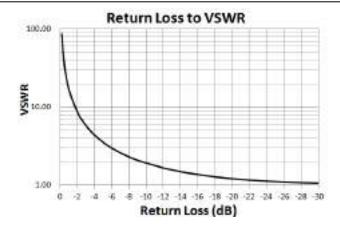


Figure 7-12. VSWR to Return Loss

As an example, suppose the signal analyzer has a VSWR specification of 1.5:1 and the DUT output return loss is -12 dB. The reflection coefficient for the signal analyzer is as follows:

$$p_{signal\ analyzer} = \frac{1.5 - 1}{1.5 + 1} = 0.2$$

The reflection coefficient for the DUT output is as follows:

$$\rho_{DUT} = 10^{-12/20} = 0.2512$$

The mismatch uncertainty is then the following:

$$Uncert = 20 \log(1 \pm (0.2)(0.2512)) = +0.43, -0.45 \text{ dB}$$

Figure 7-13 shows a graph of the mismatch uncertainty using signal analyzer and DUT return losses as parameters.

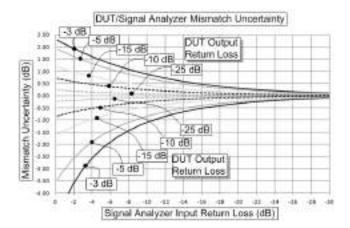


Figure 7-13. Mismatch Uncertainty. X-axis is Signal Analyzer Return Loss, Labeled Curves are the DUT Return Losses.

Most signal analyzer VSWR performance is a function of both frequency and RF input attenuator setting. In general, avoid using the lower value RF input attenuator settings for best VSWR performance.



Worst Case and Root Sum Squared Uncertainty

Based on the above descriptions, most uncertainty terms associated with amplitude measurements can be treated as statistically independent. When one uncertainty term does not depended on another uncertainty term, this independence is valid.

Worst case analysis, in which all the error terms are added, yields a very conservative overall measurement uncertainty. Empirical measurements using signal analyzers do not support the worst case error analysis.

A widely accepted method of combining uncertainty terms is the root-sum-of-squares (RSS) technique [¹⁷]. The RSS algorithm is described in **Equation 7-7**.

$$RSS = \sqrt{u_1^2 + u_2^2 + u_3^2 + \cdots}$$

where u_x are the individual uncertainty terms

Equation 7-7

The RSS technique gives a 95% confidence (2σ) that all measurements will have a combined uncertainty within the limits predicted by the equation.

Nearly all uncertainties in signal analyzer specification sheets have logarithmic units expressed in dB. One question is whether the uncertainty terms need to be converted to linear units (Watts) before applying the RSS equation. The log curve is nonlinear which may bias the mean of the individual error terms. **Figure 7-14** shows that for small value uncertainty terms, the slope of the linear power to log power curve is approximately linear. In these cases, the RSS equation, shown in **Equation** 7-7, can be used with terms expressed in dB.

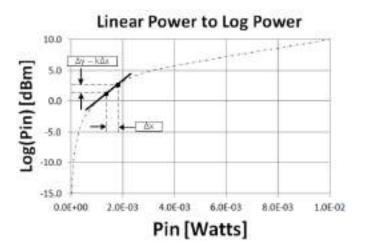


Figure 7-14. Small Value Amplitude Changes on the Log Scale are Approximately Linear.

To compute using linear terms, use the following equations:

Start with all uncertainties expressed in dB. Call these terms Δ_x . Then compute the linear term using the following equations:



$$u_x = 10^{\Delta_x/10} - 1$$

Equation 7-8

Compute linear RSS as follows:

$$RSS_{lin} = \sqrt{\sum u_x^2}$$

Equation 7-9

Then convert to log as follows:

$$RSS_{log} = 10 \log(1 + RSS_{lin})$$

Equation 7-10

Table 7-1 summarizes an experiment where six uncertainty terms, each of which has the same uncertainty value, are combined using worst case, RSS log and RSS linear.

| Individual Uncert. | Worst Case | RSS (log) | RSS (lin) | RSS Differenc e | Units |
|-----------------------|---------------|--------------|--------------|-----------------------|-------|
| 1 dB x 6 | 6 | 2.45 | 2.13 | 0.316 | dB |
| 0.5 dB x 6 | 3 | 1.22 | 1.14 | 0.089 | dB |
| 0.25 x 6 | 1.5 | 0.61 | 0.59 | 0.024 | dB |
| 0.1 x 6 | 0.6 | 0.24 | 0.24 | 0.004 | dB |

Table 7-1. Worst Case vs. RSS Uncertainty Analysis

RSS Difference column is the dB error between RSS log where the uncertainties in dB are computed using **Equation 7-7** directly and RSS linear where the individual uncertainty terms are first converted to linear using **Equation 7-8**, **Equation 7-9**, and **Equation 7-10**.

One can see how conservative the worst case uncertainty analysis is compared to RSS analysis. Use to determine whether RSS log or RSS linear is appropriate. If all uncertainty terms are less than 0.5 dB, then the difference between RSS linear and RSS log is less than 0.1 dB. RSS linear is the more accurate method.

Power Meter Leveling

From an amplitude uncertainty perspective, the definition of an ideal source is one whose source impedance is exactly equal to the characteristic impedance of the system (normally 50 ohms or 75 ohms for CATV applications). Why is perfect source output impedance so important? Recall from **Figure 7-11**, that mismatch error occurs due to the re-reflected portion of the incident signal combining with the signal transmitted to the signal analyzer. If the re-reflected signal disappears, so does the mismatch error. **Figure 7-15** shows the block diagram of the system with an ideal (from an RF sense) source.



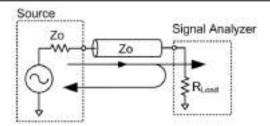


Figure 7-15. Ideal Source: Impedance Equals Characteristic Impedance of the System.

Comparing the signal flow of **Figure 7-15** with that of **Figure 7-11**, the re-reflected signal does not exist when the source match equals the system's characteristic impedance, Z_o . According to **Equation 7-6**, mismatch uncertainty drops to 0 dB if reflection coefficient, ρ_{DUT} , is zero. ρ_{DUT} in this example is the output impedance of the source. Reflection coefficient can be computed from **Equation 7-11**.

$$\rho = \frac{R - Z_0}{R + Z_o}$$

Equation 7-11

Let R be the output impedance of the source. When R equals characteristic impedance Z_o , the reflection coefficient is zero.

One method of creating an effectively perfect source match is to level the source as shown in **Figure 7-16**.

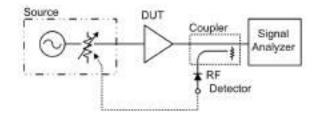


Figure 7-16. Source Leveling for Constant Impedance Source Impedance Using RF Detector.

A directional coupler siphons off a minute portion of the incident signal as it enters the RF input port of the signal analyzer. An envelope detector converts the coupled amplitude to a video signal that is then fed back to the source's amplitude control. The source creates a compensation amplitude component that, when combined with the signal analyzer's incident signal, results in a signal with flat frequency response as it enters the signal analyzer. **Figure 7-17** depicts the leveling technique.

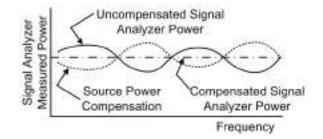


Figure 7-17. RF Detector Allows Source Power to Compensate for Signal Analyzer Mismatch.



A flat frequency response of the incident signal created using leveling effectively reproduces the ideal source that has an output impedance equal to the system characteristic impedance.

However, the leveling technique suffers from a few challenges. One is that the source may not have the ability to have its amplitude externally modulated. If the source is an over-the-air signal, this clearly cannot be amplitude controlled in this fashion. The coupler's frequency range can be limited, especially on the low end. The coupler introduces a few RF impairments, such as output match and isolation between coupler output port and the coupled arm. Finally, the envelope detector response is not perfectly flat with frequency and also not perfectly linear with power level.

The *power meter ratio* technique overcomes some of the deficiencies of the source leveling technique. This technique uses simultaneous measurements with the signal analyzer and a power meter. The power meter's inherently better amplitude accuracy can then be transferred to the signal analyzer measurement. The result, if proper care is given to the measurement technique, can yield effectively flat incident power into the signal analyzer. Flat incident power equates to no mismatch uncertainty term. However, if not properly implemented, the power meter ratio technique can introduce worse mismatch error than if nothing had been done.

Power splitters come in a few different topologies. Often times the labeling is not clear on the type of power splitter, so one must use the specifications as clues. The Wilkinson power splitter is lossless in the sense that no power (ideally) is dissipated within the power splitter. The input power is split evenly between the two output ports. Each port power is approximately 3 dB lower in power than the power at the input. The isolation between output ports can be quite large, exceeding 25 dB.

Figure 7-18 shows a setup with the Wilkinson power splitter driving a power meter with one output port and the signal analyzer with the other. The source is not limited to a true source; in general this is a DUT.

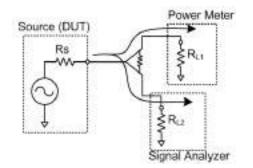


Figure 7-18. Wilkinson Power Splitter with Power Meter.

A perfectly reasonable assumption is that the DUT power evenly splits between the two output ports of the power splitter. The false assumption is that the power meter and the signal analyzer both measure the same power level. **Figure 7-19** shows measurement results with signal analyzer input return loss equal to -10 dB, source output match equal to -16 dB and power meter return loss equal to -25 dB. The power splitter is nearly ideal: 3 dB insertion loss, -30 dB return loss on all ports, and 25 dB isolation between output ports.



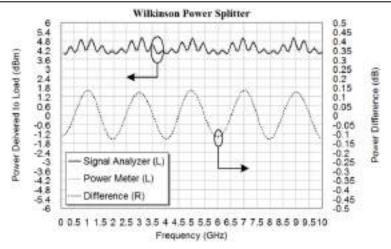


Figure 7-19. Power Meter Ratio Technique Using Wilkinson Power Splitter.

The *difference* signal represents the ratio of the power meter to signal analyzer. This difference signal is the amplitude error in the measurement that occurs when the power meter is used to determine the incident power into the signal analyzer. The uncertainty for this example is 0.3 dB (\pm 0.15 dB).

Wilkinson power splitters do not work at lower frequencies. They are often limited to operate over one or two octaves of frequency. To overcome the frequency range limitations, a resistive splitter is used. There are two flavors of resistive splitter: 3-resistor and 2-resistor. For both resistive splitters, power is dissipated within the splitter, which is the cost of broad frequency coverage.

Figure 7-20 shows the power meter ratio measurement using the 3-resistor power splitter.

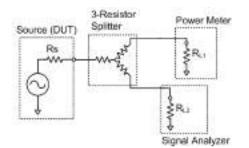


Figure 7-20. 3-Resistor Splitter.

The resistor values are Zo/3 or 16.67 Ohms each for a 50 Ohm characteristic impedance system. The insertion loss is roughly 6 dB for each port. The output port isolation is only 6 dB, which is a clear differentiation from the Wilkinson power splitter.

Figure 7-21 shows the results of using the power meter ratio technique with the 3-resistor power splitter. The source, signal analyzer, power meter load, and source matches are the same as used in the Wilkinson splitter example.



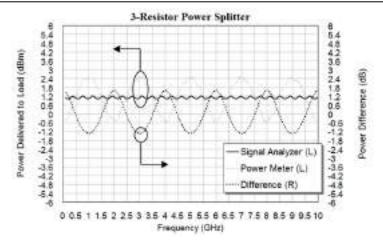


Figure 7-21. Power Meter Ratio Technique Using 3-Resistor Power Splitter.

Here the difference between the power meter and the signal analyzer is nearly 3 dB, clearly an abysmal result.

The 2-resistor power splitter has equal value resistors on the output arms, each with value equal to the characteristic impedance (50 Ohm most likely).

Figure 7-22 shows the connection for the power meter ratio technique using the 2-resistor power splitter.

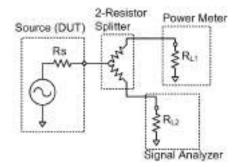


Figure 7-22. 2-Resistor Power Splitter.

Using the same load and source match values as in the previous examples, the measurement results for the 2-resistor power meter ratio technique using the 2-resistor power splitter are shown in **Figure 7-23**.



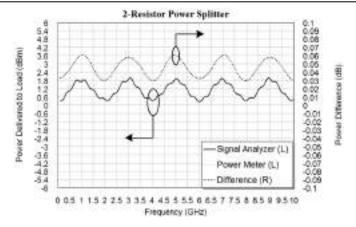


Figure 7-23. Power Meter Ratio Technique Using 2-Resistor Power Splitter.

Now the power meter and signal analyzer power measurements track to within \pm 0.015 dB. Reference [¹⁸] explains the theory behind both the 3-resistor and 2-resistor power meter ratio technique. Some of the magic with the 2-resistor power splitter when used in this fashion is that the reflected signal from the signal analyzer appears at the power meter port. The power meter then can automatically account for this reflected power. By using the ratio of the power meter and the signal analyzer powers (remember, ratio when using amplitudes expressed on a log scale is subtraction) the result is flat incident power to the signal analyzer.

The methods to reduce the signal analyzer's mismatch uncertainty for amplitude measurements are summarized as follows:

- Wilkinson power splitter is marginal. Use it with care.
- 3-resistor power splitter is unacceptable. Avoid using it.
- 2-resistor power splitter is the best solution

The Wilkinson and 3-resistor power splitters have their place in this world, just not for the use of trying to drive down signal analyzer mismatch uncertainty.



8. Spectrum Monitoring Applications

Spectrum monitoring entails over-the-air measurements for the detection and surveillance of RF signals. The signal analyzer seems a natural fit for this application where not only does a magnitude versus frequency spectrum become useful for determining whether or not a signal of interest exists, but once detected, demodulation may be required to decode the signal. Probability of intercept (POI) becomes the chief metric for signal monitoring applications. Dynamic range and measurement speed feed into the POI are of merit. The dynamic range problem for signal monitoring applications differs somewhat from that of standard R&D and manufacturing test applications. This section highlights those differences and offers some considerations to enhance the dynamic range of the signal analyzer for spectrum monitoring applications.

The Spectrum Monitoring Dynamic Range Problem

For R&D and manufacturing tests, the signal analyzer is most often connected to the DUT using a shielded coaxial cable. In many cases, the signal being measured is known in frequency. This describes a relatively benign environment.

Spectrum monitoring most often involves connecting the signal analyzer to an antenna in order to receive signals over-the-air. Near pandemonium can result with a large number of unknown signals impinging on the signal analyzer. To compound matters worse, the signal of interest is often a weak signal in the presence of large amplitude interfering signals. **Figure 8-1** alludes to this situation.

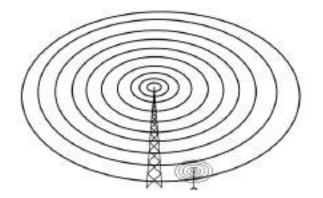


Figure 8-1. Weak Signal of Interest in Presence of Strong Interfering Signal.

In the frequency spectrum, the weak signal of interest/strong interfering signal appears as shown in Figure 8-2.

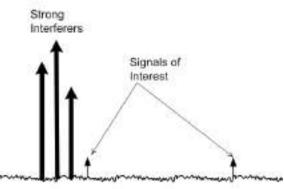




Figure 8-2. Frequency Spectrum of Low Amplitude Signals of Interest in the Presence of Large Amplitude Interferes.

Taking the spectrum depicted in **Figure 8-2** and applying it to the signal analyzer could result in the situation described in **Figure 8-3**.

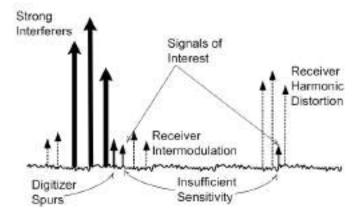


Figure 8-3. Distortion Created in the Signal Analyzer for the Spectrum Monitoring Situation.

Harmonics of the interfering signals can mask signals of interest. Intermodulation distortion created in the signal analyzer as a result of multiple interferes can mask signals of interest that fall near these IMD terms. Additionally, the signal analyzer's noise floor can bury the signal of interest in noise.

In section 6 the concept of the dynamic range chart helps to optimize the signal analyzer when measuring signals with sufficiently large amplitude. However, this is usually the wrong optimization for signal monitoring. Increasing the signal analyzer's RF input attenuation to drive down the signal analyzer's internally generated harmonic and intermodulation distortion has a severe impact on noise performance. Quite often, the SNR of the weak signal of interest cannot tolerate such a tradeoff in noise performance.

Spectrum Monitoring Receiver

As described in reference [¹⁹], a specialized receiver especially optimized for over-the-air spectrum monitoring may be warranted. **Figure 8-4** shows the additional blocks added to the signal analyzer that the spectrum monitoring receiver brings.

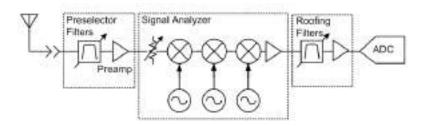


Figure 8-4. Spectrum Monitoring Receiver Block Diagram.

The primary block that the spectrum monitoring receiver adds to the signal analyzer is the preselector. The preselector can either be a tunable bandpass filter or a series switched filters placed at the RF input path. The primary function of these filters is to reduce harmonic distortion produced in the signal analyzer. And so, their bandwidths must be narrower than one octave (an octave is a doubling in frequency).



The location of the preamplifier is after the preselector filter. Recall from **Table 6-1**, a preamplifier added in front of the signal analyzer reduces dynamic range; distortion degrades faster than the improvement in noise performance. However, this is not necessarily true if the preamplifier is protected by the preselector filter. The insertion loss of the preselector filter does degrade noise figure performance as compared to the unprotected preamplifier, but the enhanced distortion performance far outweighs the noise figure degradation.

Figure 8-5 shows the harmonic distortion situation with and without preselection.

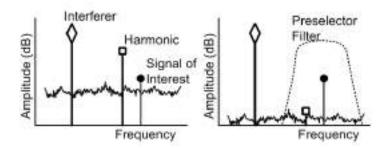


Figure 8-5. Harmonic Distortion with and without Preselection and Preamplification.

The left hand side of **Figure 8-5** shows the spectrum without preselection. The interferer generates harmonic distortion within the signal analyzer. If the harmonic distortion signal is near the signal of interest, deciphering the signal from the false distortion component becomes very difficult. POI drops appreciably if the signal analyzer's distortion gets confused for the true signal of interest.

The right hand spectrum shown in **Figure 8-5** is with the preselector included. First of all, the signal analyzer's harmonic distortion improves dramatically due to the filtering of the interfering signal. Next, with the large amplitude signals no longer impacting the signal analyzer, the preamplifier gain can be increased to improve the overall receiver's noise performance. This has the added benefit of discovering the weak signal of interest from the noise floor. Even for sufficiently large SNRs, by widening the resolution bandwidth filter setting, a faster scan of the spectrum can take place. With low system noise figure, a wider RBW setting is possible. Fast scanning also aids in higher POI.

One final component in the spectrum monitoring receiver as shown in **Figure 8-4** is the set of so called *roofing filters*. These are no more than a richer set of analog IF filters. For multiple interferers spaced too closely in frequency to be filtered adequately by the preselector filters, the roofing filters help protect the back end analog blocks and the digitizer from generating intermodulation distortion. These roofing filters behave according to the discussion related to **Figure 6-42**.

A wider range in bandwidth is offered with the roofing filters over the analog IF filters in the signal analyzer. This aids in a better tradeoff between selectivity and measurement speed.



9. Summary

Whether called spectrum analyzers or vector signal analyzers, signal analyzers that employ a superheterodyne architecture have been described in this application note. Much of the discussion centered on the internal architecture of this type of signal analyzer. The motivation behind this application note is that a fuller understanding of how the signal analyzer works under the hood allows the user to optimize the measurement for the unique needs of a particular measurement.



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Tips for Getting the Best Performance from Vector Signal Analyzers



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1. Overview

As wireless communication becomes widespread, the increased complexity of the devices and wireless protocols create an unrelenting demand for linear RF components and systems. These ubiquitous wireless devices require high performance test systems to characterize linearity. Linearity performance, particularly third order intercept point (IP3 or TOI), has become a defining characteristic as it affects power efficiency, channel density, signal coverage, and adjacent channel power ratio (ACPR).

IP3 performance is a figure of merit used to describe the linearity of components including power amplifiers, frequency mixers, switches, ADCs, DACs, and others. This measurement involves the use of a two-tone stimulus and requires the comparison of the power of two fundamental tones with the power of third order distortion products.

ACPR, also known as Adjacent Channel Leakage Ratio (ACLR), is an important metric for components designed for use in wireless standards such as UMTS or 3GPP LTE. This measurement describes the ratio of power in a modulated signal versus power emitted into an adjacent channel. In order to measure ACPR, one must provide the device under test (DUT) with a modulated stimulus – or in the case of a DAC – internally generate a modulated signal on the DUT itself.

2. Theory of Intermodulation Distortion

Intermodulation distortion describes the linearity of a device when presented with a multi-tone stimulus. An ideal component, such as a power amplifier, (PA) should have low noise figure and remain linear - so it does not distort the incoming signals. A weakly non-linear system has a voltage-in to voltage-out relationship that can be described by the power series shown in (1). When operating in the linear range of the DUT, the output voltage closely follows this power series.

$$v_{out} = a_0 + a_1 v_{in} + a_2 v_{in}^2 + a_3 v_{in}^3 + a_4 v_{in}^4 + \cdots$$
(1)

Consider the behavior of the two-tone third-order intermodulation. If two continuous wave (CW) signals are placed at the input of the DUT, the voltage is represented by (2). The two tones have different radial frequencies denoted as ω_1 and ω_2 .

$$v_{in} = A\cos(\omega_1 t) + B\cos(\omega_2 t) \tag{2}$$

By replacing the expression for the input voltage into (1), the relationship between the input and output signal is more apparent. From the power series, one can easily see that a CW signal at the input generates output distortion signals at $2\omega_1$, $3\omega_1$, and so on. Each harmonic has amplitudes proportional to v_{in}^2 , v_{in}^3 , and so forth. The output voltage is represented by (3).

$$v_{out} = a_0 + a_1\{\dots\} + a_2\{\dots\} + a_3\{A^2Bcos[(2\omega_1 - \omega_2)t] + AB^2cos[(2\omega_2 - \omega_1)t] + \dots\}$$
(3)



The DUT generates many tones due to intermodulation (one tone mixing with the other tone) as shown in **Figure 1**. The intermodulation tones of most concern are the third-order tones at $2\omega_1 - \omega_2$ and $2\omega_2 - \omega_1$ as shown in **Figure 2**.

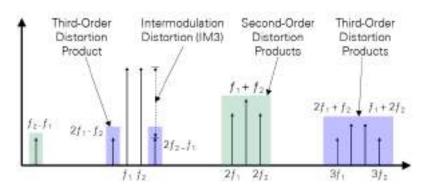


Figure 1: Intermodulation Tones Caused by Non-linear DUT

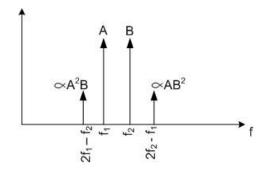


Figure 2: Two-Tone Third-Order Intermodulation Distortion

These distortion products are spaced too close in frequency to be effectively filtered. When considering IP3 performance, the fundamental tones powers are set to the same power level. From **Figure 2**, one can derive that a 1 dB change in the fundamental tone results in a 3 dB change in the third-order distortion products. When viewed on a log-power out versus log-power in graph, the fundamental tones powers increase with a slope of 1. The third-order distortion products increase with a slope of 3 as shown in **Figure 3**.



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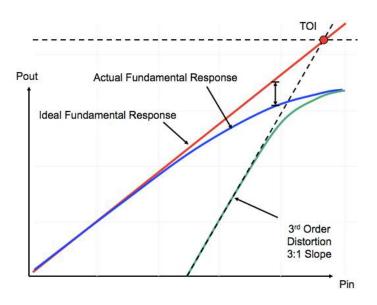


Figure 3: Power out versus Power in Plot for Fundamental versus Third-Order Intermodulation Distortion

At a theoretical power level, the fundamental and distortion level lines intersect. This intersection is referred to as the third-order intercept point (IP3 or TOI). However, the intersection is fictitious – it is realized only in mathematics and used as a figure of merit in industry. In practical use, a device goes deeply into gain compression well before the input power approaches the IP3 level. Moreover, the simple power series model that was expressed in (1) breaks down at higher power levels. In practice, the empirical formula to calculate IP3 is shown in (4), where P_{IMD3} is the third-order intermodulation distortion power.

$$IP3_{in} = \frac{3}{2}P_{in}(dBm) - \frac{1}{2}P_{IMD3}(dBm)$$
(4)

IP3 can be expressed as $IP3_{in}$ or $IP3_{out}$. For example, mixers use $IP3_{in}$ and amplifiers use $IP3_{out}$. The decision to specify either the input or output IP3 is typically driven by marketing and industry standards. The relationship between $IP3_{out}$ and $IP3_{in}$ is dictated by the gain (dB) of the DUT. For example, convention dictates that the linearity of a high-gain device is specified by its output IP3 and that the linearity of a passive device such as a mixer is specified by its input IP3.

3. Optimizing IP3 Measurements

Although IP3 measurements require complex instrumentation configurations, one can ensure the accuracy of these measurements using a few simple best practices. When measuring IP3, one must ensure that the distortion products of both the sources and the signal analyzer are significantly smaller than the distortion products introduced by the DUT. The lower the inherent IP3 introduced from the actual measurement setup, the more accurate the IP3 number. As a result, one must take care to ensure a low distortion signal source and optimize VSA settings such as reference



level, resolution bandwidth, and attenuation. As shown in **Figure 4**, the IP3 measurement requires two RF signal generators and a power combiner to produce a high-quality two-tone stimulus.

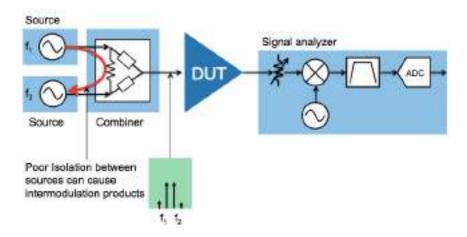


Figure 4: IP3 Measurement Setup Including Unwanted Transmission from Lack of Source Isolation

Source isolation is sometimes critical to making an accurate IP3 measurement. When using a power combiner, one must provide sufficient source isolation because without it, energy from one source can leak into the other source as shown in **Figure 5**. This leakage tone creates AM sidebands at the difference frequency, which falls at the same frequency as the measurement distortion products.

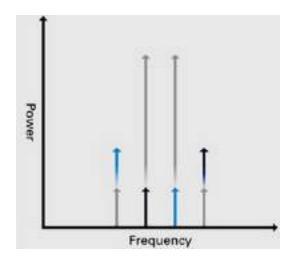


Figure 5: Distortion Caused by a Lack of Source Isolation

One way to improve the isolation is to choose a combiner with the best port-to-port isolation. Generally, purely resistive combiners feature only between 6 dB and 12 dB of isolation, depending on the resistor topology. Additional measures to increase the isolation using either directional couplers or isolators can provide up to 50 dB of isolation if used at both ports. However, couplers are often limited to a single octave and are thus not suitable for broadband applications.



One solution to the isolation challenge is to place an attenuator at the output of each signal generator. The attenuator provides additional isolation for signals traveling in the reverse direction. A good rule of thumb is that roughly 40 dB of isolation is required to measure up to IP3 numbers above +25 dBm. Another option is to place an amplifier at the output of the sources. The amplifier not only provides the extra gain needed to produce high power signals, but also offers inherently good isolation. It is important to note that the amplifier must exhibit high linearity so excess distortion isn't introduced into the measurement.

Optimizing VSA Linearity

In addition to ensuring isolation between sources in the stimulus signal, one must also configure the VSA to use optimal linearity settings. As shown in Figure 6, the signal-to-noise ratio (SNR) and internally generated distortion products are related to the mixer level of the VSA. Typically, the mixer level is optimized for the best dynamic range automatically by adjusting the front-end attenuation/gain when the reference level is set. However, to manually adjust the mixer level, one can set the attenuation to achieve the desired response.

The ideal setting (point C on Figure 6) for the VSA occurs at the point where the mixer level is high enough to reduce the noise floor but also low enough that the internally generated distortion points do not interfere with the measurement. Point A on Figure 6 shows a setting where the measurement is limited by the VSA noise floor. Essentially, more gain is required, which will reduce the noise floor enough to resolve the intermodulation tones. Point B on Figure 6 demonstrates the case where the downconverter has too much gain and more attenuation is required to reduce the internally generated VSA distortion.

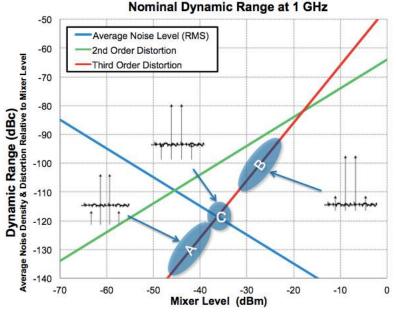


Figure 6: Dynamic Range Chart of the PXIe-5668R VSA

An easy way to determine if the VSA is contributing to the IP3 measurement is to increase the front-end RF attenuation and to observe whether the distortion decreases. If the distortion stays constant, then one can be certain that the distortion products are coming from the DUT. On the other hand, if the observed distortion decreases, one



can infer that the signal analyzer distortion decreased because the signal level at the mixer is reduced. This technique, known as "rolling the step attenuator," identifies the source of distortion. While it is important to set the reference level and attenuation appropriately, one can also reduce the effective noise floor by reducing the resolution bandwidth (RBW). **Figure 7** shows a TOI measurement being performed by the RFSA Soft Front Panel.

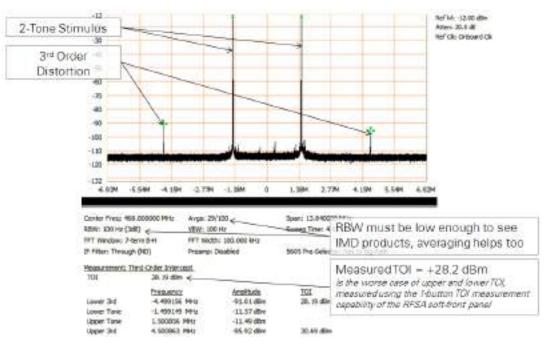


Figure 7: Anatomy of a TOI Measurement using the RFSA Soft Front Panel

Equation 5 shows how the noise floor relates to the noise figure and the RBW of the VSA. However, by decreasing the RBW, the measurement time increases. VSAs with inherently low noise figures help to minimize the measurement accuracy to speed tradeoff.

$$Noise \ Floor = -174 \frac{dBm}{Hz} + NF_{VSA} + 10 * log_{10}(RBW)$$
(5)

VSA IF Conditioning

Another consideration when setting the VSA is the IF signal conditioning. Typically, one configures the gain of the analyzer to use an IF power that is slightly less than full scale of the analog-to-digital converter (ADC) to prevent clipping. One can improve the noise floor of the instrument by setting a narrow IF bandwidth and increasing the IF power level.

To accomplish this, one must space the distortion tones so they exceed the bandwidth of the VSA's IF filter. By filtering out the two-tone stimulus, one also reduces the distortion products internally generated by the ADC, which allows for a more accurate IP3 measurement. **Figure 8** depicts a dynamic range chart including the ADC distortion versus VSA mixer level and **Figure 9** shows a graphical representation of this optimization technique.



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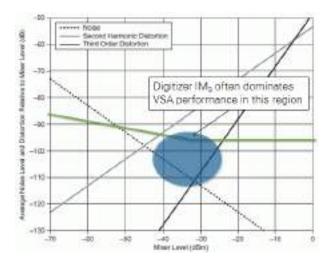


Figure 8: Dynamic Range Chart Including ADC Distortion

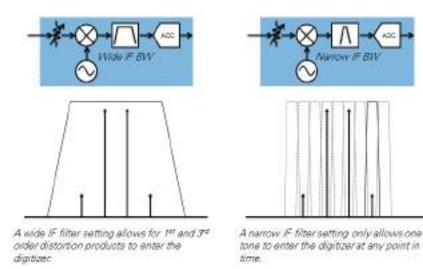


Figure 9: Graphical Representation of IF Filter Settings

Another valuable technique to improve the IP3 measurement is to apply IF gain re-ranging. The previous dynamic range charts in this paper describe a VSA setting with a constant reference level across the entire measurement span. With a narrow IF filter, one can apply more gain (lower reference level) while measuring the IMD3 tones to significantly improve the noise floor and decrease the inherent distortion. Since the two fundamental tones aren't present at the ADC, gain can be applied without overloading the backend. **Figure 10** shows this re-ranging technique where the appropriate reference level is still applied to the fundamental tones.



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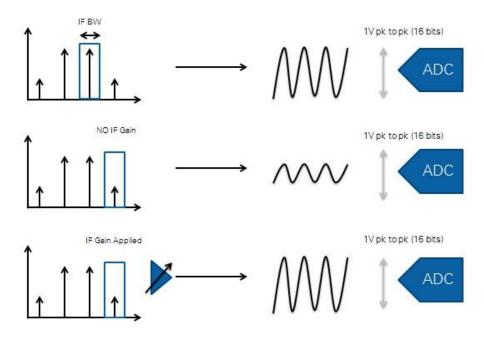


Figure 10: IF Gain Re-ranging to Improve the Noise Floor and Decrease the Inherent Distortion

IP3 is an important figure of merit used by industry to describe the linearity of a DUT, and it presents many challenges. While measuring IP3, careful consideration of the measurement setup including sufficient source isolation and optimized VSA settings is critical for accurate results. The ability to optimize an IP3 measurement is an appropriate transition to describing ACPR theory.

4. Theory of Adjacent Channel Power Ratio

The adjacent channel power ratio of a wireless communication signal describes the ratio between the integrated power in the carrier channel relative to the adjacent channel. ACPR is the wideband equivalent to the CW IP3 case. The adjacent channel power ratio of a wireless communication system is the integrated power in the carrier channel relative to the adjacent channel. ACPR is the wideband equivalent to the CW IP3 case. Figure 11 depicts this specification graphically. When a modulated wideband signal is incident on a non-linear DUT, spectral re-growth will occur. This re-growth has the unintended consequence of leaking into the adjacent channel, which can cause over-the-air interference within a different carrier's channel bandwidth.



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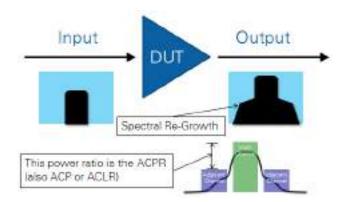


Figure 11: Graphical Depiction of ACPR

One technique to model spectral re-growth of a weakly non-linear system, such as the VSA when operating far below its gain compression limits, is to divide up the digitally modulated signal into equal frequency slices. Represent each slice as a CW whose average power is the average power of that slice. As described above, the tones will create intermodulation tones at predictable frequencies and amplitude levels.

Using the CW IP3 behavior, one can model the spectral re-growth by adding the amplitudes as voltages (assuming the phases of the individual contributors line up). To increase the model accuracy, one can continue to decrease the frequency slice bandwidths, which increases the number of CW tones. Essentially, by taking the limit as the BW of each slice approaches zero, the model accuracy improves. Primarily, the intermodulation is third-order, but fifth and even seventh-order products can affect ACPR. **Figure 12** shows the modeled spectral re-growth as a discrete spectrum of CW tones.

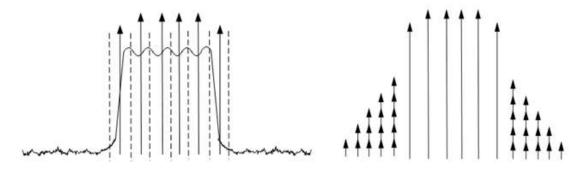


Figure 12: Spectral Re-growth Modeled as Intermodulation Tones

Through this exercise, it's apparent that the linearity of a device or system directly impacts the ACPR performance. As shown in **Figure 13**, other factors include broadband noise and phase noise.



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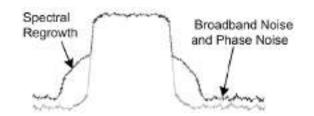
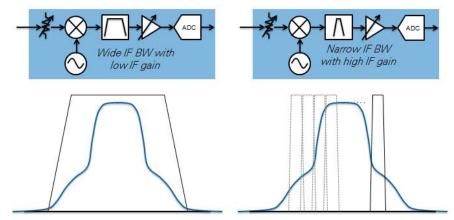


Figure 13: Broadband Noise and Phase Noise Contribution to ACPR

Although factors such as phase noise theoretically can limit the ACPR specifications of measurement devices such as a signal analyzer, modern high-performance signal generators and analyzers typically exhibit phase noise performance that well exceeds what is required for measuring ACPR on modern communications signals.

5. Optimizing ACPR Measurements

Similar to the IP3 measurement case, IF filtering plays a critical role in improving the ACPR measurement performance. Although the source is a digitally modulated wideband signal, by decreasing the IF filter and increasing the IF gain, one can achieve better ACPR numbers. The RFSA software will collect data in multi-span FFT acquisitions and stitch the spectrums together. Like before, the tradeoff for narrowing the IF is measurement speed. There is a proportionate increase in acquisitions that must happen to cover the wideband signal. **Figure 14** demonstrates this technique graphically.



A narrow IF filter not only constrains signal bandwidth at digitizer – but also reduces IF signal power. As a result, additional IF gain can be applied to improve ACPR.

Figure 14: Graphical Depiction of ACPR Measurement with Narrow IF Filtering

Another technique that can be applied to increase the ACPR measurement performance of the VSA is noise correction. When configured to use noise correction, the VSA first measures its internal noise floor, and caches this value. It then subtracts this cached "noise power" from resulting spectrum measurements to more accurately reflect the noise power of the actual signal. As a result of noise correction, one can reduce the inherent ACPR of the VSA



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by more than 7 dB. Note that additional measurement time is required to perform the computation necessary to measure and subtract the VSA's inherent noise floor. **Figure 15** shows both the inherent VSA noise and actual noise at the VSA input.

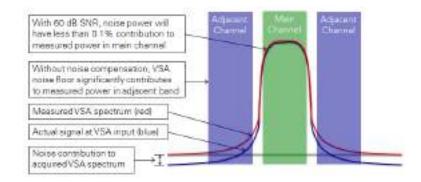


Figure 15: Graphical Depiction of Noise Correction

The modulation signal is noise-like so the power spectral density of both the signal and the noise will scale by the same amount as a function of RBW setting. For this reason, SNR is independent of RBW setting. By decreasing the RBW, the main channel power will decrease proportionally with the adjacent channel noise. Therefore, even with the optimization techniques mentioned above for IP3 and ACPR, the ACPR measurement capability of a VSA distills down to the dynamic range of the receiver. Best-in-class VSAs like the PXIe-5668R are capable of measuring ACPRs above 80 dB.

Figure 16 shows the PXIe-5668R with the RFSA Soft Front Panel measuring the ACPR of a WCDMA signal without narrow filtering and without noise correction. The main channel is centered at 468 MHz with a channel bandwidth of 3.84 MHz.



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Figure 16: PXIe-5668R Measuring WCDMA ACPR with 100 MHz Wide IF Filter and Noise Correction Disabled

By decreasing the reference level, the VSA's ACPR measurement performance increases if the signal measurement is noise limited. However, at a certain point distortion created by the VSA results in spectral re-growth as shown in **Figure 17**.



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Figure 17: PXIe-5668R Measuring WCDMA ACPR showing ADC Distortion

Figure 18 shows the PXIe-5668R with the RFSA Soft Front Panel measuring the ACPR of a WCDMA signal with the 300 kHz narrow IF filter and without noise correction. By reducing the bandwidth of the analog IF filter, one can now increase the IF gain by lowering the reference level to improve the ACPR capability.



Tips for Getting the Best Performance from Vector Signal Analyzers



Figure 18: PXIe-5668R Measuring WCDMA ACPR with 300 kHz Wide IF Filter and Noise Correction Disabled

Figure 19 shows the PXIe-5668R with the RFSA Soft Front Panel measuring the ACPR of a WCDMA signal with the 300 kHz narrow IF filter and with noise correction. Since the receiver is still noise limited, one can now enable noise correction to reveal the source's noise. It is important to note that IF gain re-ranging can also be applied for these measurements. The IF gain can be increased within the adjacent channel bandwidth to improve the VSA dynamic range. Also, to increase measurement speed, a larger bandwidth IF filter can be used in the adjacent channel without performance degradation as long as the narrow IF filter is used within the main channel. The RFSA Soft Front Panel does not perform this re-ranging, so additional ACPR performance can be achieved with the appropriate implementation.



Tips for Getting the Best Performance from Vector Signal Analyzers

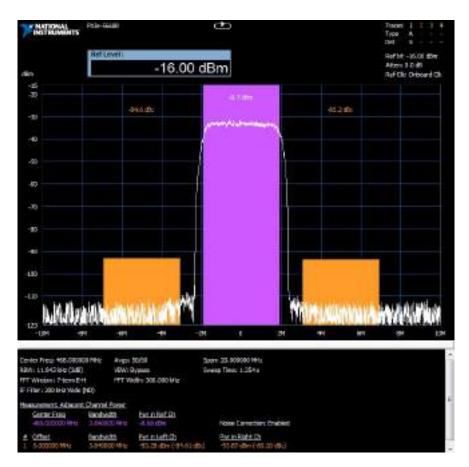


Figure 19: PXIe-5668R Measuring WCDMA ACPR with 300 kHz Wide IF Filter and Noise Correction Enabled

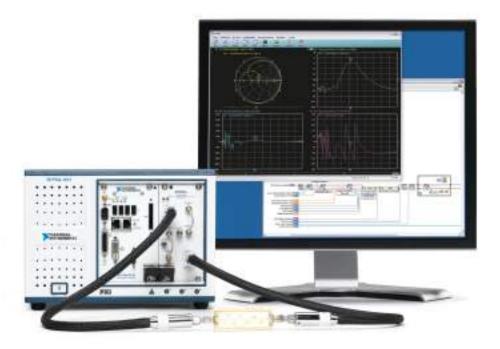
In accordance with theory, the measurement ability increased with a narrow IF filter and with noise correction enabled. The table below summarizes the ACPR WCDMA measurement capabilities of the PXIe-5668R using the RFSA Soft Front Panel at 468 MHz.

| IF Filter BW | Noise Correction | ACPR (dBc) |
|--------------|------------------|------------|
| 100 MHz | Disabled | -75.0 |
| 300 kHz | Disabled | -77.4 |
| 300 kHz | Enabled | -84.6 |

As devices continue to advance, it will become more difficult to measure linearity performance and even more important to invest in best-in-class instruments to measure these sensitive but important specifications. The PXIe-5668R High-Performance Vector Signal Analyzer offers best-in-class performance up to 26.5 GHz for linearity measurements.



Introduction to Network Analyzer Measurements Fundamentals and Background





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1. Introduction to Network Analyzer Measurements

Before the network analyzer, determining the reflection coefficient of a circuit required you to manually calculate the phase of the reflection coefficient one frequency at a time on a piece of paper. This manual process required you to first find the maxima and minima of the standing voltage wave in a slotted line and repeat this process for all frequencies of interest. It is safe to say that if the vector network analyzer (VNA) had not been created, the high-frequency community would look very different today.

Today, the VNA characterizes high-frequency passive and active devices in their linear mode of operation by measuring their network parameters, called S-parameters, as a function of frequency. Over time, VNAs have been extended in hardware and in capability to also measure noise parameters and nonlinear characteristics, including compression, intermodulation, and hot S_{22} measurements¹. As a result, the VNA evolved into an instrument enabling multiple measurements with a single connection.

VNA Basics

In its simplest form, a network analyzer is an instrument used to measure impedance. At lower frequencies, you can measure impedance with relatively simple tools, including a sine wave generator, a volt meter, a current meter, and a calculator. Using these tools, you can measure the ratio between voltage and current and calculate the resulting impedance, as shown in **Figure 1.1**.



¹ Hot S22 measurements are S22 device measurements where drive power is provided to the device input while the device is excited at port 2, typically with an offset frequency.

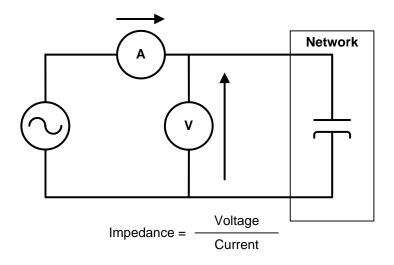


Figure 1.1. Determining Impedance

At RF and microwave frequencies, measurements of voltage and current become more complex. As a result, a VNA uses a more complex design to measure incident and reflected waves. In many ways, the VNA marries the principles of the basic impedance measurement with hardware appropriate for microwave frequencies.

When using a VNA to measure the impedance or the reflection factor, a sine generator stimulates the device under test (DUT). In addition, two receivers take the place of the combination of a volt meter and current meter. These receivers, with the help of signal separation hardware, characterize the response of the device by measuring the phase and amplitude of signals that are both incident to and reflected from the DUT. Finally, calibration capabilities are required to eliminate systematic errors and compute the appropriate ratios (similar to the impedance) necessary to produce S_{11} , which is one part of the S-parameters. The architecture of this simplified VNA is illustrated in **Figure 1.2.** To measure devices with more ports the architecture needs to be extended.



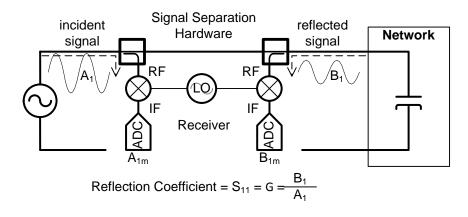


Figure 1.2. Measuring Impedance at Higher Frequencies

Modern network analyzers cover a wide range of frequencies from the low hertz (Hz) to tens of terahertz (THz).

Applications for Network Analyzers

The primary use of a VNA is to determine the S-parameters of a myriad of passive components, including cables, filters, switches, diplexers, duplexers, triplexers, couplers, bridges, transformers, power splitters, combiners, circulators, isolators, attenuators, and many more. In addition, VNAs can also characterize active devices such as transistors and amplifiers using S-parameters, as long as they are operating in their linear mode of operation. High-frequency devices can have one or two or more ports.

RF and Microwave

S-parameters provide complete insight into the linear behavior of RF and microwave components and are independent of the characteristics of the VNA itself. Through the VNA calibration, impairments of the instrument are completely removed from the measurement. Thus, S-parameters are an extremely accurate representation of the linear behavior of the component under test, describing how it behaves and how it interacts with other devices when cascaded. For this reason, you can use S-parameters in a wide range of applications, ranging from a simple metric to characterize circuit performance to a basic behavior model in electronic design automation (EDA) tools. In the latter case, S-parameters enable the simulation of complete RF circuit chains.

S-parameters are very useful because they are an integral part of different theories, such as filter, transmission line, and amplifier design theories. Also, S-parameters are essential to performing proper noise parameter measurements. While engineers are increasingly facing nonlinear problems, they mainly design for linear modes of operation, as the arsenal of engineering tools is still based on linear theories.

Other Uses

Over time, VNAs have even found their way into applications besides the pure S-parameter measurement capability. Modern VNAs are being used for the characterization of the dielectric properties of materials (both liquids and



solids) and even the human brain. Because the VNA can measure the reflection factor as function of frequency, the results it produces can be transformed into the time domain. As a result, one common application of modern VNAs is cable fault detection, which takes advantage of time-domain reflectometry (TDR).

Note that in addition to becoming the tool of choice for RF and microwave circuits, the VNA has also become widely used in signal integrity applications. In fact, the relatively high dynamic range of a VNA, in comparison to the traditional oscilloscope, makes it a useful diagnostic tool.

Additionally, electromagnetic imaging is one of the growing fields where VNA technology is fast becoming an essential tool. Similar to the detection of a cable fault, it is possible to reconstruct images from reflected waves by performing a form of scanning. VNAs have penetrated into medicine to help visualize, for example, breast tumors. Also in the security realm, microwaves are used for body scans, and some applications are taking advantage of a VNA as a basic instrument.

The growing ranges of applications for which a VNA can be used have driven their design into new form factors beyond the traditional box instrument. For example, because of the need to conduct VNA measurements in the field, VNAs are also available in portable formats. In addition, because applications such as electromagnetic imaging require significant software, the VNA is increasingly connected to the PC. In fact, in the case of VNAs in the PXI form factor, the network analyzer becomes a part of the PC itself.

History of Network Analysis

In the early 1950s, Rohde & Schwarz achieved a breakthrough in network analysis with a product called the Zg diagraph [1]. Until then, network analysis had been limited to manual measurements of standing waves in a slotted line. The Zg covered a frequency range from 30 MHz to 300 MHz, though the frequency range was later extended to 2.4 GHz. This instrument became the first complex network analyzer on the market that could directly measure the phase of S-parameters and display the S-parameters on a Smith Chart. The Zg was used to measure radios, TV antennas, and some telecommunication cables.

Between the 1950s and early 1960s, three other companies introduced a form of network analyzer based on the homodyne principle. In 1965, Wiltron (later acquired by Anritsu) introduced the Wiltron 310 Impedance Meter. During that time, other companies such as Hazeltine Corporation and Rantec Corporation [2] created some of the world's first network analyzers.

Improvements in the VNA Design

In 1966, Hewlett Packard introduced the 8405A vector voltmeter, which used step recovery diode (SRD) technology. Using SRD technology, you could measure both amplitude and phase difference between two signals up to 1 GHz. The SRD enabled sampling techniques to down convert high-frequency signals into manageable low frequencies. Note that the SRD technology in the HP 8405A was leveraged from the sampling oscilloscope product HP 185/7 that was brought to the market around the same time [3].

By combining the 8405A with couplers or bridges, you could measure the amplitude and phase difference of incident, transmitted, or reflected waves of a DUT at one frequency. Next, you would use a printed-paper Smith



chart to plot an impedance chart as function of frequencies. It is important to note that at that time, sophisticated calibration techniques were not available to remove the vast majority of measurement uncertainties. Instead, you could only compensate for the difference in length between the various measurement paths.

In 1967, Hewlett Packard introduced the 8410A network analyzer, which contained the ability to directly plot a Smith chart on the instrument's screen. One year later, HP introduced the world's first automatic network analyzer, the HP8540A. This instrument was a 2-cabinet system that included the HP 8410 network analyzer and a HP 2001A "computer". The computer contained a maximum of 32 K words of memory, a 5 MB disk drive, a digital cassette tape, and a paper tape reader or puncher. The first version had a teletype as human interface, though a later version, the HP 8542A, had a "real" CRT display and improved software [3]. In many ways, the HP 8542A revolutionized RF and microwave component and system design. For the first time, it automated the measurement process by allowing users to perform multi-band frequency sweeps by measuring and plotting impedance and transfer characteristics on a Smith Chart.

Meanwhile, scattering parameters, or S-parameters, were quickly becoming the standard format to describe passive and active devices in their linear mode of operations at high frequencies. At first, the network analyzer could only measure the transmission and reflection wave. Thus, you were forced to reverse the DUT to measure these quantities at the opposite device port. Soon, more sophisticated test sets became available that could measure the full Sparameter set of a two-port device. Thanks to the automation of the measurements, it became possible to develop and deploy more sophisticated calibration techniques to eliminate the systematic errors of the measurement system.

In contrast to VNAs, scalar network analyzers measure standing wave ratios, reflection and transmission coefficients only as scalar quantities rather than as vector quantities. The first realizations of a scalar network analyzer date from 1954 and were developed in parallel with vector network analyzers. At the time, scalar network analysis was more adequate for production test than it was for design work. As a result, scalar network analyzers were sold in much larger quantities than VNAs.

Automated VNAs

While automated vector network analyzers were available since the late 1960s, it was not until 1985 that a fully automated, relatively compact vector network analyzer first became available. The first of this category was the 8510, which was introduced by Hewlett Packard [4]. The HP 8510 network analyzer took major advantage of the microprocessor technology that had become available. Using microprocessor technology, the instrument could perform frequency sweeps and labor-intensive calibration techniques relatively quickly. For the first time, a vector network analyzer was capable of giving almost real-time feedback while tuning a DUT, representing the measurements in both the time domain and the frequency domain. Soon after the 8510 introduction, Wiltron introduced the 360 Network Analyzer [5], which benefited from similar technological advancements.

In the 1980s, engineers conducted much research on the use a set of scalar detectors or power measurements to derive the amplitude and phase information of the incident and reflected waves. One example of this is a six-port reflectometer with three scalar detectors at each port to determine the coefficients-parameters of the DUT. Using only power detectors, the setup was highly simplified compared to a VNA. However, this approach never took off commercially.



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The Modern Era

In the 1990s, growth in high-frequency systems such a cellular telephones and cable television created increasing demand for instrumentation that was capable of characterizing RF and microwave circuits. As the demand for VNAs increased, engineers began to enjoy a broader range of price points and performance capabilities for VNAs. As a result, the market saw various low-end VNAs being commercialized and deployed into test and manufacturing environments while high-end VNAs continued to target research and development. Meanwhile, the scalar network analyzer had been replaced by the VNA in many applications due to VNA benefits such as a reduced cost and technical advantages.

One key innovation in network analysis technology in the early 2000s was the evolution of the VNA towards measuring devices in their non-linear range of operation. As RF and microwave components have been pushed to their limits, an increasing number of engineers are required to characterize the nonlinear behavior of the devices under test. The need for measurements beyond S-parameters, which only describe linear behavior, has imposed a new requirement on the VNAs to measure additional device characteristics and requires the presence of the combined sources and receivers. As a result, a wide range of vendors such as Agilent Technologies, Rohde & Schwarz, and Anritsu began to develop non-linear network analysis technology.

Given the challenges of modern network analysis, the complexity of network analysis technology has elevated the role of software in VNA systems. As a result, PC-based form factors are an increasingly popular choice for VNAs. In 2011, National Instruments introduced the first PXI-based vector network analyzer, the NI PXIe-5630, which was followed one year later by the NI PXIe-5632 with full S-parameter measurement capability. A NI PXIe-5632 VNA is shown in a PXI system in **Figure 1.3**.





Figure 1.3. PXI VNA System



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2. Understanding Wave Quantity Measurements

S-parameters describe the complete linear behavior of a high frequency device or circuit in linear mode of operation. The main function of a vector network analyzer is to provide S-parameters as a function of the frequency. Although modern VNAs return S-parameters with a push of the button (after proper calibration), an S-parameter is not a quantity that can be measured directly. Instead, the S-parameters are the result of solving a set of equations after measuring ratios of wave quantities while exciting the device at its various ports.

Measuring Wave Quantities

To properly understand S-parameters, you must first understand the theory and process of measuring wave quantities. Consider, for example, a load connected to a sine wave generator through a piece of cable, as illustrated in **Figure 2.1**. The sine wave generator has a certain source impedance, and the electrical power travels across the wire from the source to the load as the combination of an electric and magnetic field. The propagation speed (v_p) depends on the dielectric properties of the medium and is a fraction of the speed of light (*c*). The wavelength, λ , is defined as a function of the propagation speed $(v_p \text{ or } c)$ and the sine wave generator frequency (f_0) in **Equation 2.1**.

$$\lambda=\frac{v_p}{f_0},$$

Equation 2.1. Wavelength as a function of frequency and propagation speed

When the frequency (f_0) is low, the wavelength is large, and the length of the cable is negligible compared to the size of the wavelength. As a result, the measured voltage and current are independent of the location on the cable. This situation is illustrated in **Figure 2.1**, and the circuit is referred to as being a lumped element circuit or lumped circuit.

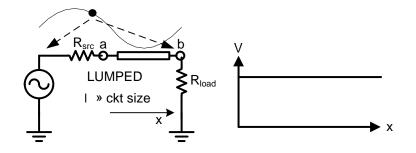


Figure 2.1. A Lumped Circuit

When the frequency (f_0) of the source increases, however, the wavelength is reduced. Thus, as frequency increases, the wavelength eventually becomes similar in size or even smaller than the length of the cable.

In a scenario where the wavelength of the signal is similar or smaller in size to the length of the cable, the measured voltage and current will depend on the position, as shown in **Figure 2.2**. Thus, measuring the voltage with a voltage



probe is invalid because the result will be dependent on the probe's position. In this scenario, the circuit must be treated as a distributed element circuit rather than as a lumped circuit.

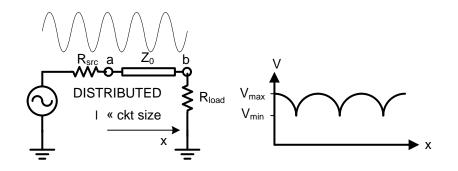


Figure 2.2. A Distributed Circuit

Analysis of a distributed circuit is more complex and involves the use of transmission line theory. Using transmission line theory, you can think of the electrical power traveling along the line starting from the source. While traveling on the line, a certain voltage (E-field) and current (H-field) relation is imposed by the electrical properties of the line. In fact, the cable itself will behave such that it is characterized by an inherent impedance that does not change as long as the properties of the line or cable do not change. This impedance is called the characteristic impedance (Z_0).

As the electrical power hits the termination (R_{load}), the voltage to current relationship is now imposed by the impedance of the load. Under the condition where the load impedance is equal to the characteristic impedance, the power is fully absorbed. To keep the reasoning simple, real impedances are assumed here. If the load impedance is different from the characteristic impedance, the ratio of voltage and current will change at the point where the transmission medium occurs. As a result, the load will not absorb all the power, resulting in a portion of the power traveling back towards the source. Given this characteristic, you can refer to the system as having both an incident and reflected wave.

The ability to separate incident and reflected waves is a critical element of network analysis not only because it identifies how much power is absorbed by the load but because it reveals information about its impedance. As a result, one of the most important functions of a VNA is to separate these incident and reflected waves by means of either couplers or bridges in order to measure each of these waves independently. Thus, while measurements from simple volt and current meters are able to determine impedances and transfer functions of circuits at lower frequencies, measuring similar characteristics at higher frequencies requires the measurement of incident and reflected waves. In fact, the VNA is able to measure the amplitude and phase differences between individual waves, using one of the waves as a reference. Modern VNAs are even able to measure these quantities in an absolute way instead of in a relative way. **Figure 1.2** clearly shows the signal separation hardware, which separates incident and reflected waves using directional couplers.



Introduction to S-parameters

Determination of scattering parameters, or S-parameters, is the basic measurement capability of vector network analyzers. The S-parameters describe the magnitude and phase relationship between incident and reflected waves and are numbered according to where a wave originates from and where it propagating to. The term *scattering parameters* comes from the "Scattering Matrix" described in a 1965 IEEE article entitled, "Power Waves and the Scattering Matrix." Today, S-parameter results, such as S_{21} for example, are displayed on the VNA display screen with the simple push of the button. Moreover, the instrument is able to apply appropriate calibration techniques to remove a wide range of inherent impairments of the instrument itself.

In order to better understand the S-parameters in an intuitive way, it is useful to first consider an example from the field of optics—the behavior of light through varying transmission mediums. Suppose a light beam is shining on a lens. In this example, some light will pass through the lens, undergoing the influence of the lens. However, some light will also be reflected back by the lens towards the direction of the light source, as illustrated in **Figure 2.3**.

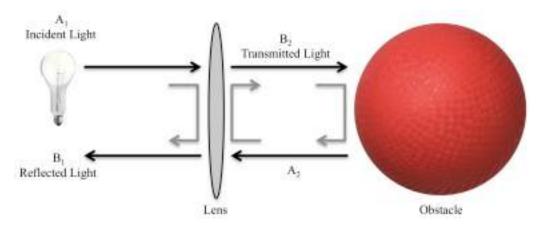


Figure 2.3. Optical Equivalent of S-parameter Behavior

In addition, it is also useful to investigate what happens once the light passes through the lens. If there is no obstacle behind the lens, the light will continue to travel forever. If an obstacle exists, however, part of the transmitted beam is reflected back towards the lens. This light will be partially reflected back towards the obstacle and will be partially transmitted through the lens. Thus, the reflection from the obstacle is added to the initial reflected by the lens and a part of the reflection from the obstacle. Similarly, the transmitted wave consists of the original beam transmitted through the lens and partially the reflected beam from the obstacle, also shown in **Figure 2.3**.

The lens and light analogy translates easily to the understanding of the way waves travel through an electrical circuit. As illustrated in **Figure 2.3**, both the B_1 and the B_2 wave can be described as a combination of both the A_1 wave and the A_2 wave. These relationships can be formally represented as S-parameter equations for a two-port device as shown in the following equations.



 $b_1 = S_{11}a_1 + S_{12}a_2$ $b_2 = S_{21}a_1 + S_{22}a_2$

Equation 2.2. 2-port S-parameters Equations

Figure 2.4 shows a two-port device under test (DUT) along with the corresponding four S-parameters and the four waves.

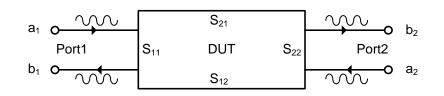


Figure 2.4. Two-port S-parameter Model

It is important to realize from **Equation 2.2**, that a wave leaving the DUT $(b_1 \text{ or } b_2)$ is a linear combination of the waves entering the DUT $(a_1 \text{ or } a_2)$. As a result, terminating the DUT in the characteristic impedance (Z_0) allows the extraction of the individual S-parameters. For example, when the characteristic impedance (Z_0) is equal to 50 Ohms, and if a 50 Ohm termination is present at port 2, a_2 is reduced to zero, resulting in equations for S₁₁ and S₂₁. This principle can be applied in the reverse direction as well. By setting a_1 to zero, you can obtain equations for S₂₂ and S₁₂, as shown in **Equation 2.3**.

$$S_{11} = \frac{b_1}{a_1}\Big|_{a_2=0} \qquad S_{12} = \frac{b_1}{a_2}\Big|_{a_1=0}$$
$$S_{21} = \frac{b_2}{a_1}\Big|_{a_2=0} \qquad S_{22} = \frac{b_2}{a_2}\Big|_{a_1=0}$$

Equation 2.3. Formal Definitions of S-parameters

Note that a common notation is to use the subscripts *m* and *n* for the general S-parameter (S_{mn}), where *m* is the receiver port and *n* is the source port. For example, S_{12} is the transmission coefficient for a wave sourced at port 2 and received at port 1. Keep in mind that these ratios are complex-valued, and the S-parameter will carry both a magnitude and phase component as a function of frequency.

Finally, the equations in **Equation 2.2** can be converted into matrix notation and generalized for any number (n) of ports as illustrated in **Equation 2.4**.



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$b = S \cdot a$ $S = \begin{bmatrix} S_{11} & \dots & S_{1n} \\ \vdots & \ddots & \vdots \\ S_{n1} & \dots & S_{nn} \end{bmatrix}$

Equation 2.4. S-parameter Equations in Matrix Notation

In theory, determining the S-parameters using a VNA is a straightforward process. In practice, however, it is impossible to terminate the port opposite to the driving port exactly in the characteristic impedance (Z_0), resulting in the presence of a reflected wave at all times. Given this challenge, a solution is to first connect the source to port 1 and terminate port 2, which allows the VNA to measure ratios b_1/a_1 , b_2/a_1 and a_2/a_1 as a function of frequency. Next, for a two-port device, the VNA switches the source to the output and terminates the input port of the DUT. With this configuration, the VNA can now measure ratios b_1/a_2 , b_2/a_2 and a_1/a_2 . Using basic linear algebra, it is possible to calculate the S-parameters as a function of frequency.

In practice, a VNA will generally measure the incident and reflected waves through a series of couplers or bridges, referred to as directional devices. The directional device is able to separate the incident from reflected waves. Using this measurement practice, the resulting measurements will be quantities that are subject to imperfections of the VNA, such as the coupling factor and directivity of the directional devices. Thus, more complicated calculations are needed in order to determine calibrated S-parameters starting from a set of raw quantities.



3. Network Analyzer Architectures

Historically, scalar network analyzers were a common tool used to characterize a network using only the magnitude of the signal. However, as network analysis technology matured, the development of digital components, such as analog-to-digital converters (ADCs), greatly simplified the design of vector network analyzers. As a result, most modern VNAs are capable of measuring both scalar (magnitude) and vector (magnitude and phase) information about a signal.

Scalar Network Analyzers

Scalar network analyzers typically capture a broadband signal and convert it to DC or low frequency AC in order to measure the power of the signal. Examples of the hardware used to accomplish this include diodes and thermoelectric devices. The main advantage of scalar network analyzers is that the hardware required for downconversion and power detection is relatively simple and inexpensive. In addition, because the detectors are broadband devices, it is unnecessary to re-tune the receiver to measure power at a different frequency. Thus, performing a frequency sweep is as simple as re-tuning the frequency of the microwave source and measuring the power at each frequency step.

Due to their relatively simple architecture, scalar network analyzers are capable of relatively fast frequency sweeps. Note that some scalar network analyzers simplify the hardware even further by removing the reference sensor. In these designs, however, a slightly more complicated measurement sequence is required. A simplified block diagram of a scalar network analyzer is shown in **Figure 3.1**.



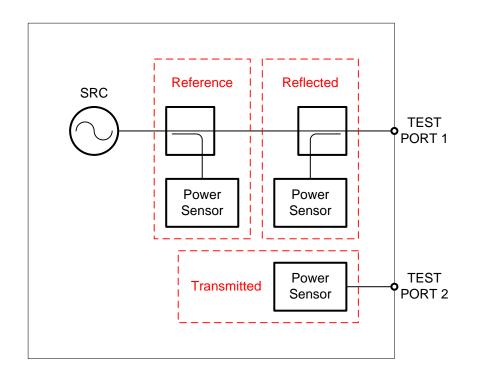


Figure 3.1. Simplified Scalar Network Analyzer Block Diagram

Scalar analyzers make transmission or insertion loss measurements by using a signal source that sweeps repetitively over the frequency range of interest. If the reference detector (as seen **Figure 3.1**) is not available, you can calculate the transmission coefficient as the power ratio of the transmitted signal with and without the device under test (DUT) in the signal path. In this case, two sweeps are required to characterize the device. If a reference detector is present, you can calculate the transmission coefficient as the ratio of the transmitted power to the incident power as measured by the reference detector. For reflection measurements, a directional device, either coupler (as seen in **Figure 3.1**) or bridge, is necessary to detect the signal reflected by the DUT. A ratio of the reflected signal power over the incident signal power at the device is the reflection coefficient.

While scalar network analyzers have the benefit of simplicity of design, they are also prone to inherent challenges. For example, broadband detectors are susceptible to spurious tones and broadband noise. In addition, because the calibration is scalar in nature, it is not as accurate as full vector calibration. Due to their lack of selectivity, scalar network analyzers tend to have limited dynamic range compared to vector network analyzers.

Vector Network Analyzers

Vector network analyzers generally use full heterodyne receivers to measure both the phase and magnitude of signals and are often significantly more complex than scalar network analyzers. Measurements made with vector network analyzers are often more accurate, and the narrowband nature of the receivers provides better rejection of



broadband noise and spurious tones, allowing for improved dynamic range. Furthermore, calibration can use more complex error models, which provide greater accuracy.

Due to the complexity of the heterodyne receiver architectures of vector network analyzers, these instruments generally perform frequency sweeps more slowly than broadband scalar network analyzers. In addition, the added complexity often makes them more expensive.

The fundamental principle of a vector network analyzer is to measure the amplitude and phase of both incident and reflected waves at the various ports of the DUT. The general design of a VNA is to stimulate an RF network at a given port with a stepped or swept continuous wave (CW) signal and to measure the travelling waves, not only at the stimulus port but at all the ports of the network terminated with specific load impedances, typically 50 Ohms or 75 Ohms. A typical but simplified VNA architecture is illustrated in **Figure 3.2**.

As observed in **Figure 3.2**, the block diagram of a VNA can be decomposed into different elementary blocks. First, note that one (or more) signal source with controllable frequency is used to drive the DUT. In addition, the test set includes two test ports. Each test port contains some signal separation hardware to split out the incident and reflected travelling waves. Optionally, the test set can contain switches to route the signal source to the different test ports and terminates other ports with specific load impedances. Finally, one or more receivers (marked ref 1, ref 2, test 1, and test 2 in **Figure 3.2**) capture both the phase and magnitude of the RF signals before processing them.



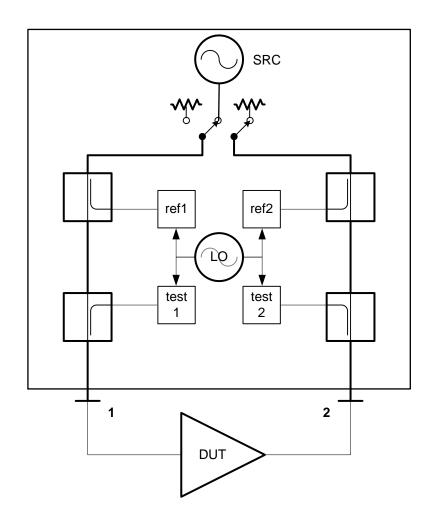


Figure 3.2. Typical VNA Block Diagram

VNA Source

The source of a VNA provides the stimulus that is used to characterize the response of a DUT. The source must be tunable for output frequency and power, and its performance is dictated largely by the types of measurements the VNA will make. The output signal is typically a sine tone for single-frequency S-parameter measurements. However, some advanced applications require multi-tone or modulated signals to provide the most complete characterization of a device's behavior.

In general, the most significant requirement of a VNA signal source is the speed and accuracy of its tuning capabilities. For most measurements, the source will rapidly sweep across frequencies to collect the frequency response of a device. However, for measurements such as gain compression, the source will sweep across output power at a fixed CW frequency.



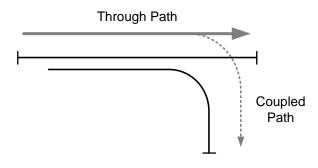
When measuring the linear behavior of the device, the requirement for spectral purity of the output signal is more relaxed than general-purpose signal generators. Because the frequency of the source is known, the VNA's receivers can selectively tune to the appropriate frequency. Thus, spurious tones and other undesirable frequency components can often be present without affecting the accuracy of a measurement. One caveat is that phase noise of the source signal contributes to the noise measured at the output. As a result, phase noise cannot be neglected entirely.

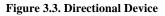
The spectral content of the source's output becomes much more critical when making nonlinear measurements. Multi-tone measurements, such as intermodulation distortion, are negatively impacted by undesirable signals, as are frequency-translation measurements.

Signal-separation Hardware

The signal-separation hardware is a key element of vector network analyzers. Each test port of the VNA makes use of signal-separation hardware, such as a directional coupler or bridge, to separate incident from reflected waves travelling the test port. Each wave is then captured with a dedicated receiver, allowing the instrument to measure the phase and magnitude of both waves independently.

As illustrated in **Figure 3.3**, a generic directional device separates the incident wave from the reflected wave, while a small portion of the incident wave travelling along the through path is diverted into the coupled path. Note that forward coupling is shown in **Figure 3.3**.





Key characteristics of a directional device are its insertion loss, coupling factor, isolation, and directivity. Insertion loss is a metric of the loss of signal power when it is traveling through the directional device (through path). A high quality directional coupler can present insertion loss as low as 1 dB up to microwave frequencies. You can calculate insertion loss using **Equation 3.1**.

Insertion Loss $(dB) = -10 \log_{10} \left(\frac{output \ power}{input \ power} \right)$



With any coupling device, only a portion of the incident signal going through the directional device is diverted to the coupled path. The coupling factor is a metric that describes the ratio of the coupled power to the total forward signal



power. In typical VNAs, the coupling factor varies anywhere from 10 dB to 20 dB, and the coupling factor is calculated using **Equation 3.2**.

Coupling Factor
$$(dB) = -10 \log_{10} \left(\frac{forward \ coupled \ power}{forward \ power} \right)$$

Equation 3.2. Calculating Coupling Factor

Ideally, a radio frequency signal travelling in the reverse direction of a directional device should not appear in the coupled path. In practice, however, some power traveling in the reverse direction is diverted in the coupled path due to the finite isolation between each port of the directional device. Isolation is one of the most critical characteristics of a coupling device, and performance in the range of 30 dB to 40 dB is required for accurate network analysis measurements. Isolation is expressed in **Equation 3.3**.

 $Isolation(dB) = -10 \log_{10} \left(\frac{reverse \ coupled \ power}{reverse \ power} \right)$

Equation 3.3. Calculation of Isolation

Directivity is a metric that describes the overall performance of the ability of a directional device to separate out forward and reverse signals traveling through the directional device. This metric combines elements of isolation, coupling factor, and insertion loss and is expressed in **Equation 3.4**.

Directivity (dB) = Isolation (dB) - Coupling Factor (dB) - Insertion Loss (dB)

Equation 3.4. Directivity As a Function of Isolation, Coupling Factor, and Insertion Loss

The directivity of a directional coupler has a significant impact on the measurement accuracy of a network analyzer. A large value for directivity implies a high performance VNA. In essence, directivity accounts for the amount of measurement error that is introduced by a wave traveling in the reverse direction of the coupling device.

Challenges with Poor Directivity

In scenarios such as a return loss measurement, the directivity error can cause artifacts in the measurement result. For example, consider what happens when the directional device diverting the travelling wave reflected back from the DUT has poor directivity. When both the reflected wave and the directivity error vector (introduced by the incident wave) are in-phase, the electromagnetic waves combine at the coupled port level, resulting in a measured return loss that is worse than the actual DUT performance. Alternatively, the directivity error vector can cancel out the captured reflected wave, resulting in return loss measurement that is actually better than the DUT performance. In addition, the phase difference between the coupled waves and directivity error vector is frequency dependent. Thus, poor directivity manifests itself with typical ripple patterns in many measurements of DUT response. **Figure 3.4a** and **Figure 3.4b** illustrate the impact of directivity on a measured response both where the directivity error is combined and where the directivity error cancels out.

The finite directivity of a VNA also has a direct impact on the precision of the instrument. Though the raw directivity can be compensated for with calibration (see **Vector Network Analyzer Calibration**), the finite directivity of the couplers will still impact the stability of the calibration. In fact, as the estimated directivity error



vector is subtracted from the raw DUT measurement during correction, a small drift in the actual directivity error vector may significantly influence the final result.

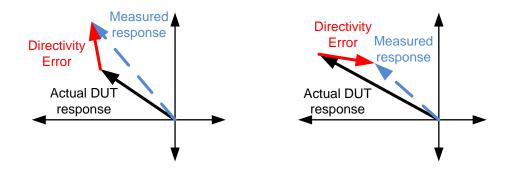


Figure 3.4a and Figure 3.4b. Impact of Directivity on Measured Response:

(a) Directivity Error Adds Up, (b) Directivity Error Cancels Out

The concept of directivity has implications to connectivity of the DUT. Because a signal separation device is required to separate a relatively low-power reflected wave from a relatively high-power incident wave, the directivity of the coupling device is one of the largest sources of measurement uncertainty. This uncertainty is further exacerbated when significant loss is added between the output of the directional device and the DUT through the use of poor quality test cables or attenuators. Because loss between the coupler and the DUT will further attenuate the reflected wave, the raw directivity of the directional device will be degraded by twice the attenuation value.

Note that both directional couplers and the directional bridges are common signal separation devices used in vector network analyzers. The choice between one or the other is frequently determined by the frequency range of the instrument. By design, directional bridges have higher insertion loss than directional couplers, but the former can operate at low frequency (down to kilohertz) while maintaining compact physical dimensions. At higher frequencies, directional couplers tend to have better performance than directional bridges, both in terms of insertion loss and directivity, and are generally preferred in microwave VNAs.

Receivers

Measurements of the phase and magnitude of a traveling wave in a VNA are determined by a receiver, which translates analog signals into the digital domain. **Figure 3.5** illustrates a generic block diagram of a single-stage downconversion heterodyne receiver.



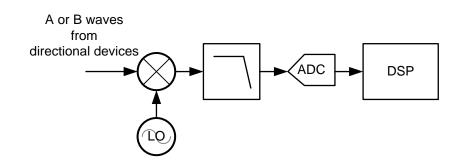


Figure 3.5. Generic VNA Receiver Block Diagram

Digitization and processing of the captured incident and reflected waves first require downconversion of the RF signal to an intermediate frequency (IF). As illustrated **Figure 3.5**, a local oscillator (LO) is used to drive the receiver's downconverter.

Although the receiver would achieve the best phase noise and spur performance using only the fundamental frequency of an LO, this design is often impractical. In fact, a broadband LO source able to cover the entire bandwidth of a microwave VNA is often challenging and costly in practice. As a result, it is more common to use harmonic-conversion process to limit the required frequency and power range of the LO source.

Sampling Receivers

The downconversion process from RF to IF can occur using either high-frequency mixers or sampling downconverters, or samplers. Historically, VNA receivers were mainly based on sampling downconverters due to simplicity and cost. In fact, such a downconverter requires only a relatively low frequency LO drive. The LO frequency limitation is then compensated by using a highly nonlinear circuit, such as shockline or step recovery diode, to generate a very short pulse with very high harmonic content. The sampling diode, acting as a switch, is then driven by this low frequency periodic pulse. As a result, the LO is effectively converted into a multi-tone stimulus, as shown in the following figure, which allows for simultaneous downconversion of multiple frequency bands at once.



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Introduction to Network Analyzer Measurements

Fundamentals and Background

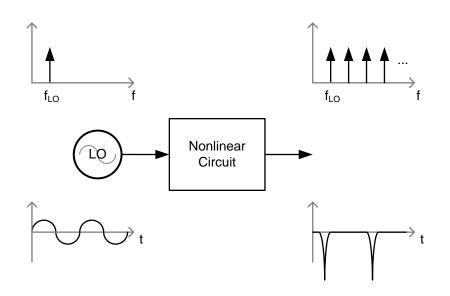


Figure 3.6. LO Source Converted into a Multi-tone Stimulus

Using sampler technology, the resulting downconversion process is then able to cover RF frequencies much higher than the LO drive frequency. Also, as the order of nonlinearity introduced by the pulse generator is very high, the harmonically related tones can have almost flat power over the full downconverter bandwidth. This results in about the same conversion efficiency for each LO harmonic. Unfortunately, the process of simultaneously downconverting multiple bands negatively impacts the dynamic range of the receiver. In this case, broadband noise and potential spurious signals are also downconverted into the same IF bandwidth.

Modern VNA Receiver Design

Today, almost all VNAs use mixers to perform the RF to IF downconversion since broadband LO sources have become less expensive. While several high-frequency mixer topologies exist, the most commonly used design is the doubly balanced diode-ring mixer [6]. These mixers use ring diodes, driven by the LO source, to perform the RF to IF downconversion.

Because broadband LO sources that cover the full bandwidth of microwave VNAs are quite expensive, the LO signal is typically clipped to look more closely like a square wave. Clipping results in generation of odd harmonically related tones. The captured RF signal is then mixed with the appropriate harmonic to produce the desired IF signal. The unused harmonics still contribute to the IF output, but in contrast with samplers, the power of the generated LO harmonics decays rapidly with the harmonic order *N* (proportional to 1/*N* for an ideal square wave). Because downconversion at higher harmonic order is less efficient, the IF bandwidth is less impacted by unwanted broadband noise and spurious signals. As a result, harmonic mixers typically present better dynamic range than sampling downconverters.

The downconverted IF signal is further converted and detected in the IF processing chain. Modern network analyzers use a high-speed analog-to-digital converter (ADC) to perform direct sampling of the IF signal. Also, an



anti-aliasing filter is placed just before the ADC input for proper IF image rejection. Note that signal conditioning hardware, generally an amplifier, can be used to optimize the signal-to-noise ratio in the ADC.

Finally, the digitized IF signal is processed using various digital signal processing (DSP) techniques to improve the ADCs as well as digitally downconvert, decimate, and filter the IF signal to retrieve the magnitude and phase (or real and imaginary part) of the measured signal. The results are then sent to the main computer processing unit (CPU) to apply error correction and calculate the DUT S-parameters or other derived quantities before displaying them.

Common VNA Architectures

Today, modern network analyzers rely on a combination of receivers and coupling devices to perform S-parameter measurements. In general, the theory of operation for both VNA architectures are similar, though each design will have a varying degree of complexity. The following sections discuss two VNA architectures: the transmission/reflection (T/R) test set and the full S-parameter test set.

T/R Test Set

One of the simplest network analyzer test sets is the transmission/reflection (T/R) test set, which uses one source fixed to port one to measure forward S-parameters, as shown in **Figure 3.7**. Because the source is fixed, the reverse S-parameters may not be measured without reversing the orientation of the DUT using either external switching or physical disconnection and reconnection, both of which can be difficult to calibrate for. However, the reduced complexity of the T/R test set makes it a less expensive option than a full two-port test set when measuring reciprocal devices or when only the forward measurements are necessary.



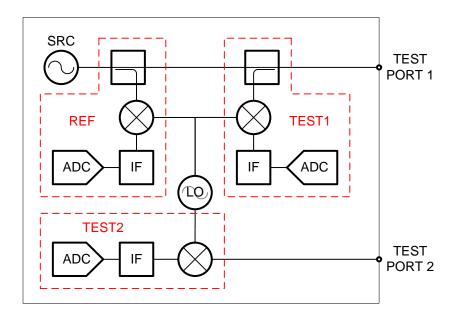


Figure 3.7. Transmission/Reflection Network Analyzer Block Diagram

As shown in **Figure 3.7**, the T/R test set contains three receivers for measuring the components of the forward Sparameters. The reference receiver (REF) measures the incident power from the source, a_{1F} . The test receivers (TEST1 and TEST2) measure the forward reflection and transmission parameters, b_{1F} and b_{2F} , respectively. Using these measurements, S_{11} and S_{21} are calculated.

In the T/R VNA architecture, the test receiver on port 2 is usually directly connected to the test port, in contrast to the coupled receivers on port 1. As a result, T/R test sets can often have higher dynamic range of S_{21} measurements, because the receiver is not subject to the coupling factor in the port 2 receiver path. However, this design may also lead to problems with load matching on port 2. Because a load mismatch cannot be fully corrected during user calibration, the resulting measurements will be affected by an uncorrected return loss at the reference plane of test port 2.

Full S-parameter Test Set

In contrast to the T/R test set, full S-parameter VNA architectures introduces an RF stimulus at each test port of the VNA. In its simplest version, shown in **Figure 3.8**, a full two-port VNA contains three receivers, an RF source and a transfer switch to route the RF stimulus to either test port 1 or 2. With the switch in the port-one switch position, the VNA measures forward reflection and transmission parameters a_{1F} , b_{1F} , and b_{2F} , and in the port-two switch position, it measures the reverse reflection and transmission parameters a_{2R} , b_{2R} , and b_{1R} without the need to physically disconnect and reverse the DUT.



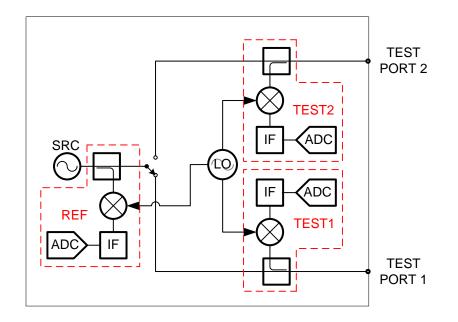


Figure 3.8. Simple VNA Using One Source and Three Receivers

As observed in **Figure 3.8**, the port 1 reference receiver (REF1) measures the forward and reverse reference parameters a_{1F} and a_{2F} . The port 2 test receiver (TEST2) measures the forward transmission parameter, b_{2F} , and the reverse reflection parameter, b_{2R} . The port 1 test receiver (TEST1) measures the forward reflection parameter, b_{1F} , and the reverse transmission parameter, b_{1R} .

In the simplest VNA architecture, the transfer switch is part of the network analyzer measurement path. As a result, VNA calibration accounts for the switch variation. However, slight differences in the two switch positions can lead to measurement uncertainty. In addition, switch contacts may wear over time, requiring more frequent VNA calibration. To resolve this, in alternative VNA architectures, the switch has been moved directly at the source output. Here the forward and reverse reference parameters are measured using two reference receivers, REF1 and REF2 respectively, as shown in **Figure 3.9**.



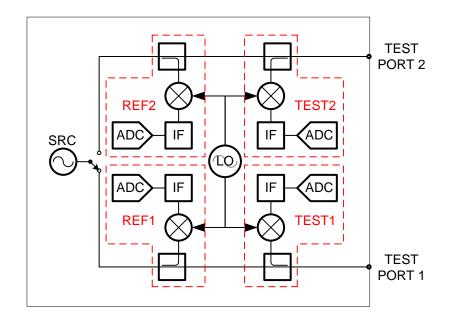


Figure 3.9. VNA Using One Source and Four Receivers

An alternative solution is to eliminate the switch entirely and simply use one RF source per port. This approach, as illustrated in **Figure 3.10**, provides the best performance but at the highest cost. Note that this architecture allows the VNA to perform advanced measurements beyond S-parameters, such as hot S22 measurements.



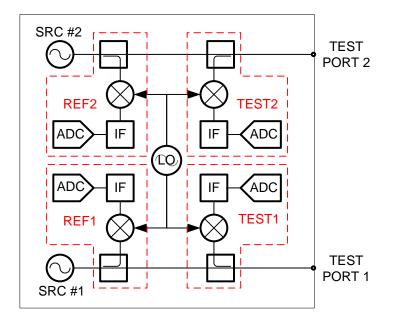


Figure 3.10. VNA Using Two Sources and Four Receivers

As previously mentioned, higher-performance VNA architectures often come with additional cost and complexity. Therefore, consider the application, performance, required accuracy, and cost among other factors, in order to select the proper VNA architecture.



4. Vector Network Analyzer Calibration

In order to understand VNA calibration, first consider the purpose of a vector network analyzer (VNA). The VNA is designed to *accurately* characterize the linear behavior of a device by evaluating the phase and magnitude of incident and reflected waves. By measuring the phase and magnitude of these waves, you can determine a wide range of device characteristics including impedance, return loss, insertion loss, and even group delay. Thus, the accuracy with which a VNA can determine the linear behavior of a device under test (DUT) is fundamentally determined by the accuracy with which the instrument can measure the phase and magnitude relationship of the incident or reflected waves.

Why Calibrate?

Although it is possible to calibrate certain elements of the VNA at manufacturing time, such as source power and receiver accuracy, much of the calibration information required to perform an accurate measurement depends on the measurement setup. For example, consider the example of measuring the impedance of an open circuit. In this example, illustrated in **Figure 4.1**, the open circuit is connected to the instrument through a cable.

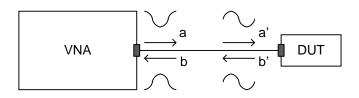


Figure 4.1. Effect of Test Port Cable on Impedance Measurement

It is well known that the open circuit has an infinite impedance and that the reflected wave will have the same magnitude as the incident wave and will be in-phase with the incident wave. However, as illustrated in **Figure 4-1**, the length of the cable between the VNA port and the open circuit will affect the phase as measured by the VNA. At certain frequencies, the open will even show up as a short when using only the factory calibration performed at the test ports.

In order to prove this phenomenon mathematically, assume the test port cable is lossless and introduces a delay (τ) that corresponds to the length of the cable. As a result, the incident wave (a'), at the end of the test port cable, will be a delayed version of the incident wave (a), at the test port, as shown in **Equation 4.1**.



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 $\begin{cases} a' = e^{-j2\pi f\tau} \cdot a \\ b = e^{-j2\pi f\tau} \cdot b' \end{cases}$

Equation 4.1. Impact of the Test Port Cable on the Incident and Reflected Waves

$$\Gamma_{test \ port} = \frac{b}{a} = e^{-j4\pi f\tau} \cdot \frac{b'}{a'} = e^{-j4\pi f\tau} \cdot \Gamma_{DUT}$$

Equation 4.2. Impact of the Test Port Cable on the Measured Mismatch

As observed in **Equation 4.2**, the open at the end of the test port cable will show up as a short at the test port for all frequencies, such that $4\pi f\tau = \pi + k \cdot 2\pi$ or $\tau = 1/(4f) + k/(2f)$ for $k \in \mathbb{Z}$. More specifically, an open will appear as a short when the cable corresponds to a quarter wavelength.

Although a factory calibration might be useful to verify the accuracy of certain elements of a VNA, such as the power of the stimulus and the frequency response of the receivers, other characteristics related to the measurement setup, such as cable length, are impossible to account for during a factory calibration. As a result, the user calibration is an important step that not only accounts for variables due to measurement configuration, but user calibration can also account for instrument variation over temperature and frequency.

Sources of Error in VNA Measurements

Although the location of the reference plane on VNA measurements is one factor that affects measurement results, VNAs are also subject to a range of inherent instrument impairments that can also be accounted for through calibration. In general, there are four main contributions to this imperfection, including:

- port match
- directivity
- frequency response
- isolation

The sources of error listed above are often referred to as *systematic* sources of error because they systematically affect the measurement at all times. The effects of systematic errors on a measurement result, shown in **Figure 4.2**, can largely be removed through calibration.



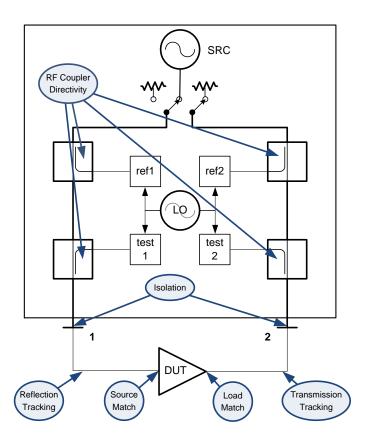


Figure 4.2. Major Contributions to Systematic Errors

Port Match

The port match of a VNA port has significant impact on the measurement accuracy of the instrument. Ideally, the test ports of a 50 ohm VNA would present itself as exactly 50 ohms. However, in practice, the precise impedance of the test port of a VNA can vary according to the quality of the hardware design. Moreover, even well designed hardware has a return loss typically around 20 dB. As a result, the effect of the impedance of a VNA test port cannot be neglected while performing accurate VNA measurements and is referred to as source match and load match.

Directivity

A VNA utilizes separation hardware such as directional couplers or bridges to independently measure the incident and reflected waves at both the input and output of the DUT. These measured quantities are referred to as the *raw* quantities. Although a perfect coupler would only sense the wave it is expected to sense, this is not absolutely true. For example, as observed in **Figure 4.3**, some portion of the reflected wave traveling in the reverse direction of the through path will still present itself in the coupled path.



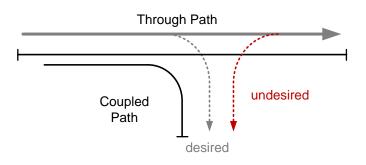


Figure 4.3. Imperfect Directional Coupler

Thus, the isolation of the coupled arm and the signals traveling in the reverse direction is finite. As a result, the coupled arm will also sense some of the wave traveling in the other direction. This phenomenon is referred to as *imperfect directivity*. Imperfect directivity is a systematic error and can be compensated for as part of the calibration process.

Frequency Response

The fundamental design of a VNA involves multiple receivers. Because the signal path corresponding to each receiver is not identical, each receiver can have a slightly different frequency response. The coupler can actually contribute to the frequency response variation between receivers, because each coupler has a slightly different coupling factor as a function of frequency. The phenomenon of each receiver having a slightly varying frequency response is often referred to as a *reflection and transmission tracking* within the context of a VNA calibration.

Isolation

Although a VNA would ideally have infinite isolation between ports, this is not always the case in practice. In fact, some of the signal measured at one port can leak into the receiver channels corresponding to another port. This phenomenon is sometimes referred to as *crosstalk* and can also be compensated for as part of the calibration process.

Types of VNA Calibration

Given the wide range of impairments that can prevent a VNA from performing an accurate network analysis measurement, calibration is required in order to measure each of these impairments and correct the measurement result. There are a wide range of methods to calibrate a VNA, depending on the number of ports, the frequency range, the available calibration standards, and even the port type of the DUT. The port type of the VNA can be coaxial, on wafer, in fixture, or wave guide. Another differentiation of VNA calibration types is the trade-off of accuracy versus speed. For a two-port, full S-parameter VNA, the following calibration types are most common:

- Frequency response calibration
- One-path, two-port calibration
- Full S-parameter calibration



The frequency response calibration (or transmission normalization) is the simplest of all calibration types and merely corrects for the frequency response of the instrument. In general, you should employ this calibration type only when a rough idea of the linear performance (S_{21} or S_{12}) of a DUT is sufficient. Transmission frequency response calibration does not compensate for directivity and mismatch errors. As a result, the quality of this basic calibration depends on the raw directivity and mismatch performance of the VNA.

A second calibration type, the one-path, two-port forward calibration, allows you to perform forward S-parameter measurements. This calibration type is often sufficient when measuring the input match (S_{11}) and the forward transmission (S_{21}) of the DUT. It is important to note that the forward transmission is only partly corrected using this calibration type, since port 2 of the VNA is not fully characterized as part of the calibration process. This calibration assumes that port 2 is perfectly matched and allows for reasonably accurate S_{11} and S_{21} measurements. S_{11} not only includes the input match of the DUT but also the (load) match at port 2 in combination with S_{12} of the DUT. In the case of a unilateral device, where the amplitude of $|S_{12}|$ is very small, the negative effect of a non-perfect match at port 2 will be minimal.

Finally, the full two-port calibration is the most common and complete calibration type of a two-port VNA. With this method, all four S-parameters are fully corrected using various calibration techniques. Note that correction for imperfect isolation is optional and can often be omitted for most commercially available VNAs. Generally, the isolation portion of the calibration process is only required when measuring high-isolation or high-dynamic range devices.

Modeling Two-Port Calibration

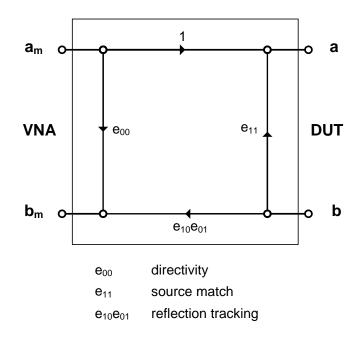
In order to understand exactly how the calibration process is used to correct for impairments in the instrument, it is first worthwhile to understand how to model a vector network analyzer and each source of error. For the purposes of this discussion, consider the full two-port calibration model, as it is the most common. Without diving into the mathematics, and not considering stochastic contributions such as measurement noise or connection repeatability, calibration models are based on linear relationships between the true wave quantities at the calibration planes and the raw measured wave quantities "inside" the VNA.

The classical representation of systematic errors uses an error adapter model, as shown in **Figure 4.4**, and is easily mapped onto the imperfections of the hardware as described above. For instance, $e_{10}e_{01}$ represents the frequency response of the reflected wave (b) path relative to the incident wave (a) path, as S-parameters are ratios.



Introduction to Network Analyzer Measurements

Fundamentals and Background





At each (source) port, the calibration algorithms are designed to compensate for coupler directivity, reflection tracking, and source match. In addition to these errors, two-port calibration introduces three additional error contributions at the other (load) port, including load match, transmission tracking, and isolation. Applying both sets of errors to both the forward and reverse model results in 12 error contributions in total, hence the term *12-term error model*. In most scenarios, the contribution of the isolation can be neglected, which reduces the model to a 10-term error model. Finally, the VNA applies the error terms of the forward and reverse model onto the raw S-parameters to obtain the corrected S-parameters

In the case of a two-port VNA that has four receivers, you can mathematically describe the relationship between the raw measured waves and the corrected waves using matrix notation. The mapping of raw S-parameters onto calibrated S-parameters is referred to as a *wave formalism* approach. In order to understand this notation, it is important to review the block diagram of a two-port VNA, as illustrated in **Figure 4.5**.



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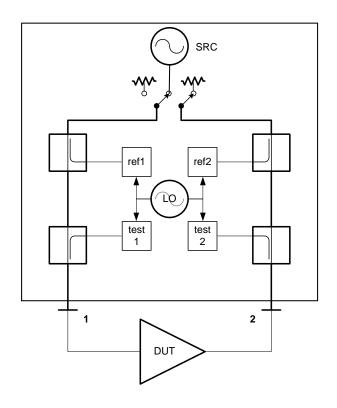


Figure 4.5. Typical VNA Block Diagram

As illustrated in **Figure 4.5**, the raw incident wave (a_{1m}) is measured at reference receiver 1 (ref 1) and the corresponding raw reflected wave (b_{1m}) at test receiver 1 (test 1). Meanwhile, the waves, a_1 and b_1 , are first measured at the reference plane at port 1 and then corrected using error coefficients. A similar scenario applies for port 2.

Assuming the crosstalk (due to imperfect isolation) between both ports can be neglected, the resulting matrix will contain eight non-zero elements within a 4x4 matrix. Because S-parameters only require ratios of wave quantities, one of the elements of the error correction matrix can be independently chosen and the value set to 1. The result is a *7-term error model*, as shown in **Equation 4.3**.

$$\begin{pmatrix} a_1 \\ b_1 \\ a_2 \\ b_2 \end{pmatrix} = K \begin{pmatrix} 1 & \beta_1 & 0 & 0 \\ \gamma_1 & \delta_1 & 0 & 0 \\ 0 & 0 & \alpha_2 & \beta_2 \\ 0 & 0 & \gamma_2 & \delta_2 \end{pmatrix} \begin{pmatrix} a_{1m} \\ b_{1m} \\ a_{2m} \\ b_{2m} \end{pmatrix}$$

Equation 4.3. Wave Formalism (7-Term Error Model)

Note that the ultimate error correction model is the *16-term error model* and is based on the determination of all 16 terms. Using a full (16-term) matrix, all possible systematic errors (resulting in contributions of any raw measured quantity to any corrected quantity) are taken into account.



Two-Port Calibration Methods

The general process of calibration involves connecting a set of known calibration standards to the appropriate reference plane. In general, think of a calibration standard as a device that has a precisely characterized and well-known impedance. The reference plane is the location at which you connect the DUT. Thus, with the calibration standard connected, the VNA measures the phase and magnitude relationship of incident and reflected waves with a well-known impedance attached to the reference plane. Using these results, the VNA calibration process compensates for internal impairments of the instrument and returns a more accurate measurement result at that reference plane.

In practice, there are several VNA calibration methods, each of which has unique benefits in various applications. The following two VNA calibration methods are discussed in the following sections:

- Short-Open-Load-Through (SOLT)
- Through-Reflect-Line (TRL)

Note that the number and type of standards, and the accuracy with which these standards need to be known, depend on the calibration methods. It is also important to recognize that because user calibration of a VNA takes characteristics from the measurement setup into account, you must take great care not to disturb the calibration plane or test setup after calibration. In addition, it is essential to use high-quality, phase-stable test port cables to minimize the impact of various bending of the cables during the measurement of the DUT compared to calibration.

SOLT Calibration

The most popular coaxial calibration method and probably the first manual calibration performed by a new VNA user is referred to as either SOLT (Short-Open-Load-Through) or TOSM (Through-Open-Short-Match). This calibration method involves the process of connecting well-known calibration standards to each port. Next, both ports are physically connected to each other. In a simple case, when preparing to measure an insertable coaxial device, which has mating connectors, both calibration ports can be directly connected to each other resulting in a zero-length through or flush through.

Dealing with non-insertable devices requires a slightly more complex variant of the SOLT calibration method. This situation is one in which the DUT has identical genders at both ports (for instance, two female SMA connectors) or different port types (for instance, 3.5 mm and 2.4 mm). The SOLR method assumes the use of a reciprocal through, hence the "R" instead of "T". In general, there are two primary techniques you can choose to deal with non-insertable devices: usage of phase-equal adapters or the application of the SOLR calibration technique. The SOLR technique is becoming more widely accepted, and hence much more common, as opposed to the traditional usage of phase equal adapters when dealing with non-insertable devices.

One of the simplest approaches to handling non-insertable devices is to use phase-equal precision adapters. Phase-equal adapters are precision matched adapters for various gender combinations, and each phase-equal adapter has identical dimensional and therefore identical electrical characteristics. Phase-equal adapters allow for a zero-length through during calibration, even when the test ports have identical genders, as shown in **Figure 4.6**.



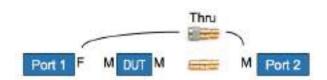


Figure 4.6. Usage of Phase-Equal Adapters to Support Non-Insertable Devices

As illustrated in **Figure 4.6**, a male-female adapter extends the test port cable at port 2. As a result, the reference plane at port 2 now corresponds to the male side of this adapter and thus the short, open, and load must be connected there as part of the calibration at port 2. Because the adapter has been included as part of the calibration process, a zero-length through connection is now possible between the original female connector of the test port cable at port 1 and the male connector of the test port cable at port 2.

After calibration, the male-female adapter is swapped by a phase-equal female-to-female adapter to appropriately connect to the male-to-male DUT. Given that both adapters have identical electrical characteristics, the calibration remains valid even though the genders of the adapters are different. Note that phase-equal adapters are typically provided as part of a calibration kit or as an option.

SOLR Adaptation to SOLT

A second and preferred solution to adaptor removal in SOLT calibration is the use of a slightly different calibration method, which is referred to as SOLR (Short-Open-Load-Reciprocal), also known as *Unknown Through*. This technique only requires a reciprocal ($S_{21} = S_{12}$) structure to be used during through calibration. In some situations, it is necessary to provide an estimate of the delay of the unknown through. When performing a SOLR calibration, the first step is to perform a one-port calibration at both ports using short, open, and load. Next, use a forward and reverse measurement to estimate both the S-parameters of the reciprocal structure (S_{11} , S_{22} and $S_{21}=S_{12}$). As a result of these measurements, you can determine the scaling factor which links the error coefficients at port 2 to those at port 1.

A final solution that allows calibrated S-parameter measurements of non-insertable devices is referred to as adapter removal calibration. Adaptor removal requires two two-port calibrations. The first two-port calibration is performed using the adapter at port 1 and a second calibration is performed using the same adapter at port 2. Combining both calibrations yields a single full two-port calibration that allows you to accurately measure the non-insertable device.

Challenges with SOLT Calibration

The biggest drawback of SOL-based calibration techniques is the assumption that the load is ideal (generally 50 ohms). Fortunately enhancements do exist. One option is to measure the load after a TRL calibration (explained in the following section) and provide its S-parameters as part of the calibration kit. By providing the S-parameters for the load, it is no longer required to assume that the load is perfect. An alternative is to use an *ideal* load only at low frequencies (typically up to a couple of gigahertz) and to instead use a sliding load above that frequency. At higher frequencies, the combination of a precision air line and the *sliding load* allows you to create a *sliding* mismatch which is applied at different distances with respect to the reference plane. The corresponding mismatch results in a circle and the estimation of the center of that circle allows a better determination of the directivity.



A second challenge with the SOLT technique is the practical challenge of manufacturing the calibration standards. For example, in transmission mediums such as a wave guide, it is difficult or impossible to come up with an open. This difficulty can be somewhat circumvented by replacing the open with a second short, which has a different length and is referred to as offset short. Using a single additional offset short, you can accurately calibrate the VNA over a limited frequency range.

If you have a need for (significantly) more than two ports, it is worthwhile to remember that there is a QSOLT calibration technique, which only requires a short-open-load calibration at a single port and a known through connection between that port and every other port. The "Q" in QSOLT stands for "quick" and is a useful calibration tool given the complexity of the standard SOLT technique.

TRL Calibration

A second popular calibration technique is TRL (Through-Reflect-Line). TRL calibration is typically used for onwafer or in-fixture measurements, where the use of precise calibration standards is impractical or impossible. In general, using the TRL calibration technique requires you to create your own in-circuit calibration standards that appropriately mate with the measurement fixture. Here, the *through* is essentially a zero-length microstrip that directly connects each end of the test fixture. The *reflect* standard is generally a crude (but repeatable) short or open circuit. Finally the *line* is similar to the through but has a non-zero length.

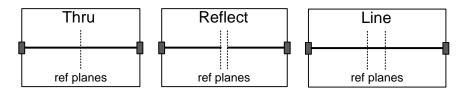


Figure 4.7. TRL Calibration Standards

The TRL calibration method first involves the connection of a reflect to each port. Note that the accuracy of this calibration technique requires that the reflect standard presented to each port be identical. In order to ensure that each port is presented with an identical reflect standard, one option is to use a single reflect standard and connect it to both ports. Next, measure a through and then a line. Here, the assumption is that the through and the line are perfectly matched and as a result, define the characteristic impedance of the system.

In practice, it is often easier to fabricate a well-matched line than it is to fabricate a well-matched broadband load. Note that the characteristic impedance of the reflect does not need to be known exactly. When performing a TRL calibration, you only have to indicate whether it is approximately a short or an open for the frequency range of interest.

One of the fundamental limitations of the TRL approach is that it can only be used over a relatively small frequency range. The usable frequency range is defined by the length of the line according to **Equation 4.4**.



$$l \neq n \cdot \frac{\lambda}{2}$$
 with $\lambda = \frac{v_p}{f_0}$ and $v_p = \frac{c}{\sqrt{\varepsilon_r}}$

Equation 4.4. Impact of the Line Length on the Frequency Range of the TLR Calibration

In fact, at specific frequencies, the line looks too much like a through, such that it is impossible to extract meaningful error coefficients at these frequencies. As a rule of thumb, the frequency range is determined by a 20° to 160° region as shown in **Equation 4.5**.

$$f_{min} = \frac{v_p}{18I}$$
 and $f_{max} = 8 f_{min}$.

Equation 4.5. Calculating Frequency Range of the TLR Calibration

One solution to the frequency range limitations of the TRL technique is to use two or more lines, each of which have an appropriate length to cover different overlapping frequency bands. The corresponding calibration technique is referred to as *multi-line TRL*. In the situation where a zero-length through is impractical, a short line can be used instead, resulting in an LRL (Line-Reflect-Line) calibration. Note that it is important to accurately specify the propagation constant (delay and loss) of the shortest line in order to properly define the reference planes.

It makes sense to briefly comment on the usage of an automatic calibration module. The automatic calibration module contains different impedance standards that are carefully calibrated by the manufacturer, typically using a TRL-grade calibration. So, on top of the ease of use, an automatic calibration module also provides a high-quality calibration.

Tools for VNA Calibration

When calibrating a network analyzer, engineers can utilize a wide range of calibration and verification tools. For example, the load, short, or open used to calibrate a network analyzer is often part of a manual calibration kit. Verification kits are typically used to verify the performance of the VNA after a calibration and can be used to identify problems with the calibration procedure or even the calibration standard.

A calibration kit is the primary tool used to calibrate a VNA. Today, VNA manufacturers provide a wide range of calibration kits for various connector types and for both manual and automatic calibration. Calibration kits contain calibration elements, or calibration standards that have been carefully characterized based on their mechanical dimensions or based on measurements using a golden network analyzer, that was calibrated using traceable calibration standards.

A calibration kit also contains precise information about each calibration element behavior, and the VNA uses this information as part of the calibration routine. Calibration element data can be based on coefficients corresponding to polynomial models or based on a sufficiently dense set of S-parameter values as a function of frequency. Although connector care is considered to be standard good practice at RF and microwave frequencies, calibration elements should be treated extra carefully.



Manual Calibration Kits

Calibration kits can be either manual (mechanical) or automatic. When using a mechanical calibration kit, connect the correct calibration element to the appropriate port when requested to do so by the VNA software. Typically, you first connect each of the required calibration elements to each port one by one and then connect the VNA cables to realize a through connection between both ports. You should generally connect calibration standards through a direct connection if the genders of the calibration planes allow it. In scenarios where the gender of the calibration standards does not allow a direct connection, you can use a reciprocal ($S_{21} = S_{12}$) structure or a different technique based on phase-equal adapters or adapter removal.

Electronic Calibration Kits

In contrast to a manual calibration, electronic (or automatic) calibration modules are generally much easier to use. Calibrating a VNA with an automated device simply requires you to connect the ports of the calibration module to the ports corresponding to the calibration planes before pushing the "start calibration" button. In most cases, the calibration module is advanced enough to discover which ports are connected to which calibration planes. When using an electronic calibration module, the through connection can either be realized internally, inside the calibration module, or externally. When the through connection is realized externally, the connection is made in a manner similar to the manner used for a mechanical calibration kit.



Figure 4.8. Automatic VNA Calibration Kit Provided by National Instruments

Calibration Verification

Verification kits can be used not only to verify the accuracy of the calibration process, but also to verify that the VNA is still operating within its specifications. A verification kit typically contains precision air lines, mismatched air lines, and precision fixed attenuators. These elements are measured after carefully performing a calibration and compared to their known characteristic to determine if the VNA is still within specification.

It is also important to consider the frequency of the calibration interval. Assuming the current VNA settings (start and stop frequency, number of frequency points, IF bandwidth, source power, and receiver step attenuator values) can be used, the need for calibration depends on the time and the variation of environmental parameters, such as temperature, since the last calibration. Again, if you require extremely accurate measurements, it is worthwhile to



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take the time and effort to perform a new calibration to eliminate all doubts about the validity of the error coefficients that are currently being used.

Calibration Considerations

When calibrating a VNA, it is important to use identical hardware and software settings during both the calibration of the VNA and the measurement of the DUT. For example, typical measurement settings options that should remain constant include frequency settings, IF bandwidth and averages, source power, and receiver step attenuator settings.

Frequency Settings

Basic frequency settings simply require a start and stop frequency and number of frequency points, although more complex frequency settings are possible. In fact, more advanced configurations allow for multiple frequency segments, each allowing for varying frequency resolution for selected frequency bands. Using this feature, you can zoom in to certain frequency ranges for which the behavior of the DUT is expected to change more rapidly with frequency than for other frequency ranges.

IF Bandwidth and Averaging

IF bandwidth and averaging can significantly affect both the measurement speed and accuracy of a VNA measurement. Decreasing the IF bandwidth will decrease the uncertainty on the measurements at the cost of increased measurement time and is an alternative to averaging. In general, calibration is most accurate when performed using a large number of averages. In fact, it is possible to use a large number of averages for the calibration routine and then reduce the number of averages used when measuring the DUT.

One final note about IF bandwidth is that various commercial VNAs indicate that the calibration is no longer valid when selecting a different IF bandwidth during a DUT measurement. When changing the IF bandwidth, be aware that the IF bandwidth setting used for the measurement must be the same or wider than the setting used for calibration, because the uncertainty on the calibrated measurement is defined by both the uncertainty of the raw measurement and the uncertainty of the error coefficients. The latter depends on the IF bandwidth that was selected during calibration.

Source Power

In principle, the source power should be chosen to be as high as possible to achieve the smallest uncertainty for a given IF bandwidth. However, remember to make sure that both the DUT and the receivers of the VNA are used in their linear mode of operation.

Receiver Step Attenuator Settings

Certain VNAs include receiver step attenuators in order to guarantee linear operation of the VNA receivers. In general, never change these attenuator settings in between calibration and DUT measurement.



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Advanced Calibration Techniques

For some measurements, advanced calibration techniques are required due to the complexity of the test setup or the implementation of an advanced measurement. The following section discusses two advanced calibration techniques: de-embedding and power calibration.

De-embedding

Unfortunately, it is not always possible to perform a calibration and thereby establish a reference plane at the exact desired location. While the ideal reference plane is typically at the input and output of the DUT, it is often impossible to connect the appropriate calibration standards at this location. In these cases, use de-embedding techniques to move the calibration plane from one location to another, ideally as close as possible to the DUT.

Generally, de-embedding requires the knowledge of the S-parameters of whatever structure is between the reference plane and the DUT. For example, suppose you are testing a DUT in a test fixture. In this case, one of the most difficult tasks is to create an accurate model of the test fixture. In many cases, the S-parameters for the test fixture are provided by the manufacturer. In other instances, it is important to first extract the S-parameters using simulation tools in order to be able to de-embed the fixture. Today, most commercially available VNAs support simple corrections of fixture effects by means of port extensions.

Power Calibration

Today, vector network analyzer technology is moving beyond standard S-parameters by adding nonlinear measurement capabilities. Although the behavior of linear devices is, by definition, independent of the level of the input signal, the same is not true for the nonlinear behavior of the device. As a result, the precise power level applied to the DUT is an important element of making basic nonlinear measurements.

As already mentioned, a VNA is typically calibrated at manufacturing time at its test ports, with respect to both the source and the receivers. However, it is generally impossible to directly connect the DUT to the test port. Thus, in order to properly compensate for losses introduced by test port cables or other means to connect the DUT to the VNA, you must perform a *source calibration*. By comparing the power level applied by the source to the power level measured using a power meter at the location where the DUT will be connected, the VNA is able to build an internal correction table such that the desired source power is applied to the DUT.

Finally, in order to measure the compression characteristic of a nonlinear device, a *receiver calibration* can be performed, typically in combination with a source calibration, in much the same way as the source calibration. Here, the receiver is used to measure the absolute power levels of incident and reflected waves at the excitation frequency and its harmonics, instead of only the S-parameters that correspond to ratios of wave quantities.

A Final Word of Advice

Performing a calibration is only worth the effort when applying good measurement practices. Good measurement practices include the use of high-quality components, such as phase-stable test port cables, proper connector care, and the correct use of a torque wrench.



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5. Glossary

| A_1 | incident power wave |
|-----------------------|----------------------------|
| A_2 | reverse incident wave |
| B_1 | reflected power wave |
| B ₂ | reverse reflected wave |
| CRT | Cathode Ray Tube |
| CW | Continuous Wave |
| DUT | Device Under Test |
| IF | Intermediate Frequency |
| RF | Radio Frequency |
| S | S-parameter Matrix |
| SNA | Scalar Network Analyzer |
| SOLT | Short-Open-Load-Through |
| SOLR | Short-Open-Load-Reciprocal |
| SRD | Step Recovery Diode |
| TDR | Time Domain Reflectometry |
| TOSM | Through-Open-Short-Match |
| TRL | Through-Reflect-Line |
| VNA | Vector Network Analyzer |
| Z_0 | Characteristic Impedance |



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6. References

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